

LI-STAR tool DEMO

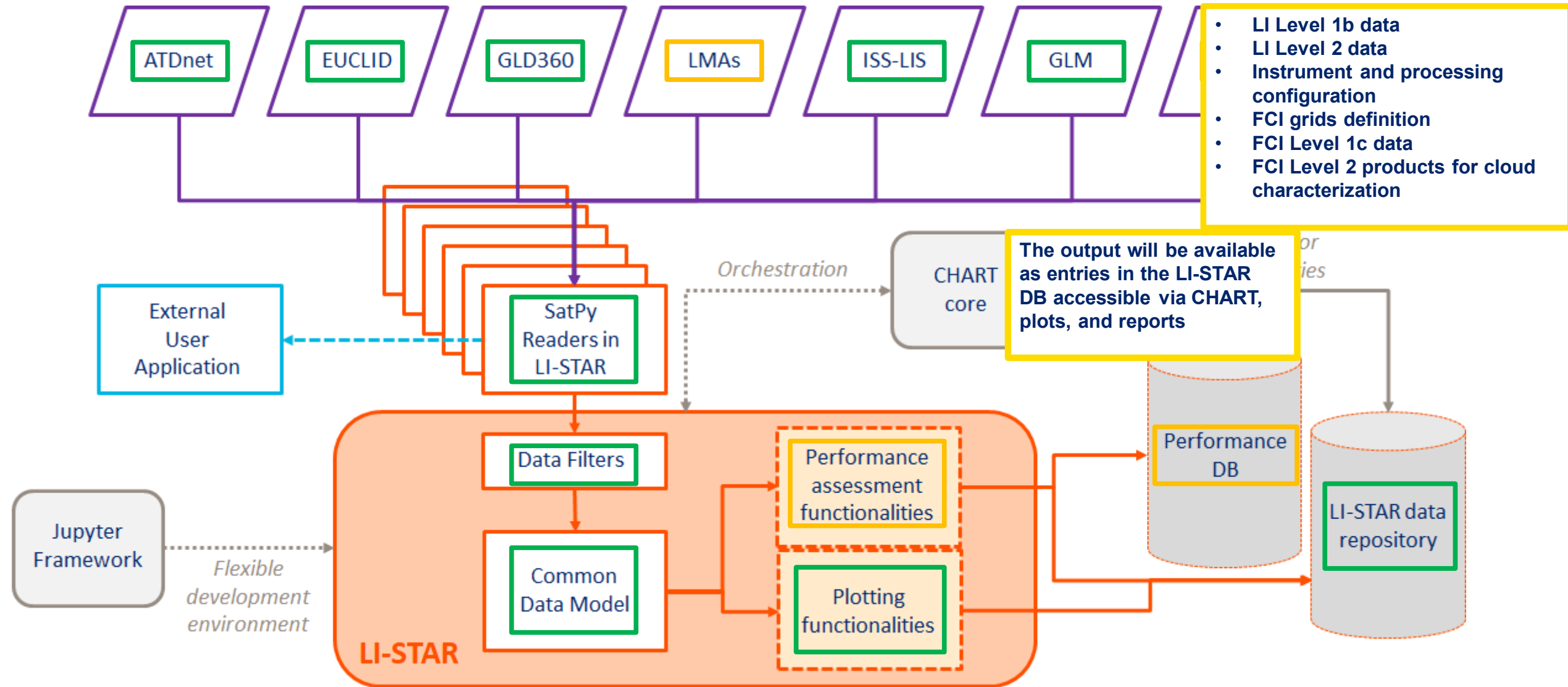
Sven-Erik Enno, Bartolomeo Viticchiè,

Francesco Pignatale

EUMETSAT LI IFCT



Introduction – LI-STAR at a glance



LISTAR available test data

Network	Data period(s)	Data area	Data level(s)
ATDnet	01/01/2018 - now*	All but ATDnet spatial range is mostly limited to Europe, N Africa, the Atlantic, S-America	Groups***
EUCLID	09/2018	All but EUCLID only detects in Europe.	Groups***, flashes
GLD360	09/2018, 01/07/2020 - 04/01/2021 15:00UTC	75N-75S, 75W-75E**	Groups***, flashes
GLM	09/2018, 01/10/2020 - now*	Full GOES-16 GLM FOV	Groups, flashes
ISS-LIS	09/2018, 01/07/2020-29/11/2020	All but consider ISS-LIS short view times from LEO	Events, groups, flashes

* now normally means until the end of the last hour, but keep in mind that the data is received and processed during the next hour, i.e. is not immediately available.

** 09/2018 is also available for the Americas (60N-60S, 135W-15W, contains full GOES16 GLM FOV).

*** coded as `group_level` in LISTAR but scientifically strokes/pulses for ground-based LLSs.

LISTAR output types

Daily filtered lightning data files

Common data model

Matching

Simple flash/group stats

- Time evolution of the number of flash/group detections.
- Comparison of time evolution between different networks.
- Spatiotemporal flash/group distribution maps.
- Flash/group accumulation/climatology maps.

Matching stats

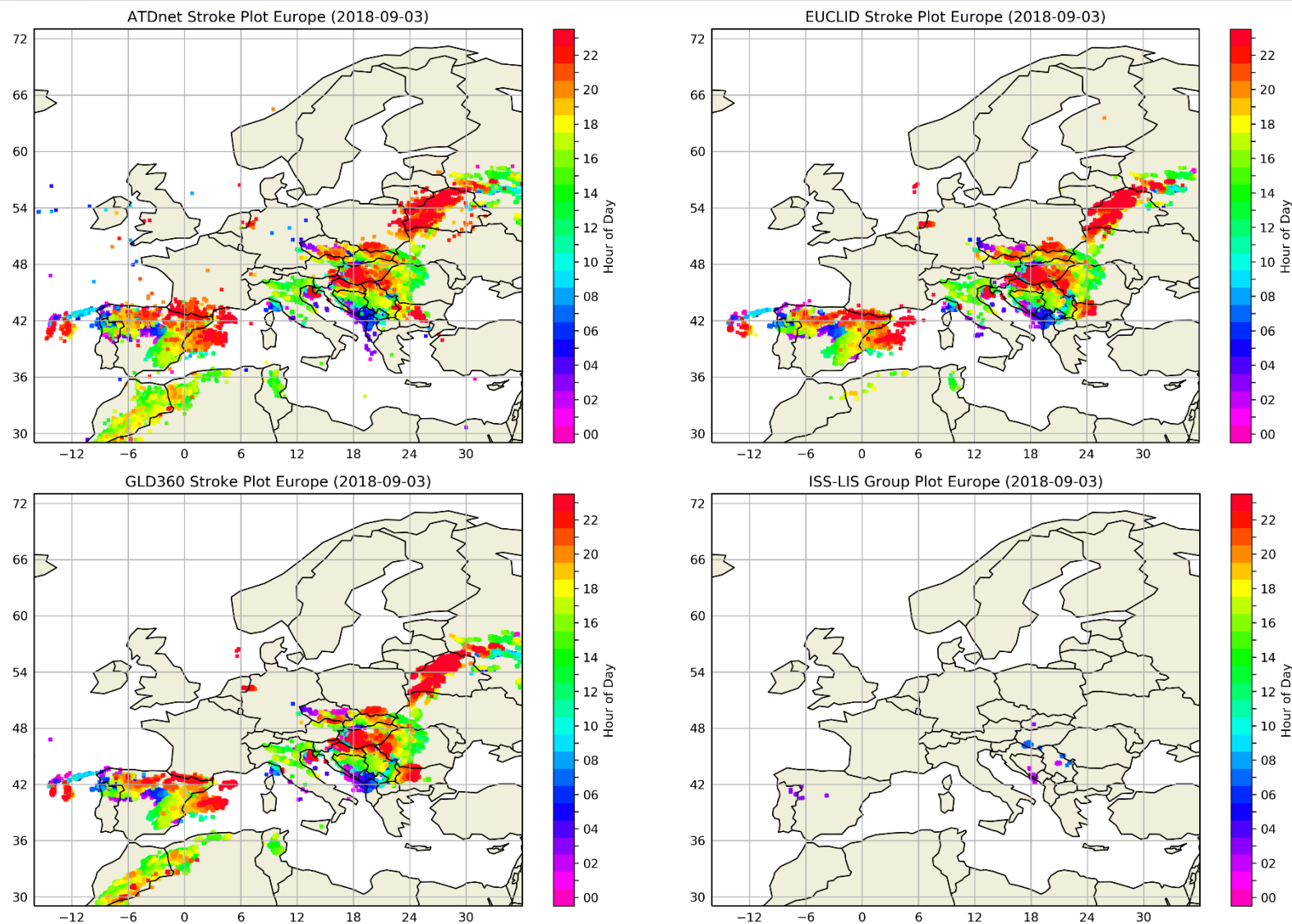
- RDE heatmaps.
- Test-ref spatial offset (histograms and maps).
- Test-reference spatiotemporal offset histograms.
- Diurnal changes in RDE.
- ...

- Instrument and processing configuration
- LI background images
- FCI grids definition and data

Additional performance stats

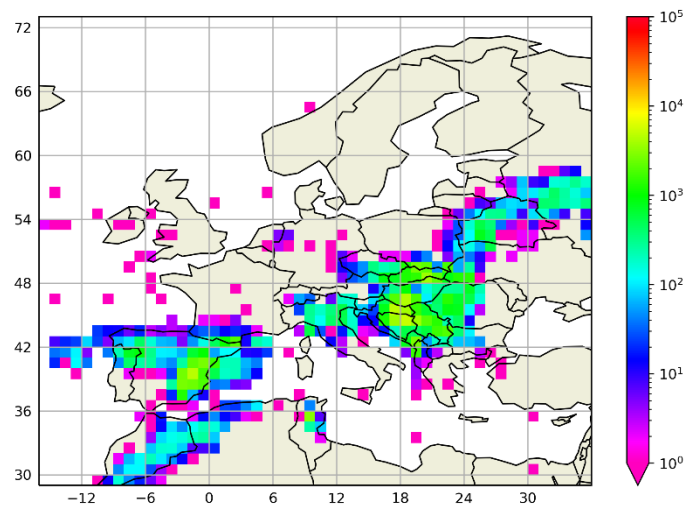
- Radiometric performance assessment.
- Meeting the LI geodisk coverage requirement.
- ...

Spatiotemporal group/stroke distribution maps

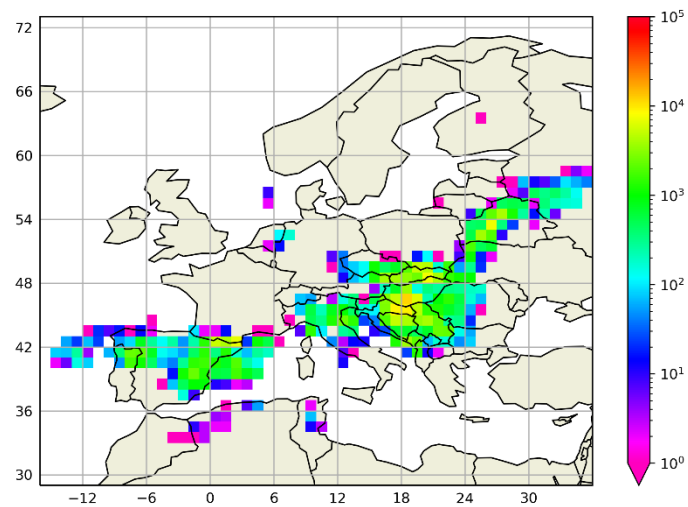


Group/stroke accumulation maps

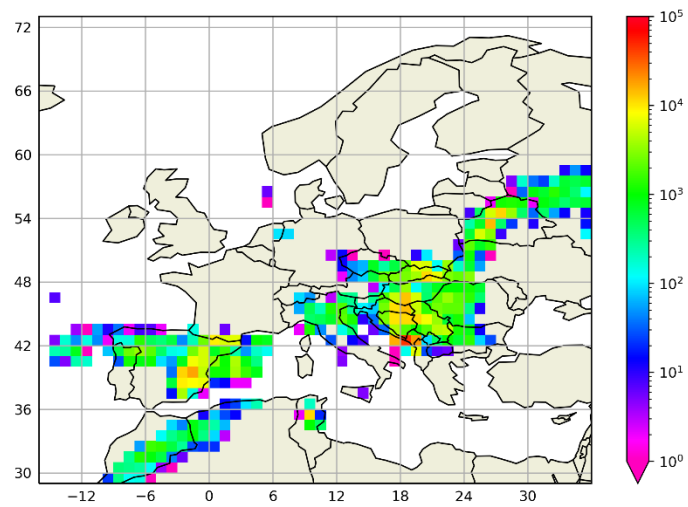
ATDnet Accumulated Strokes Europe 2018-09-03



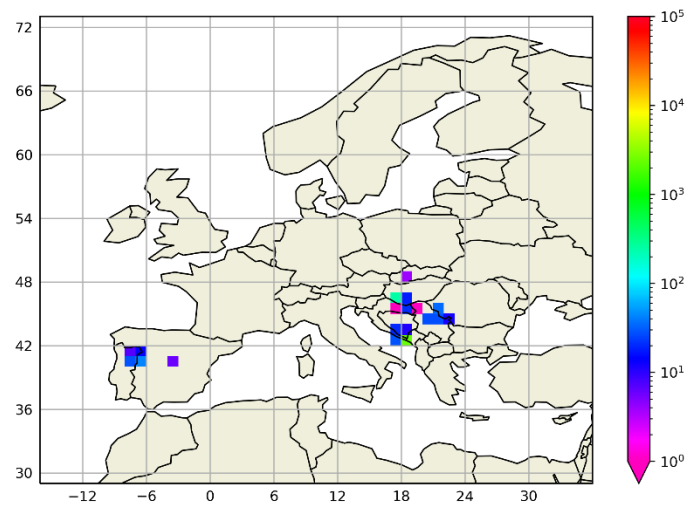
EUCLID Accumulated Strokes Europe 2018-09-03



