

Status of GMES Sentinel 5

2nd Post-EPS User Consultation Workshop

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Presentation Outline

- The GMES atmospheric chemistry missions objectives
- Sentinel 5 observation requirements
- Overview of instrument concepts
- Technology developments
- Programmatics

GMES Sentinels 4 and 5 primary objectives

- The GMES Atmosphere Service will provide coherent information on atmospheric variables in support of European policies and for the benefit of European citizens. Services are proposed to cover:
 1. Air quality
 2. Climate change/forcing
 3. Stratospheric ozone and solar radiation
- The primary objectives of Sentinel 4 and 5 are:
 - **High temporal and spatial resolution** measurements of tropospheric composition, including Planetary Boundary Layer (PBL), for *air quality applications*
 - **High spatial resolution and high precision monitoring** of tropospheric climate gases (CH₄, CO₂ and precursor CO) and aerosols with sensitivity to PBL concentrations for *climate protocol monitoring*

Implementation of Sentinels 4&5 on EUMETSAT platforms

According to current ESA-EUMETSAT working assumptions on GMES and following recommendations expressed by the GMES Atmospheric Service Implementation Group (GAS-IG), the Sentinels 4 and 5 missions will be accommodated as additional payloads on EUMETSAT platforms:

- Sentinel 4 (GEO component) will consist of
 - A UVN spectrometer embarked on the MTG-S platform;
 - TIR data from the IR sounder onboard the same platform; and
 - Imagery data from the imager on-board MTG-I.
- Sentinel 5 (LEO component) will consist of
 - A UVNS spectrometer embarked on the Post-EPS platforms;
 - TIR data from the Post-EPS IR sounder; and
 - Imagery data from the imager on-board Post-EPS.

Sentinel 5 geometric requirements

Global Daily Coverage

The Sentinel 5 flight segment shall allow a global daily coverage of the Earth atmosphere.

Spatial Sampling Distance (SSD)

UV-1 15 km (G) / 50 km (T)

UV-2, VIS, NIR, SWIR 5 km (G) / 15 km (T)

Geolocation knowledge < 0.1 SSD (G) / 0.2 SSD (T)

Inter-band spatial co-registration < 0.01 SSD (G) / 0.05 SSD (T)

Observation Zenith Angle (OZA) < 66°

Sun Zenith Angle (SZA) < 92° (requirements to be met up to 80°)

Sentinel 5 spectral and radiometric requirements

Band ID	Spectral bands [nm]	Spectral Resolution [nm]	Spectral Sampling Ratio	SSD [km]	SNR @ Lref_TR	SNR @ Lref_HL	SNR @ Lsun	Priority
UV-1	270-300	1.0	3	15 (G), 50 (T)	100	100	1000 (270 nm)	1
UV-2a	300-308	0.5	3	5 (G), 15 (T)	100	100	3000 (208nm)	1
UV-2b	308-400	0.5	3	5 (G), 15 (T)	1000	1000	3000	1
VIS-1	400-405	0.5	3	5 (G), 15 (T)	1000	400	3000	2
VIS-2	405-500	0.5	3	5 (G), 15 (T)	1050	450	4500	1
VIS-3	590-640 / 610-680 / 710-750	0.4	3	5 (G), 15 (T)	700	700	2000	2
NIR	750-775	0.05 (G), 0.4 (T)	3	5 (G), 15 (T)	200/1000	200/1000	1500	1
SWIR-1	1593-1672	0.25	3	5 (G), 15 (T)	190		1000	2
SWIR-2(*)	1940-2030	0.25	6 (G), 2 (T)	5 (G), 15 (T)	100		1500	3
SWIR-3	2305-2385	0.25	6 (G), 2 (T)	5 (G), 15 (T)	90		1500	1

- The CAMELOT study provided scientific support to the consolidation of the requirements. EUMETSAT also involved.
- Concepts trade-off were based on the priorities: SR, OZA, SSD and SNR
- In yellow the bands implemented in the Sentinel 5 Precursor mission