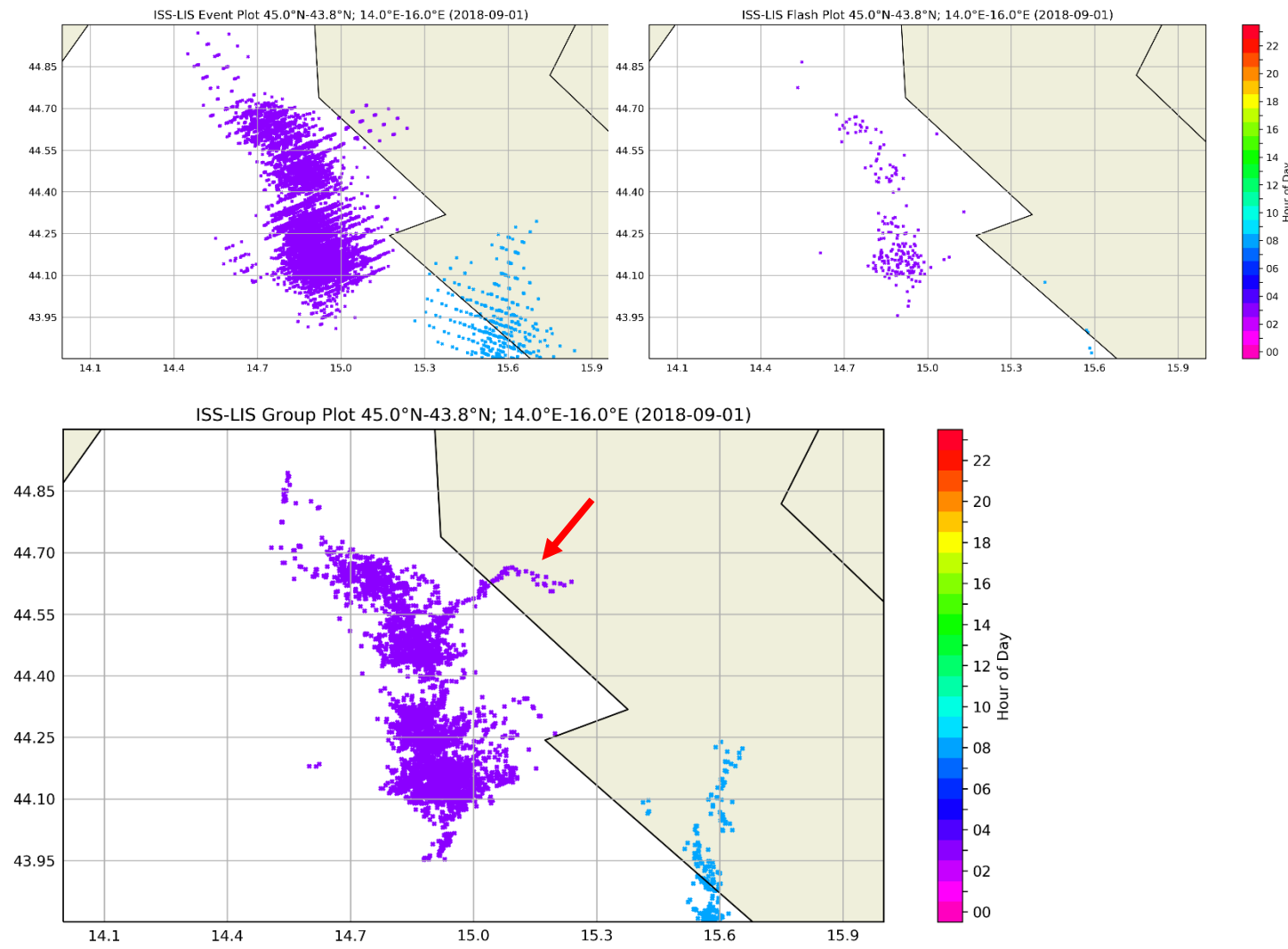
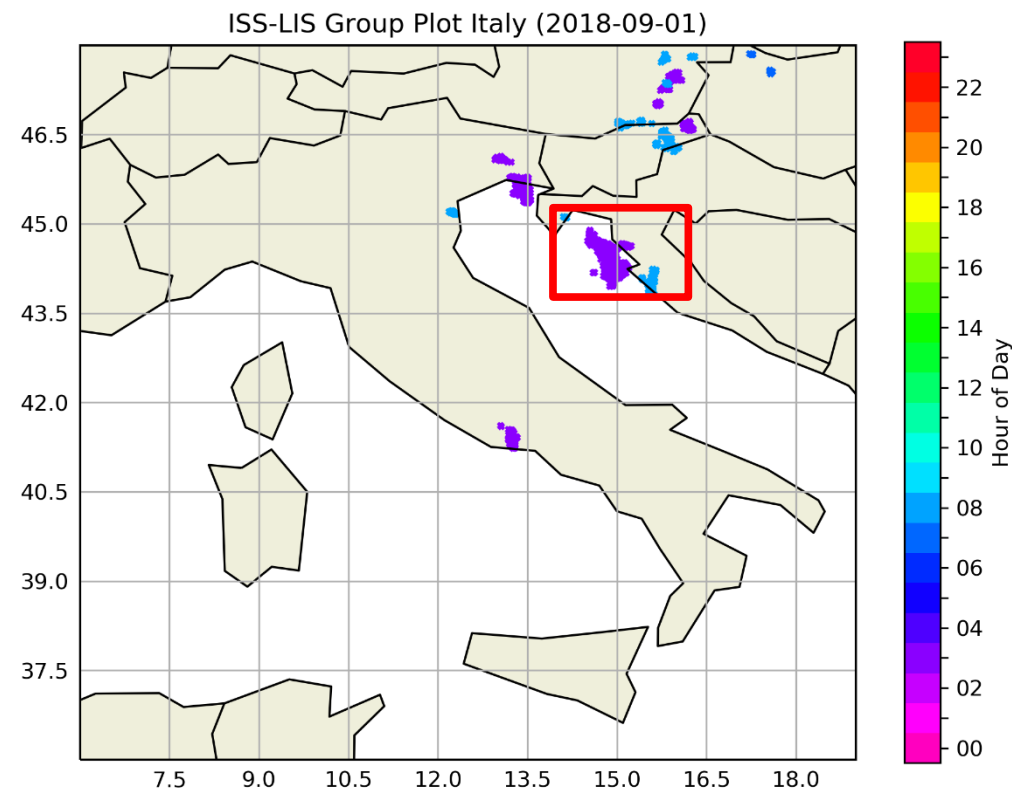
GLD360 Accumulated Strokes Europe 2018-09-03



ISS-LIS Accumulated Groups Europe 2018-09-03

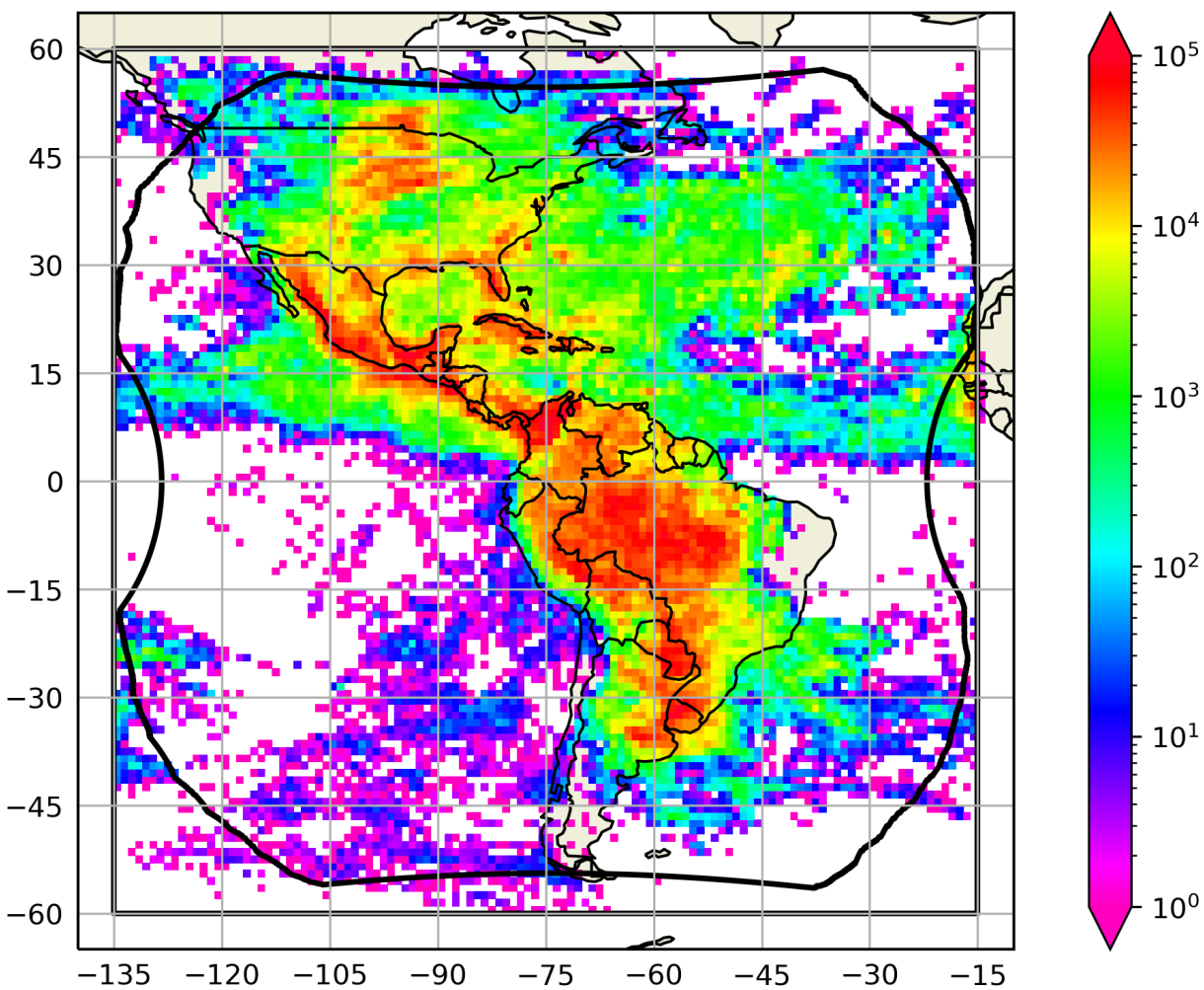


Spotting interesting features in ISS-LIS data

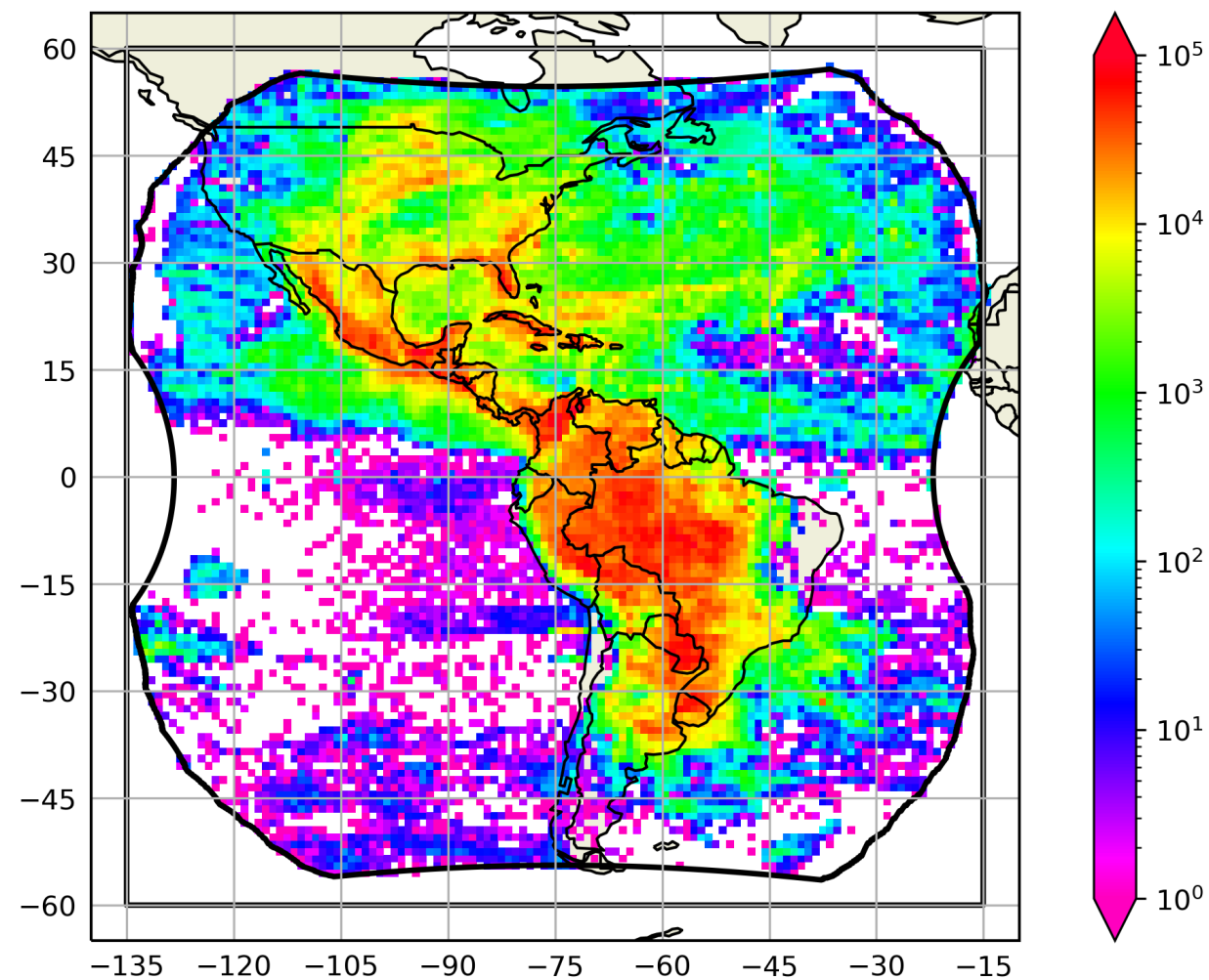


GLD360 and GLM flash accumulation September 2018

GLD360 Accumulated Flashes Gld 2018-09

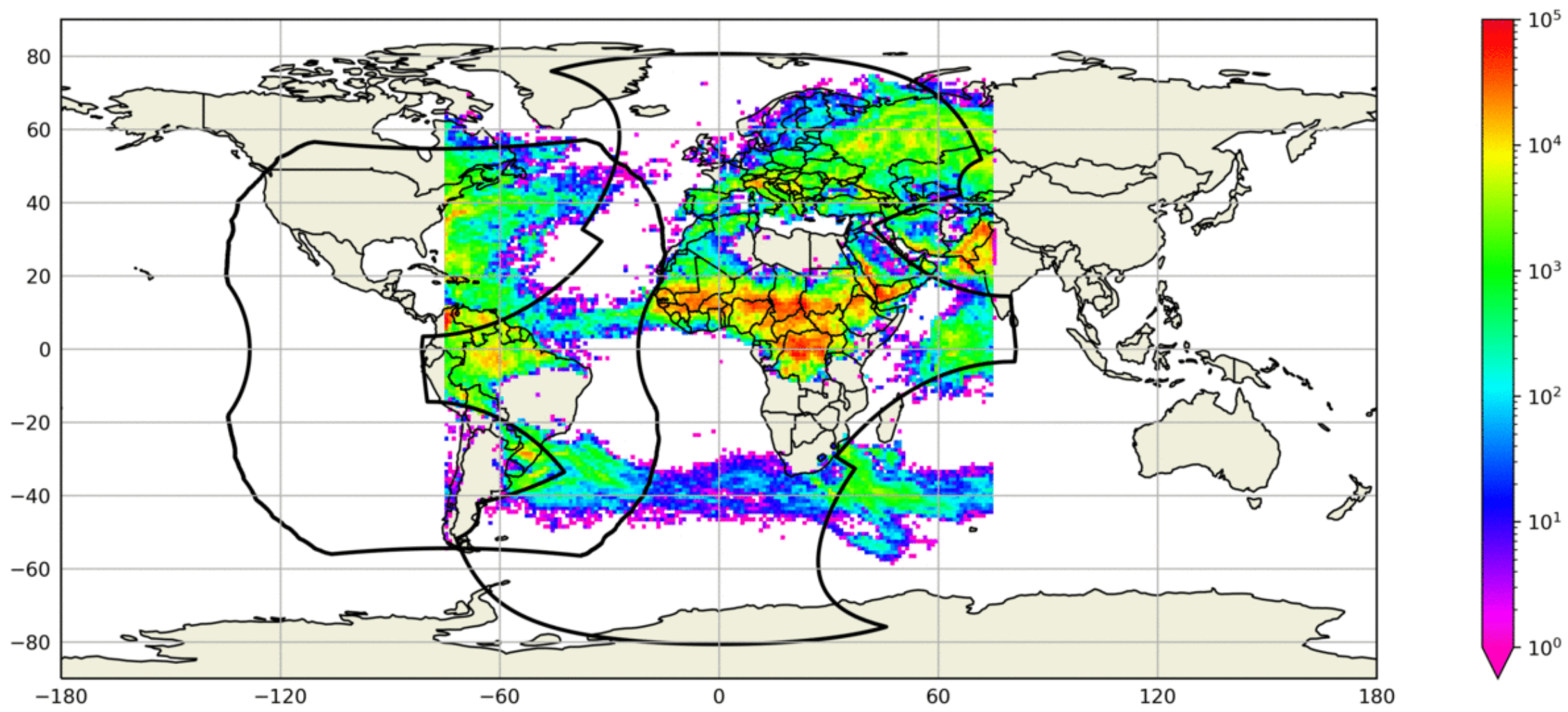


GLM Accumulated Flashes Gld 2018-09



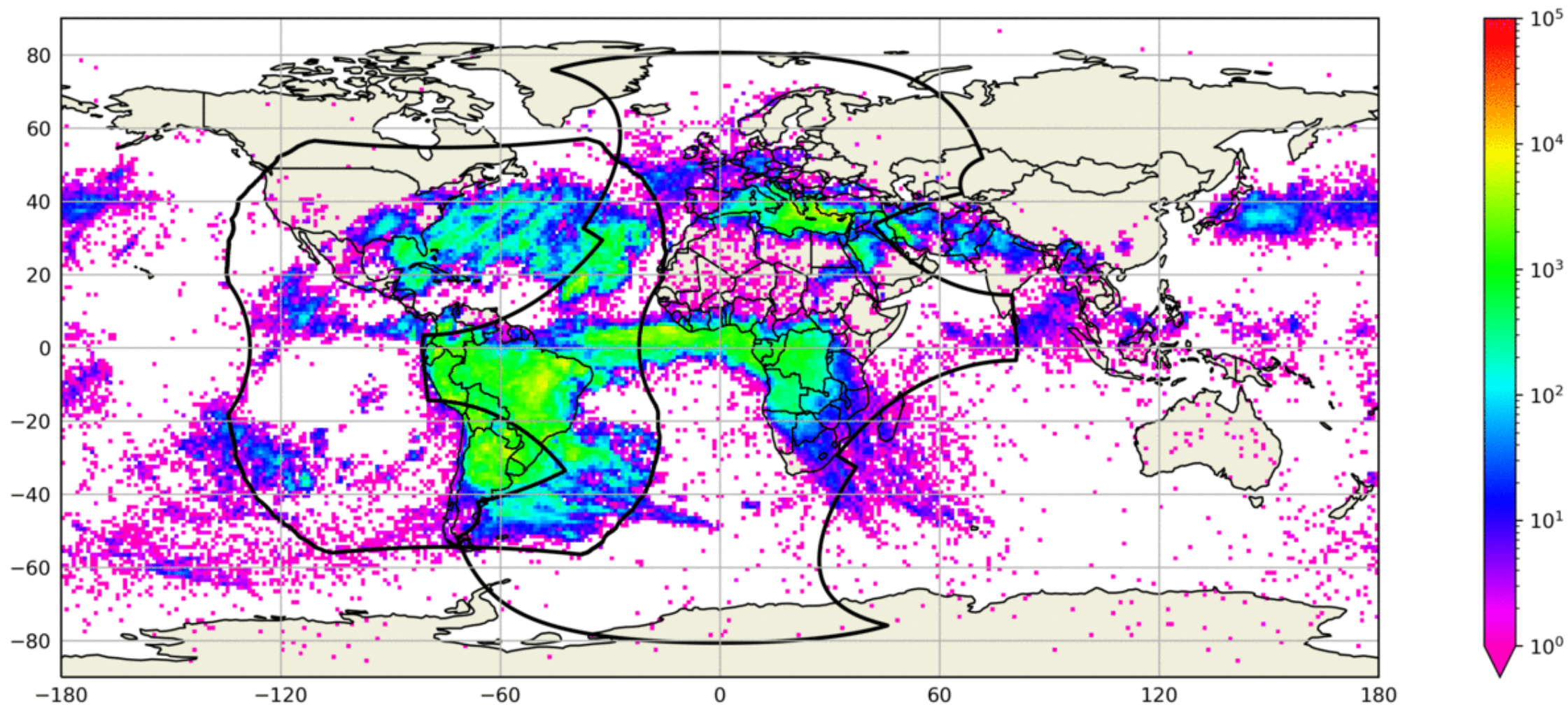
GLD360 monthly flash accumulation July-October 2020

GLD360 Accumulated Flashes Global 2020-07

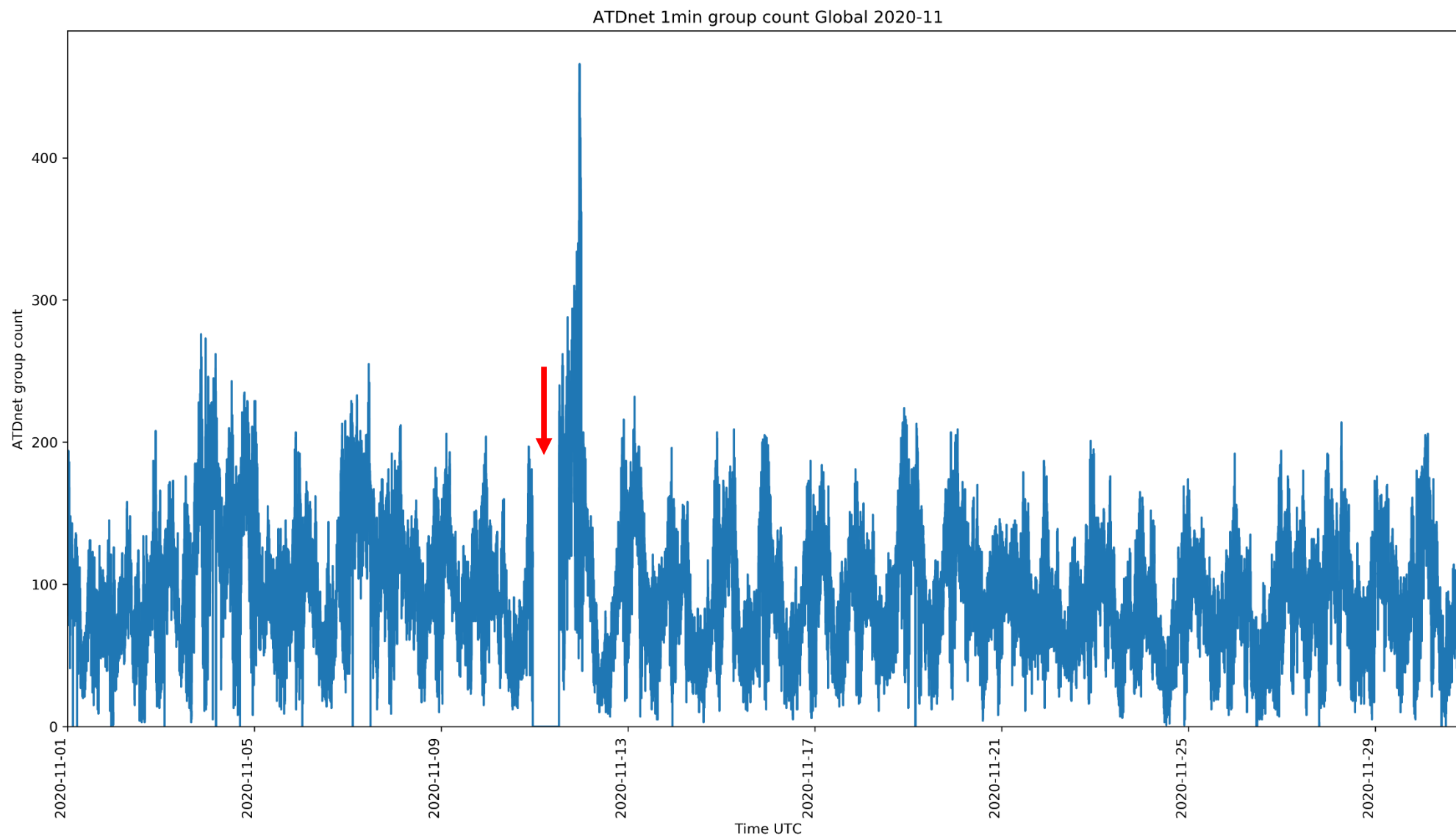


ATDnet monthly stroke accumulation 01-2019...10-2020

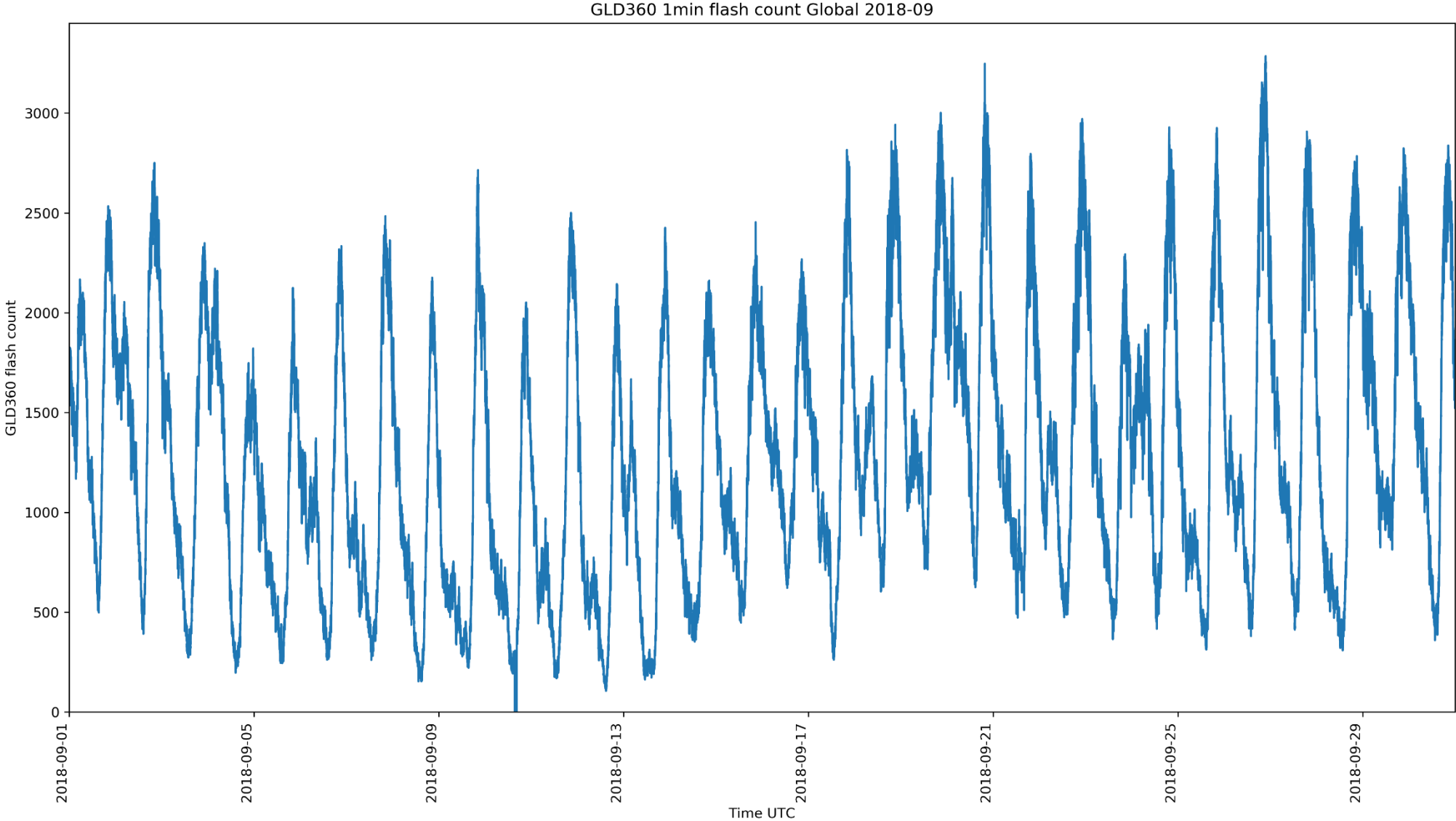
ATDnet Accumulated Strokes Global 2019-01