*) SWIR-2 band was not implemented because of priority 3

Sentinel 5

Instrument concepts

Sentinel 5: General characteristics

- Two instruments concepts (A and B) have been studied by industry in the frame of the Sentinel 4 and 5 phase 0 studies
- Instruments are compatible with the Post-EPS orbit at LTAN of 9:30
- Global daily coverage from the Post-EPS altitude of 817 km
- Wide swath (~2657 km) pushbroom spectrometer concept
- Field of view: 108° ACT and < 4° ALT
- 6 grating spectrometers: 4 UVN, 2 SWIR
- Detector technology: CCD for UVN, CMOS for SWIR
- Separate telescopes for UVN and SWIR
- Passive cooling concepts
- Irradiance measurement in anti-flight direction to reduce diffusers contamination
- The instrument is expected to be always kept ON for temperature stability
- Data rate acquired during ~70% of the orbit (including 15% for calibration)
- Instruments aperture of ~3 mm
- Both instrument concepts are an evolution of the Sentinel 5 Precursor concept

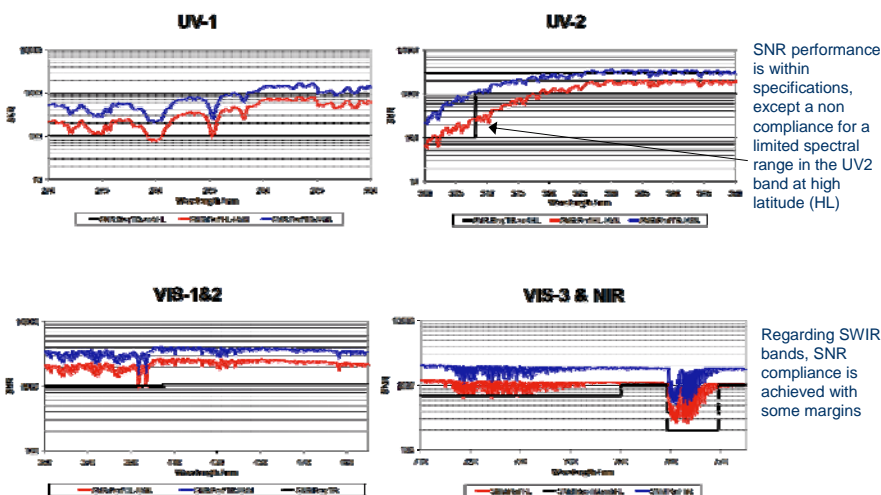
Sentinel 5: General performance (1/2)

- Performance presented are applicable to both concepts unless explicitly stated

Spatial sampling distance (Level 0 data)	15 km (UV1) 5 km (UV2/VIS/NIR/SWIR)			
Spectral bands	Band ID	Spectral range (nm)	Spectral resolution (nm)	Spectral sampling Ratio ⁽¹⁾
	UV-1	270-300	1	3.5
	UV-2	300-350	0.5	3.5
	VIS-1&2	350-500	0.5	3.5
	VIS-3&NIR	710-775	0.4	3.8
	SWIR-1	1593-1672	0.25	3.1
	SWIR-3	2305-2385	0.25	3.1
Interband spatial co-registration ⁽²⁾	Concept (single/double telescope) dependent			
Sensitivity to polarisation ⁽³⁾	< 0.5% in VNIR < 5% in SWIR			
Radiometric accuracy	Compliance TBC (needs further analysis)			

- 1) Concept B has a sampling ratio of 3.2 in UVN
- 2) The inter-channel spatial co-registration is achieved in both concepts
- 3) Polarisation compliance is dependent on the position of the scrambler:
 - In concept A scrambler is at intermediate pupil => specs are not met
 - In concept B scrambler is at entrance pupil => specs are met

Sentinel 5: General performance (2/2)



SNR estimate based on 3x3 on-ground spatially binned pixels

Sentinel 5: Calibration and Detectors

Calibration

- The instrument is expected to operate with 100% duty cycle on the bright side of the orbit
- Once per day (TBC) this is interrupted for ~1 min at the North pole to observe the sun via the dedicated sun diffuser (solar irradiance measurement)
- At the dark side of the orbit, the instrument will regularly perform the following instrument calibrations:
 - Detector dark current measurements (every orbit)
 - LED and White Light Source (WLS) measurements (once a week TBC)

Detectors

- Both industrial teams are confident about the availability of both (UVN/SWIR) detectors in time for a launch in 2019