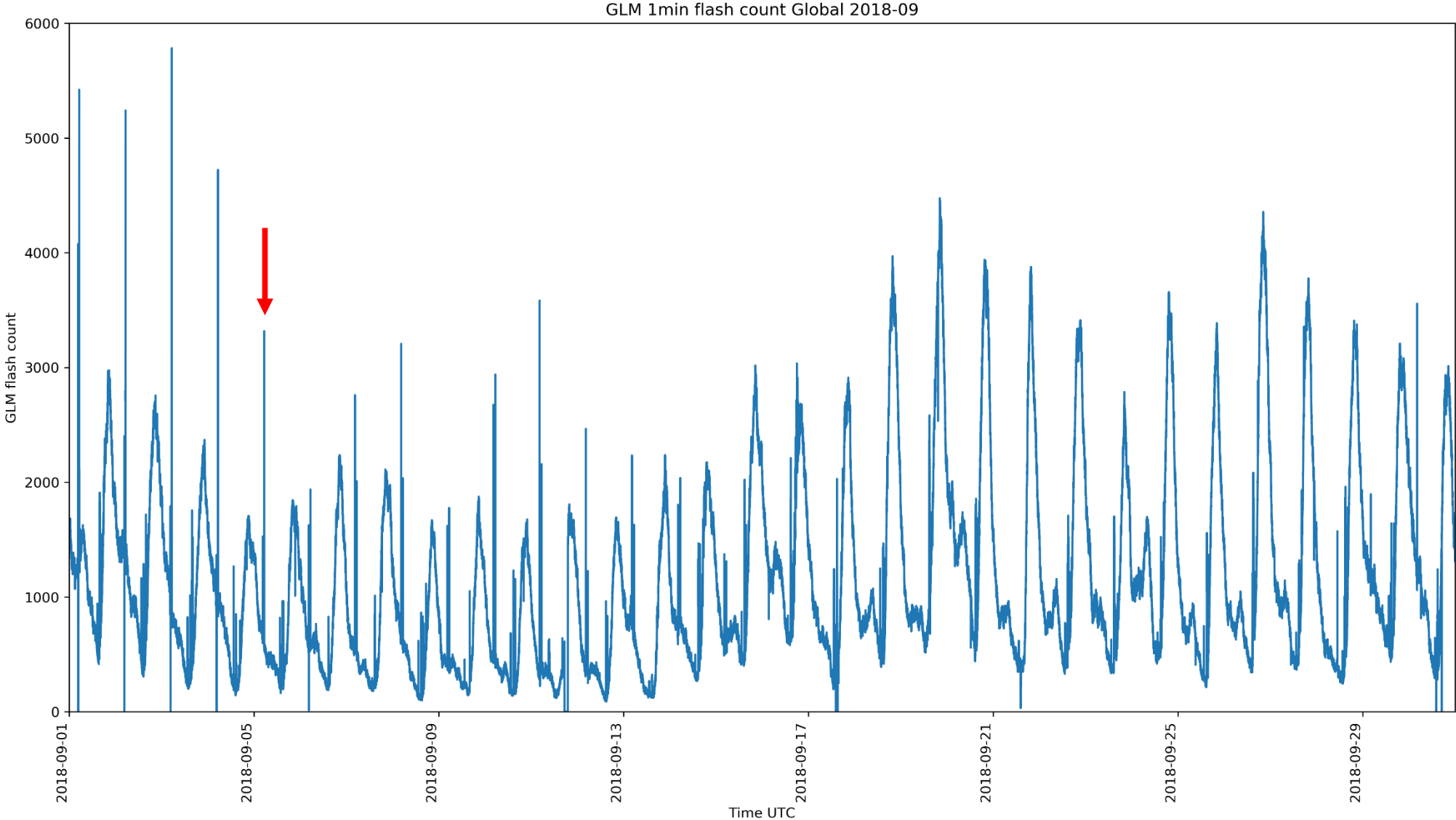
ATDnet flash counts by minute November 2020



GLD360 flash counts by minute September 2018

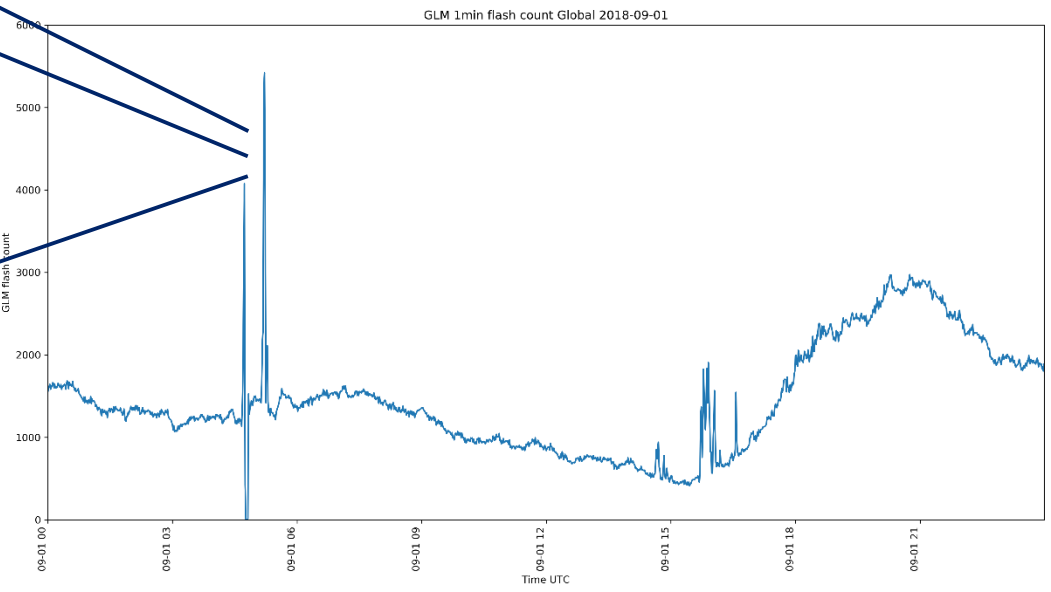
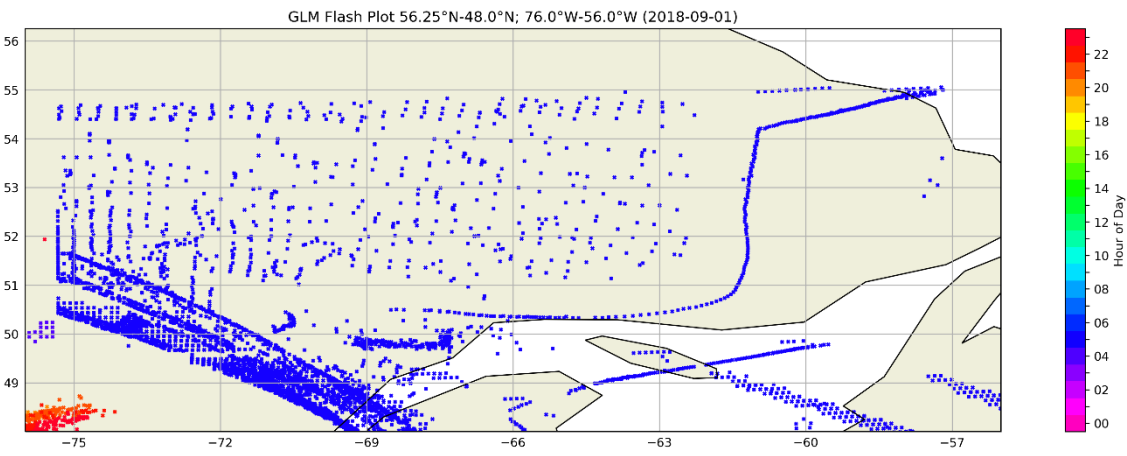
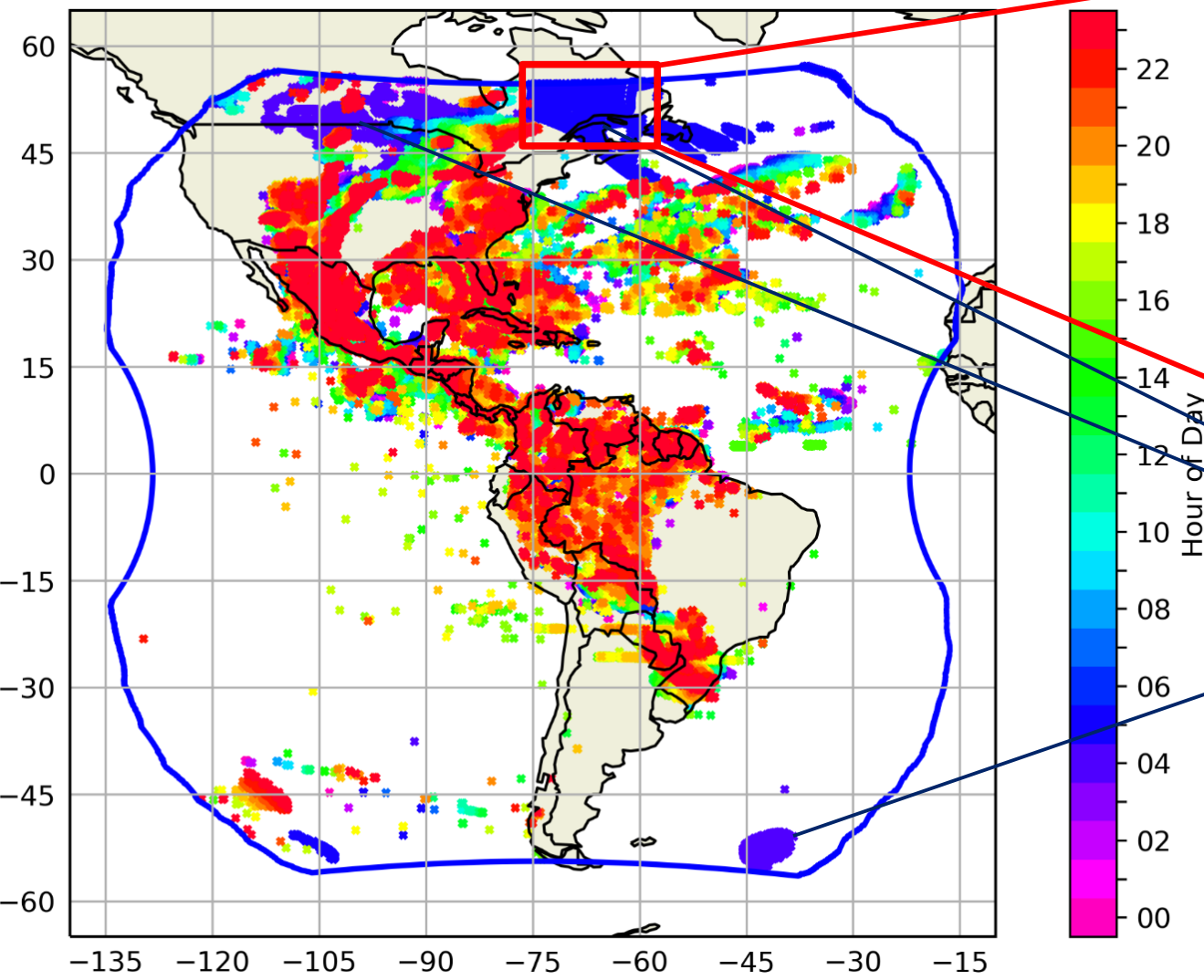


GLM flash counts by minute September 2018



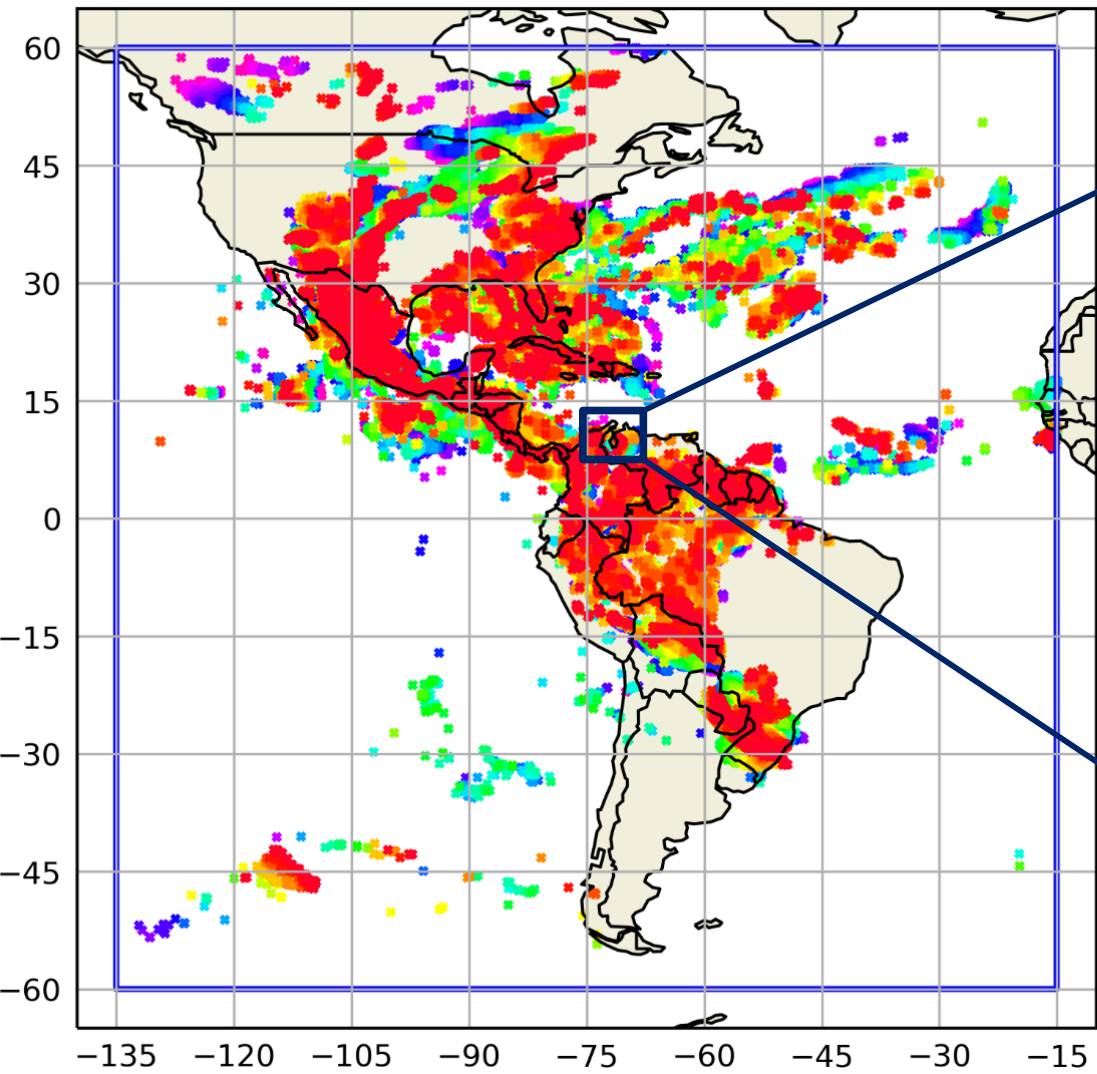
GLM flashes 01-09-2018

GLM Flash Plot Gld (2018-09-01)

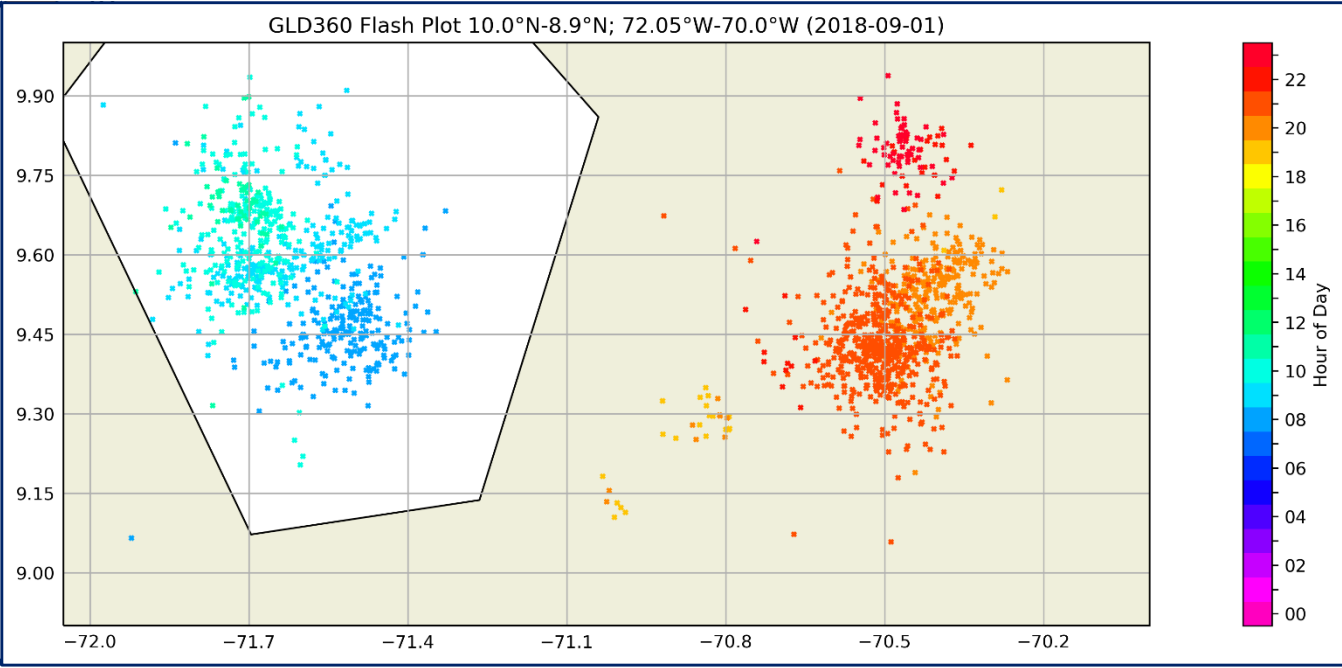


Space vs ground based differences at high zoom level

GLD360 Flash Plot Gld (2018-09-01)

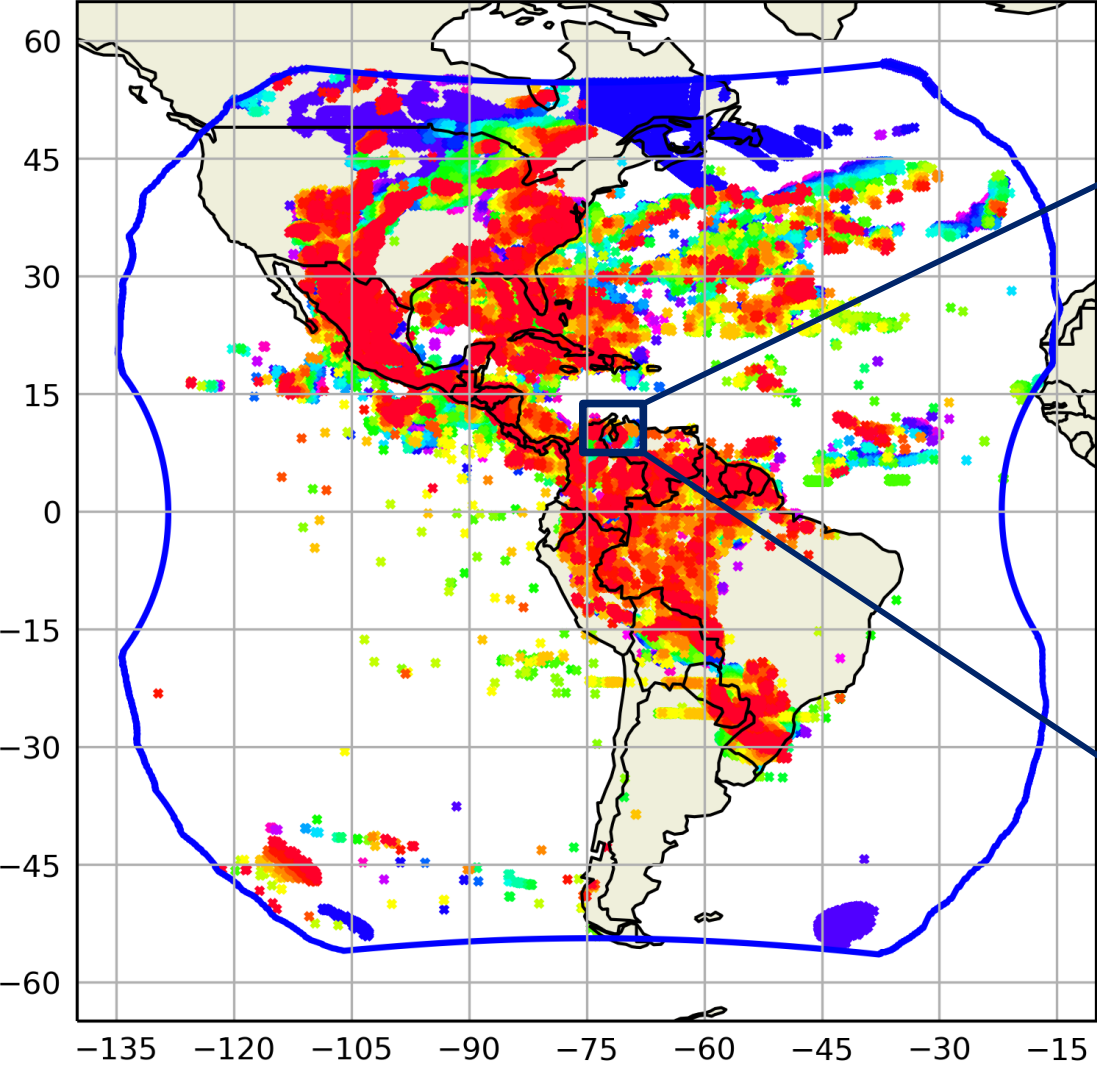


GLD360 Flash Plot 10.0°N-8.9°N; 72.05°W-70.0°W (2018-09-01)

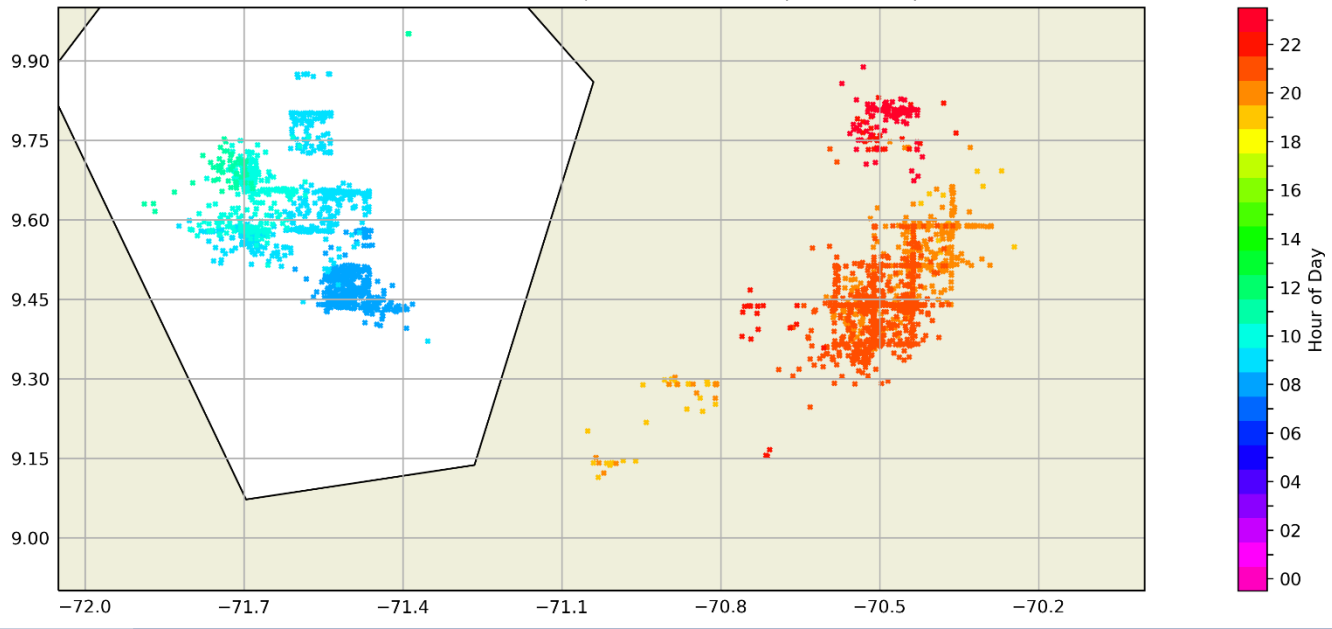


Space vs ground based differences at high zoom level

GLM Flash Plot Gld (2018-09-01)

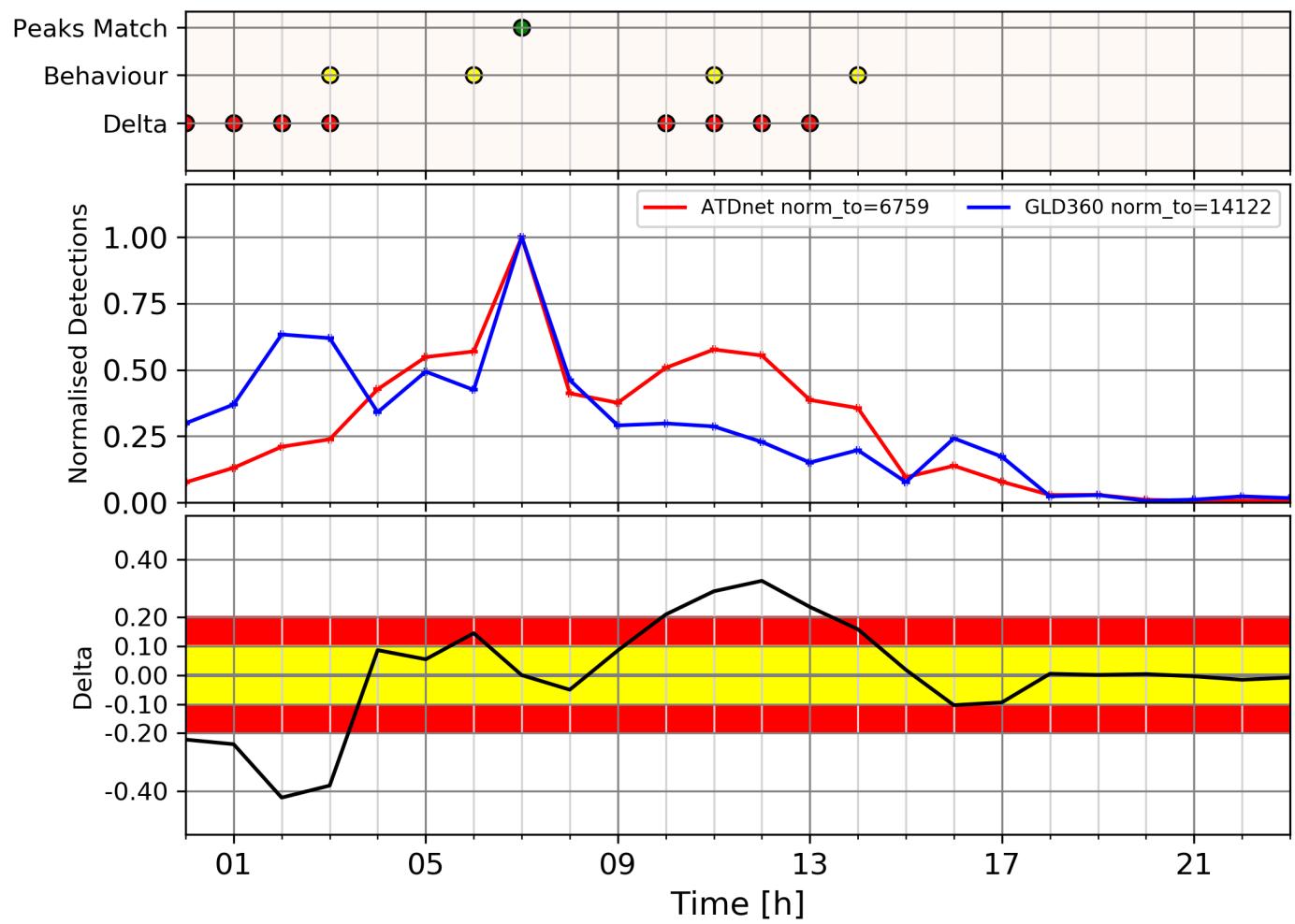


GLM Flash Plot 10.0°N-8.9°N; 72.05°W-70.0°W (2018-09-01)