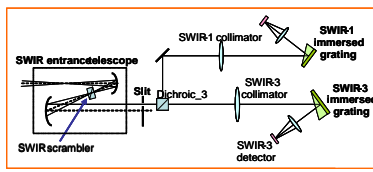
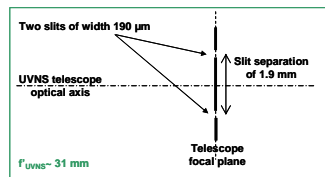
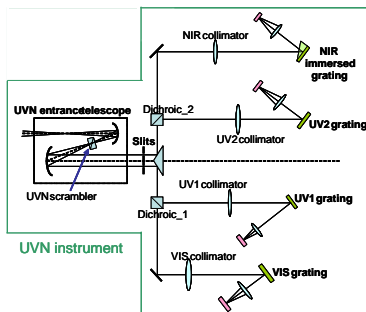
	CCD for UVN	CMOS (MCT) for SWIR
Pixel pitch	13 / 15 μm	20 / 30 μm
Format	1.25 x 1.2 k / 1 k x 1 k	1 k x 1 k / 1000 x 768
Temperature	210 K in UV 250 K in NIR	200 K in SWIR-1 160 K in SWIR-3
Notes	CMOS was discarded after trade-off	Baseline derived from existing Sofradir Saturn

Sentinel 5

Concept A

Sentinel 5: Optical layout

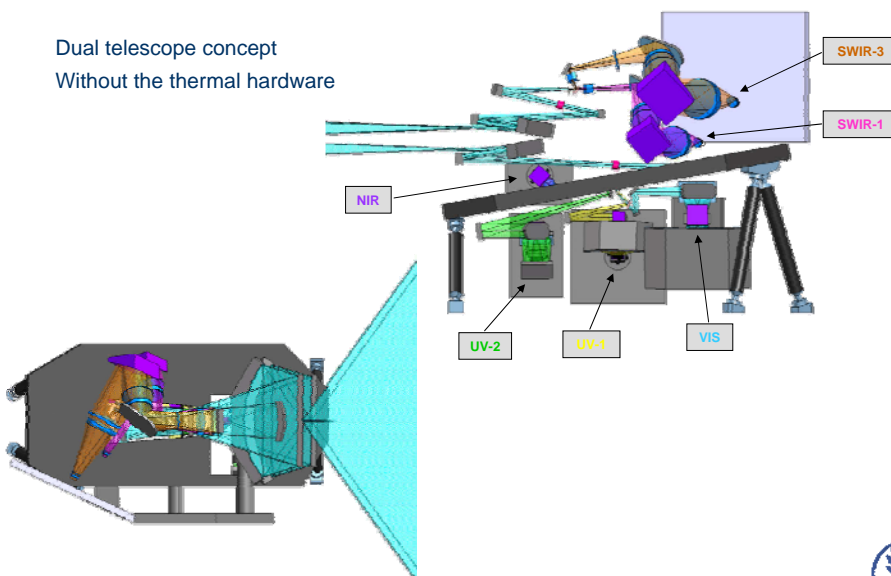
- UVN/SWIR aperture 3.1 mm
- Two slits in-field separated at UVN telescope focal plane (\Rightarrow ALT FoV $\sim 0.5^\circ$) (\Rightarrow 8 s of temporal corr.)
- Two dichroics allow to separate the UVN bands before spectrometers
- One dichroic allow to separate the SWIR bands
- Immersed gratings used for NIR, SWIR-1 and SWIR-3



SWIR instrument

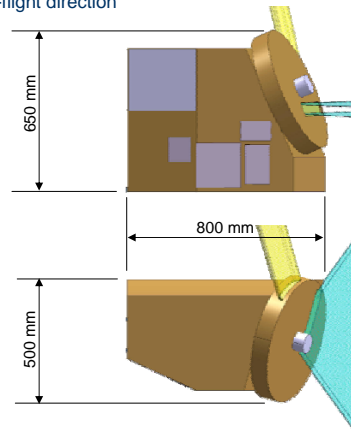
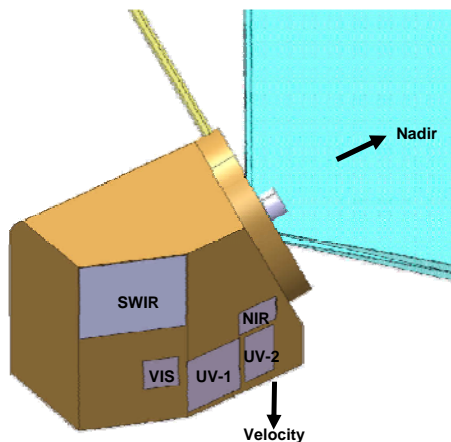
Sentinel 5: Mechanical implementation (1/2)