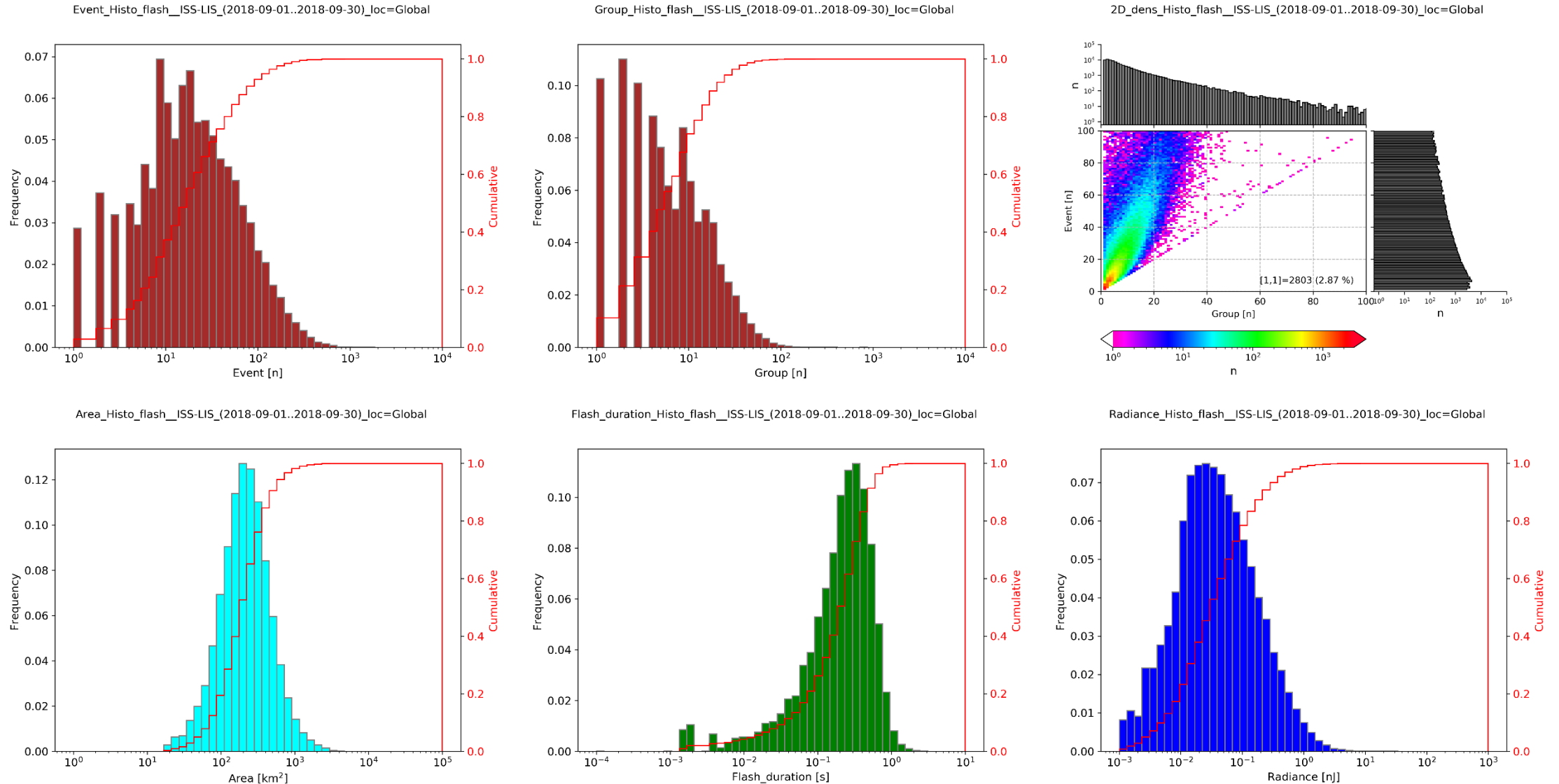


Relative time evolution ATDnet and GLD360 groups

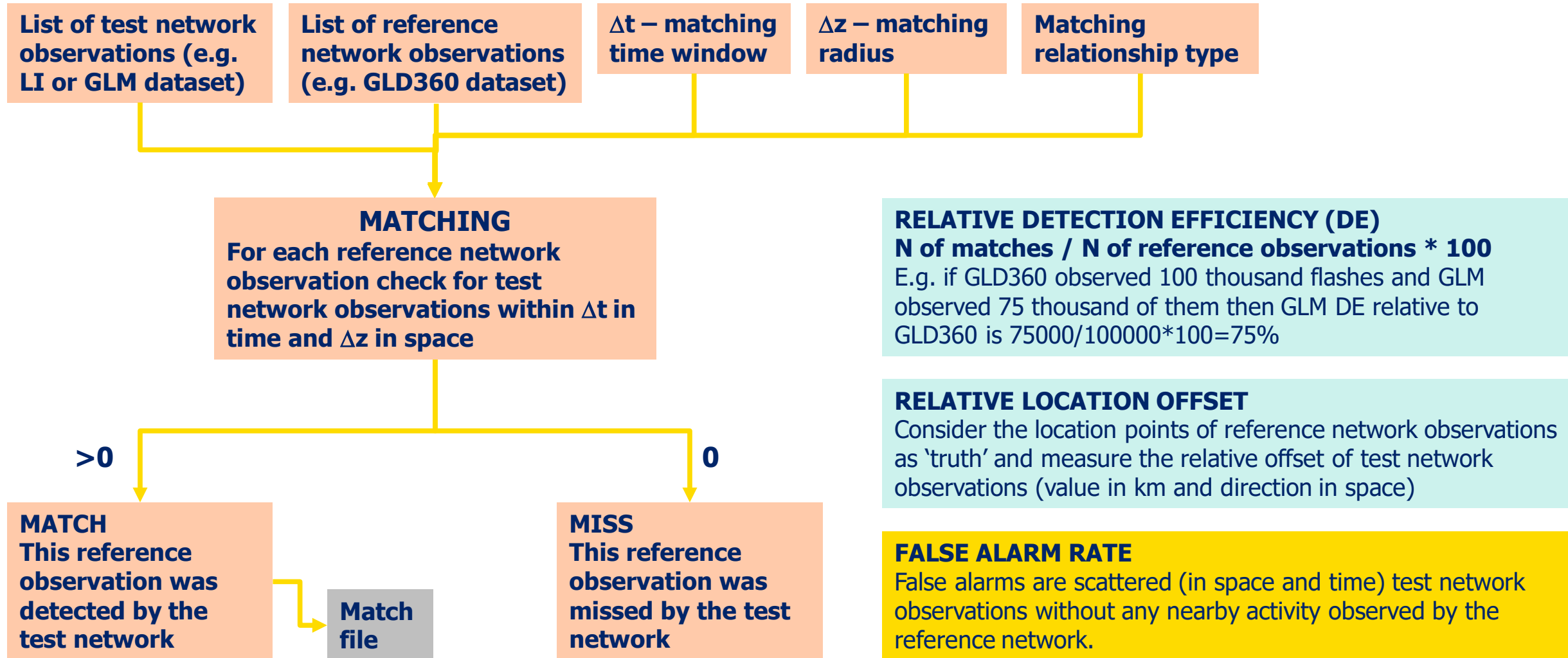
ATDnet GLD360 Tracking Strokes Europe 2020-10-13



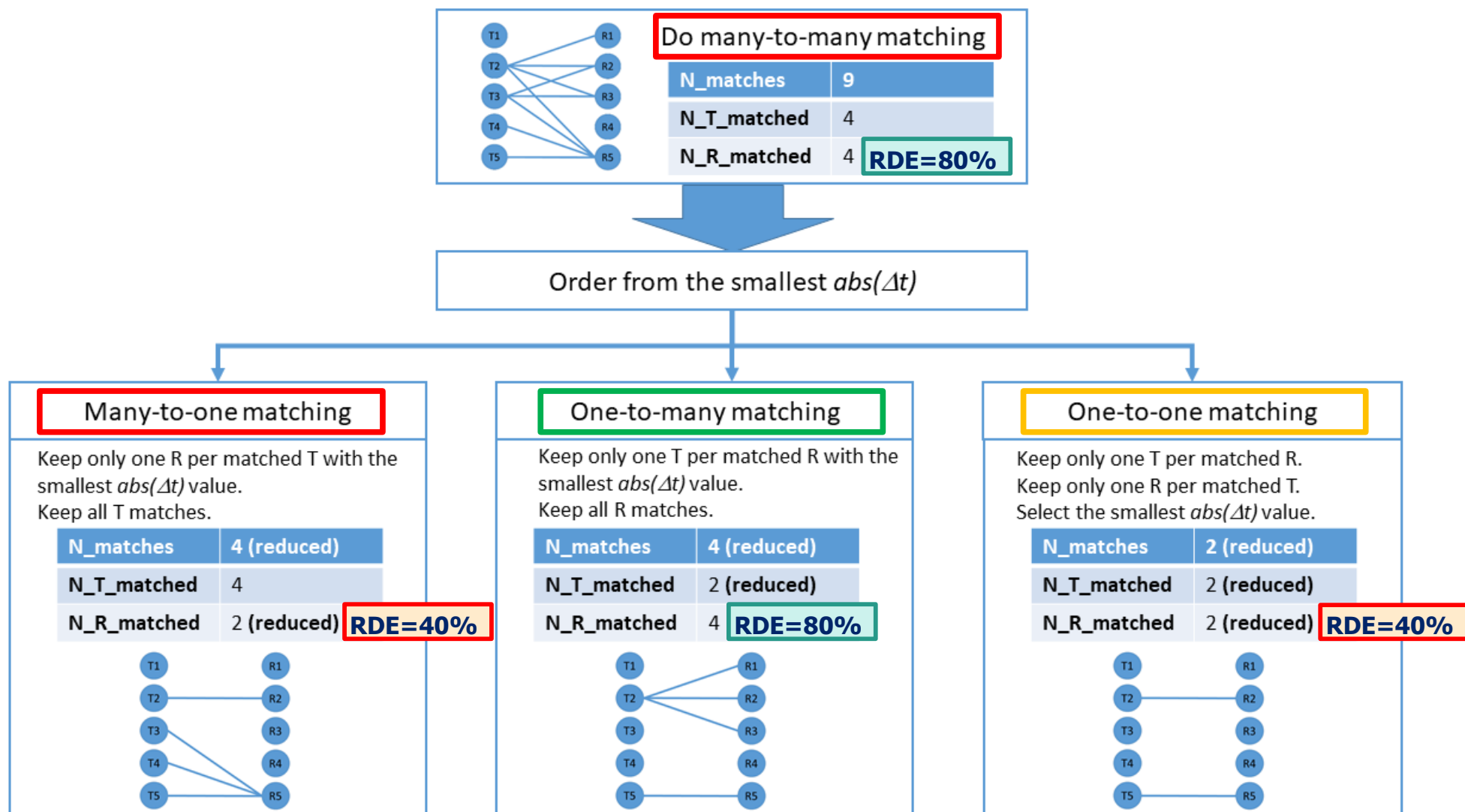
Exploring ISS LIS flash and group properties



LISTAR matching and comparison of networks



LISTAR matching scheme and relationship types

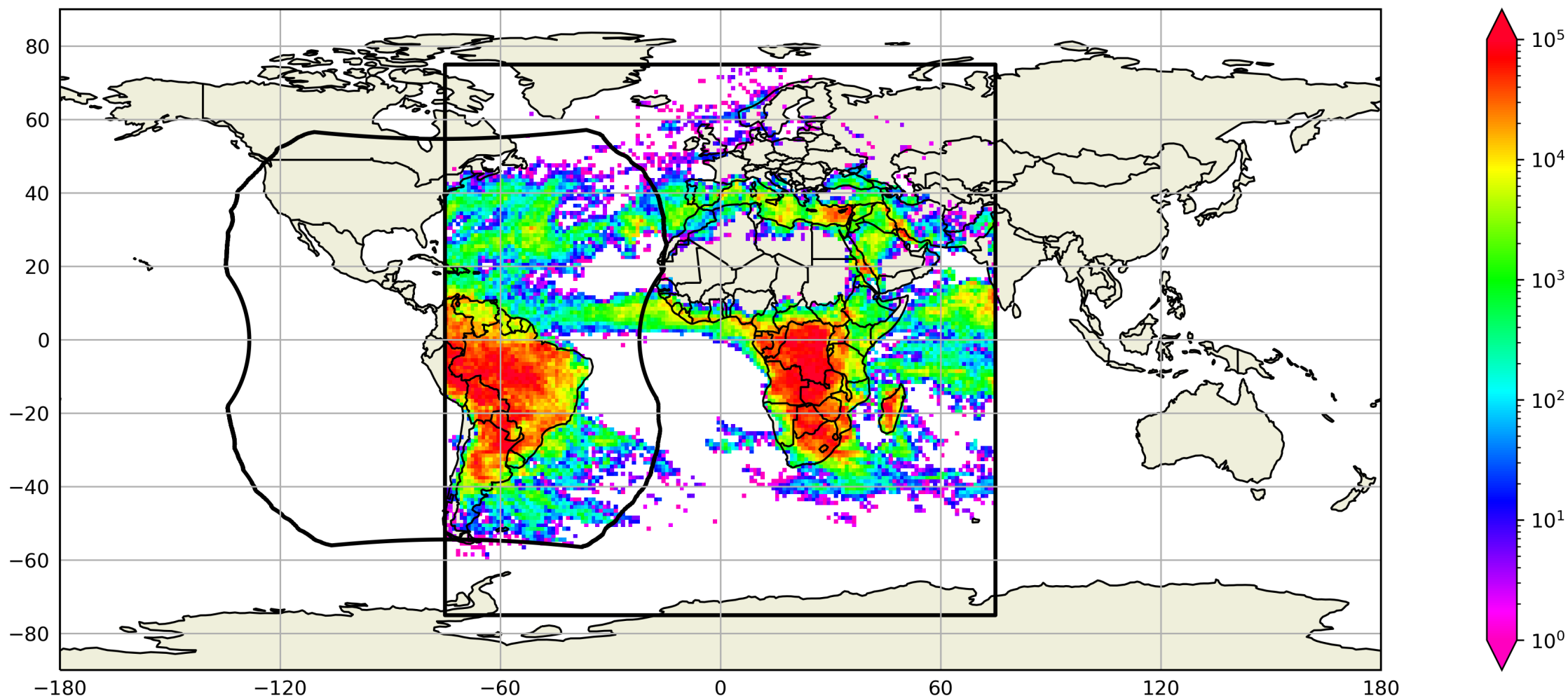


One-to-many vs one-to-one matching

- Matters if the reference network observes much more, e.g. GLD strokes against GLM groups:
 - One-to-one gives **notably better** time distribution and slightly better distance distribution.
 - One-to-one gives **much lower** relative DE.
- Much smaller impact if the reference network observes much less, e.g. GLM groups against GLD strokes:
 - One-to-one can give slightly better time and distance distribution.
 - One-to-one can give slightly lower DE.
- Now showing some GLD-GLM examples.

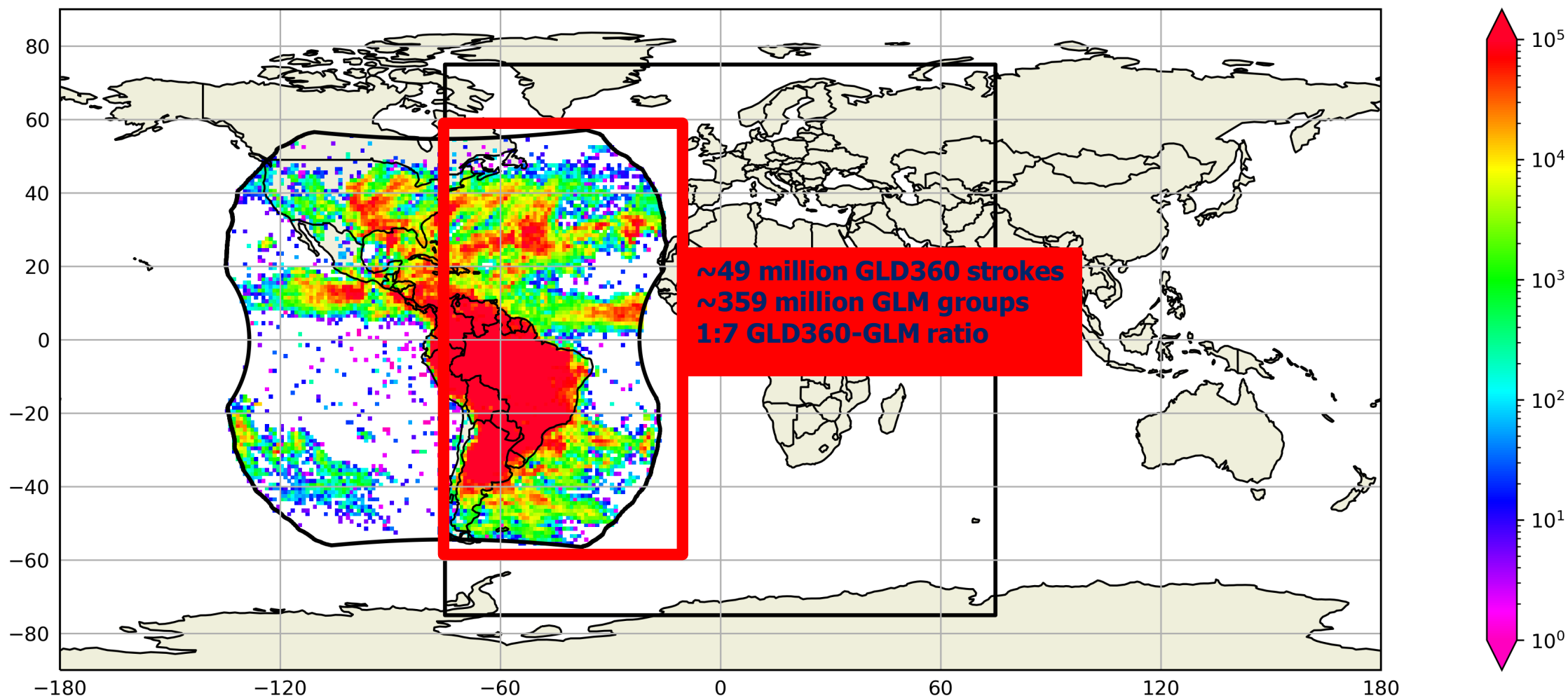
GLD360 monthly stroke accumulation 11/2020

GLD360 Accumulated Strokes Global 2020-11
~96 million GLD360 strokes/pulses



GLM monthly group accumulation 11/2020

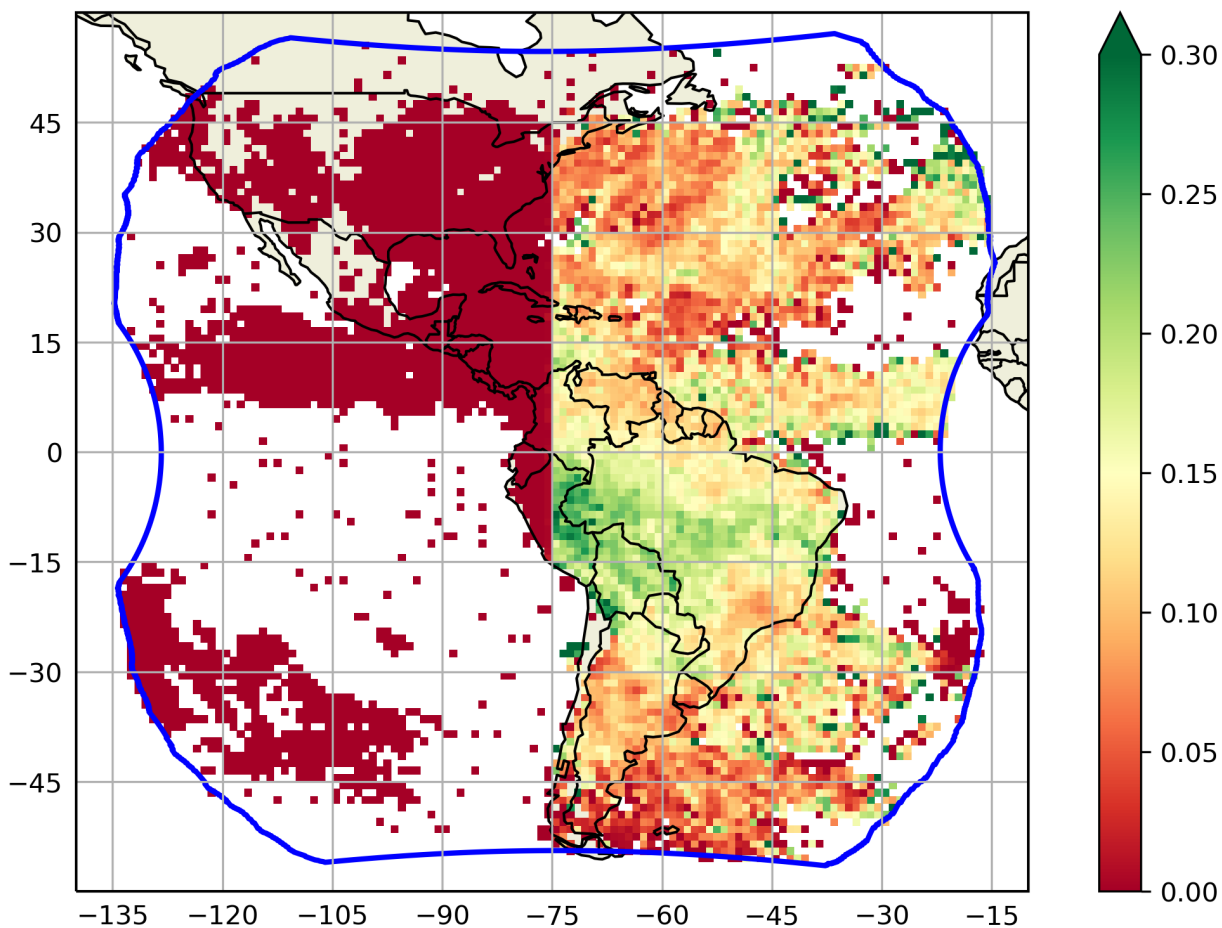
GLM Accumulated Groups Global 2020-11
~380 million GLM groups



GLD-GLM 11/2020 one-to-many matching

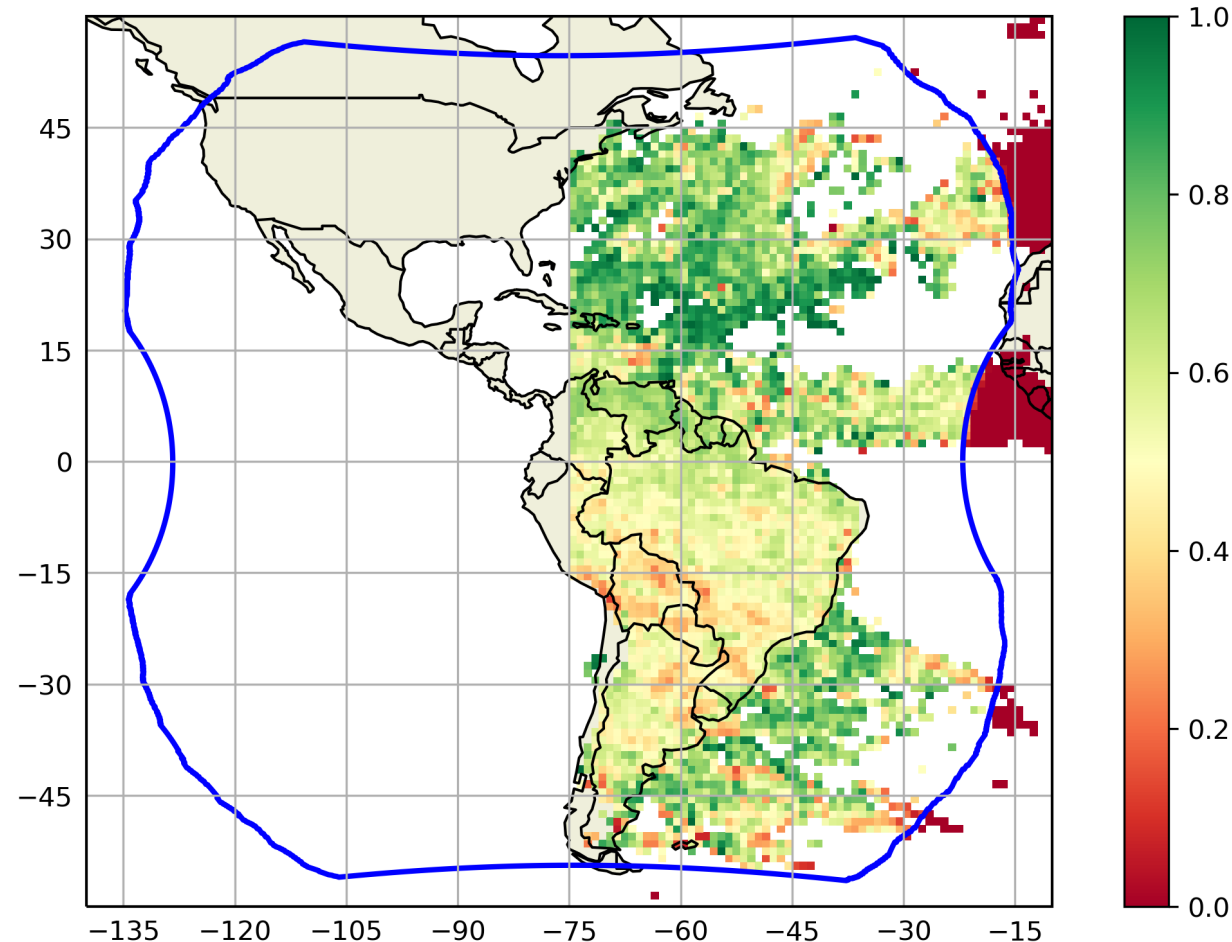
P(GLD360 | GLM) Strokes Americas 2020-11

$\Delta t_b=10\text{ms}$, $\Delta t_a=5\text{ms}$, $\Delta z=50\text{ km}$, 1-to-many matching