Dual telescope concept
Without the thermal hardware



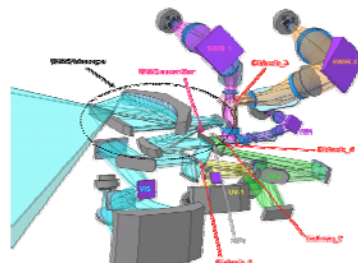
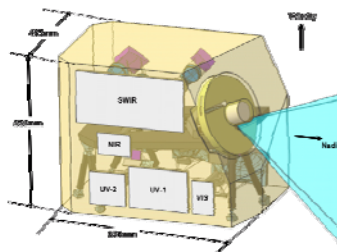
Sentinel 5: Mechanical implementation (2/2)

- Sun irradiance measurement direction is 37.5° w.r.t. anti-flight direction
- Accommodation of all radiators on the anti-sun face



Sentinel 5: Single/Dual telescope trade-off

- The evolution analysed consists in suppressing one telescope, and thus in implementing all the UVNS bands with a single telescope



	Single telescope	Dual telescope
Mass (*)	+	-
Volume	+	-
Spatial co-registration	+	-
Polarisation sensitivity	-	+
Calibration	-	+

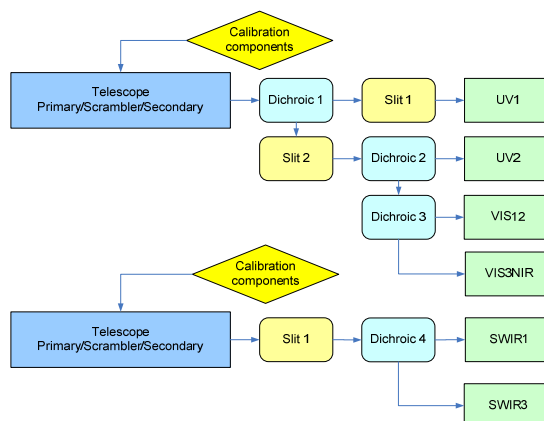
*) The mass savings compared to dual telescope concept are ~10 kg

Sentinel 5

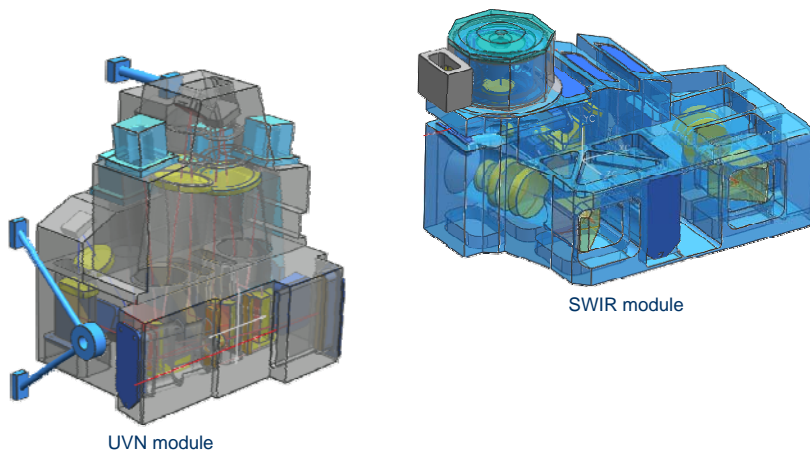
Concept B

Sentinel 5: Optical layout

- UVN aperture 3.6 mm / SWIR aperture 2.1 mm
- Three dichroics allow to separate the UVN bands before the UVN spectrometers
- One dichroic allows to separate the SWIR bands
- Immersed gratings used for SWIR-1 and SWIR-3



Sentinel 5: Mechanical implementation (1/2)

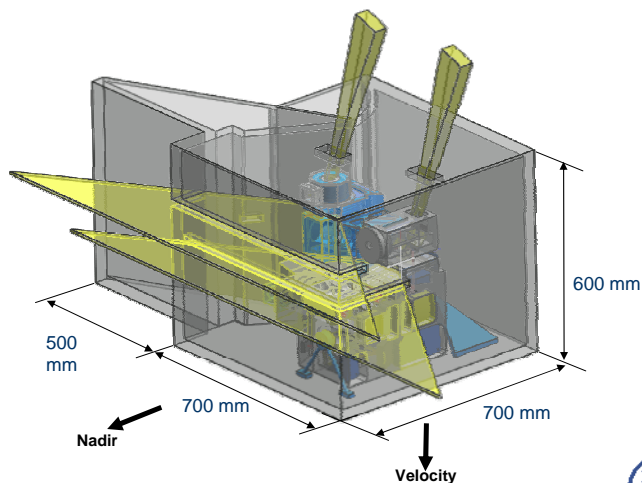


UVN module

SWIR module

Sentinel 5: Mechanical implementation (2/2)

- Sun irradiance measurement direction is 37.5° w.r.t. anti-velocity direction
- Accommodation of all radiators on the anti-sun face



Sentinel 5: Concepts summary

- The baseline is the dual telescope concept

	Concept A	Concept B
Payload mass (*)	181 kg	152 kg
• Optical unit mass	160 kg	128 kg
• Electronic unit mass	21 kg	24 kg
Payload dimensions		
• Optical unit	800 x 500 x 650 mm ³	700 x 600 x 700 mm ³
• Electronic Unit	23 dm ³	75 dm ³
Power budget (*)	96 W	190 W
Data rate (**)	23 Mbps	21.6 Mbps
Data volume per orbit (***)	96 Gbits	90 Gbits

*) Mass and Power budgets include 20% margin

**) Data rate corresponding to a native SSD of 5 x 5 km²

**) Data volume => a 15% overhead is estimated for calibration

Sentinel 5: Technology developments

- The **NIR immersed grating technology**, in order to achieve high dispersion (goal: 0.05 nm spectral resolution) with reasonable grating sizes
- SWIR immersed grating technology** for both SWIR-1 & 3 spectrometers to reach high grating efficiency with reasonable grating sizes
- SWIR MCT detector development**: optimize the ROIC, hybridation and packaging of the detector to the Sentinel 5 needs
- Calibration diffuser technology** and manufacturing process to minimize the spectral features effect
- Specific reflective **coatings** for UVN telescope and UV-1, UV-2 and VIS spectrometers
- Polarization scrambler** device to achieve the required depolarization efficiency
- Specific **dichroics** optimizations in transmission and reflection
- Technological survey on high speed 14 bit ADC

Technology development actions to cover this are initiated or under preparation

Sentinel 5: Programmatic status summary

- Sentinel 5 to be embarked on Post-EPS with launch date ~2019, requires observations in all bands, from UV to TIR => Synergy with Post-EPS IRS is essential: high priority of atmospheric composition requirements needs to be kept in Post-EPS
- Instrument design concepts from Sentinel 5 Phase 0 studies has been fed into Post-EPS Phase 0 studies
- System level aspects (accommodation, envelope, end-to-end data flow,...) are addressed in the frame of the on-going Post-EPS phase 0 system studies
- Sentinel 5 Phase A to take place in 2010, in parallel with Post-EPS phase A
- GMES Space Component (GSC) Segment 2 includes phase B1 and complementary pre-developments
- Phase B2/C/D are for CMIN-2011, synchronized with Post-EPS phase B/C/D

Thank you

Backup Slides

Sentinel 5: Concepts Summary

Concept A

Spectrometer bands	UV-1	UV-2a, UV-2b	VIS-12	VIS-3, NIR	SWIR 1	SWIR 3
Wavelength range (nm)	270 - 300	300 - 350	350 - 500	710 - 775	1593 - 1672	2305 - 2385
Spectral resolution (nm)	1.0	0.5	0.5	0.4	0.25	0.25
Spectral sampling ratio	3.5	3.5	3.5	3.8	3.1	3.1
Native SSD ALT x ACT (km)	45 x 5	5 x 5				
Integration time (s)	6.82	0.76				
SSD binning ALT x ACT	1 x 9	3 x 3				
Product SSD ALT x ACT (km)	45 x 45	15 x 15				

Concept B

Spectrometer bands	UV-1	UV-2a, UV-2b	VIS-12	VIS-3, NIR	SWIR 1	SWIR 3
Wavelength range (nm)	270 - 300	300 - 350	400 - 500	710 - 775	1593 - 1672	2305 - 2385
Spectral resolution (nm)	1.0	0.5	0.5	0.4	0.25	0.25
Spectral sampling ratio	3.2	3.2	3.2	3.2	3.0	2.0
Native SSD ALT x ACT (km)	15 x 5	5 x 5				
Integration time (s)	2.27	0.76				
SSD binning ALT x ACT	3 x 3	3 x 3				
Product SSD ALT x ACT (km)	45 x 15	15 x 15				