P(GLM | GLD360) Groups Americas 2020-11

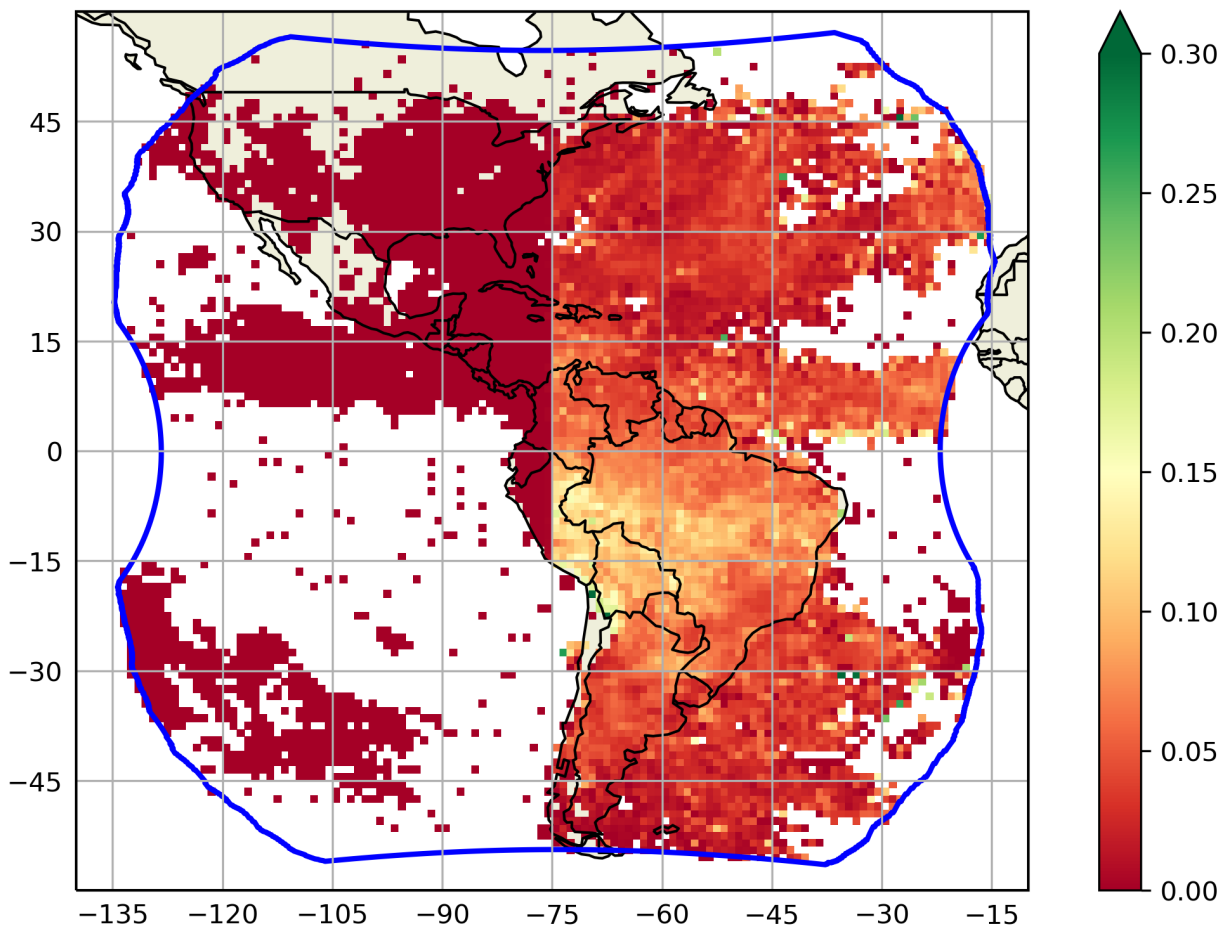
$\Delta t_b=5\text{ms}$, $\Delta t_a=10\text{ms}$, $\Delta z=50\text{ km}$, 1-to-many matching



GLD-GLM 11/2020 one-to-one matching

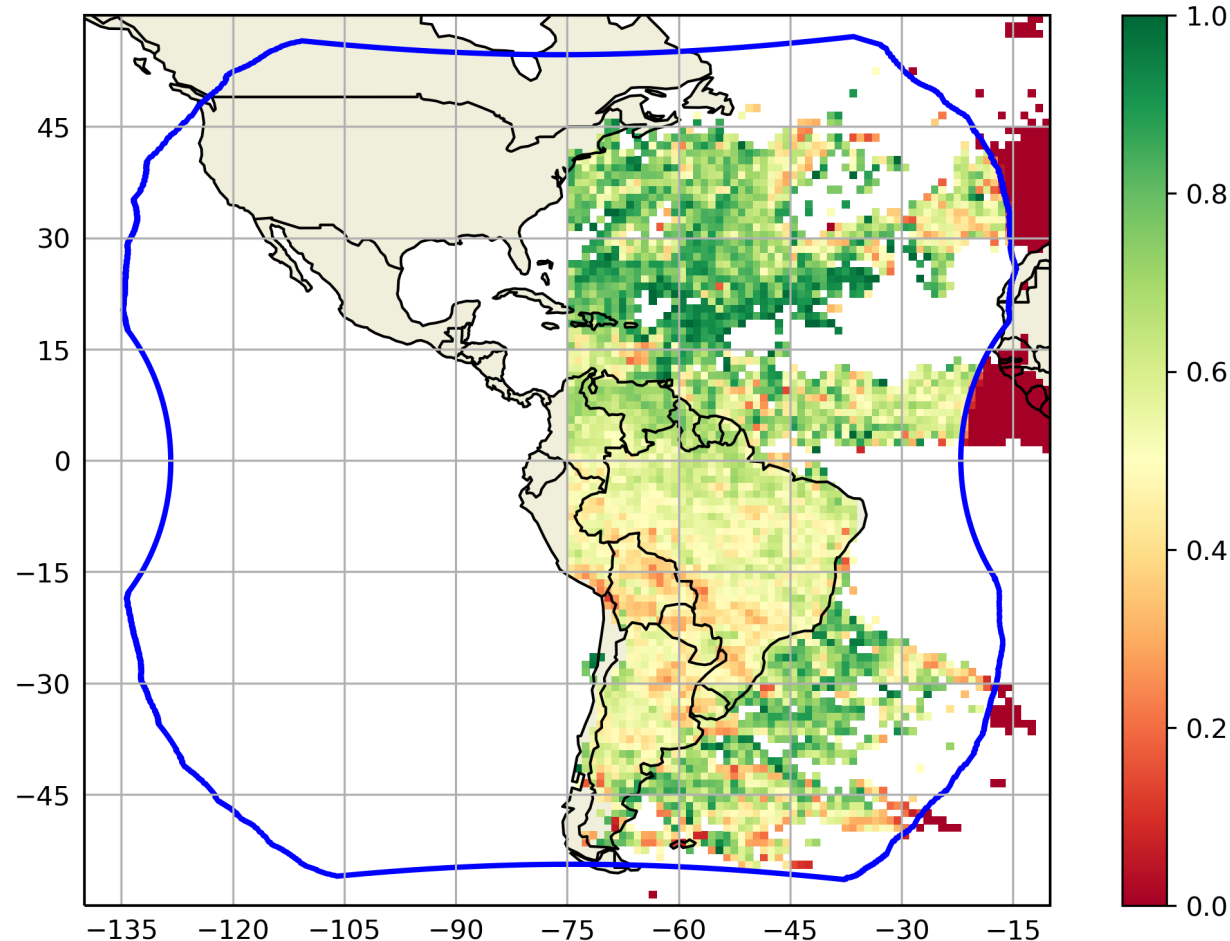
P(GLD360 | GLM) Strokes Americas 2020-11

$\Delta t_b=10\text{ms}$, $\Delta t_a=5\text{ms}$, $\Delta z=50\text{ km}$, 1-to-1 matching



P(GLM | GLD360) Groups Americas 2020-11

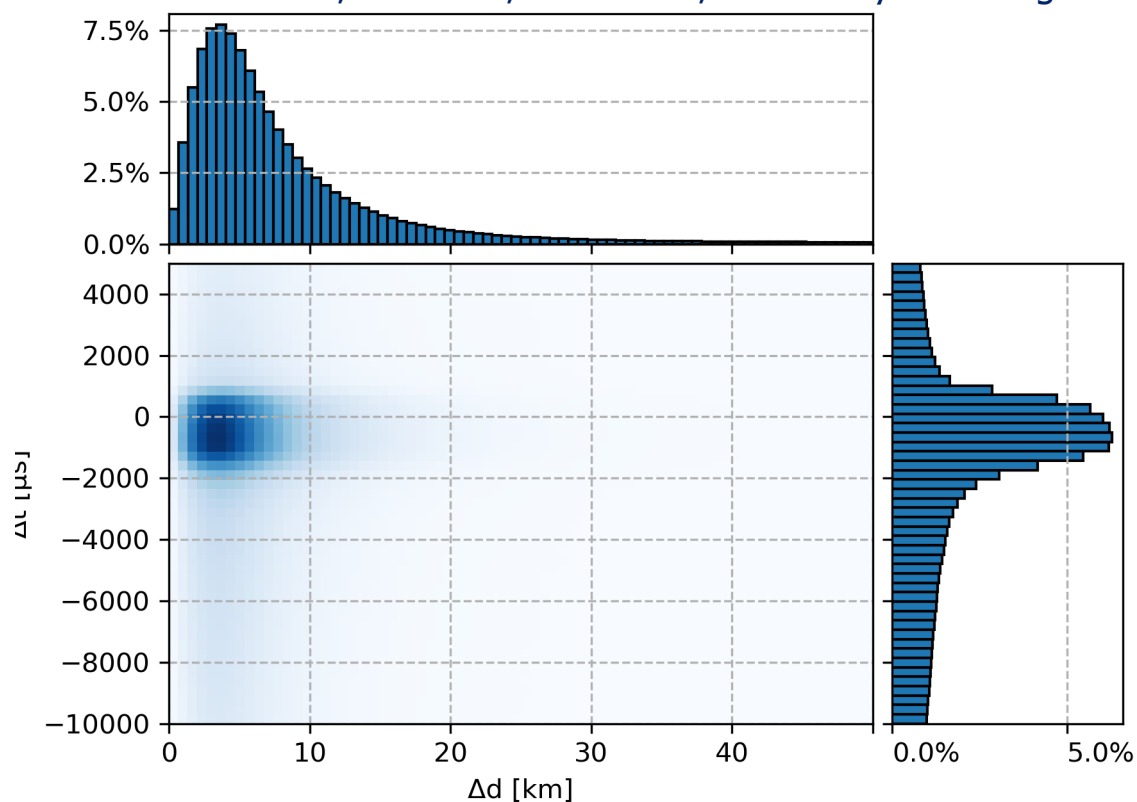
$\Delta t_b=5\text{ms}$, $\Delta t_a=10\text{ms}$, $\Delta z=50\text{ km}$, 1-to-1 matching



GLD-GLM 11/2020 one-to-many matching

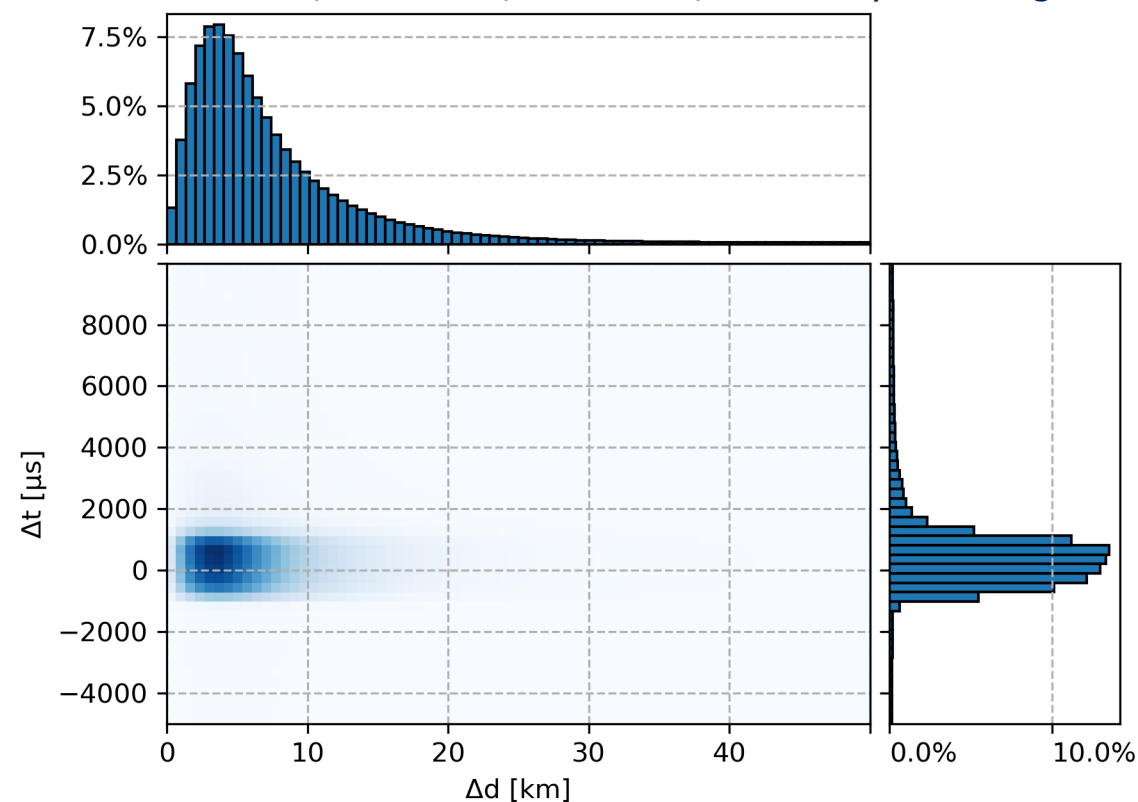
Spatio-Temporal Accuracy GLD360 vs. GLM Strokes Global 2020-11

$\Delta t_b = 10\text{ms}$, $\Delta t_a = 5\text{ms}$, $\Delta z = 50\text{ km}$, 1-to-many matching



Spatio-Temporal Accuracy GLM vs. GLD360 Groups Global 2020-11

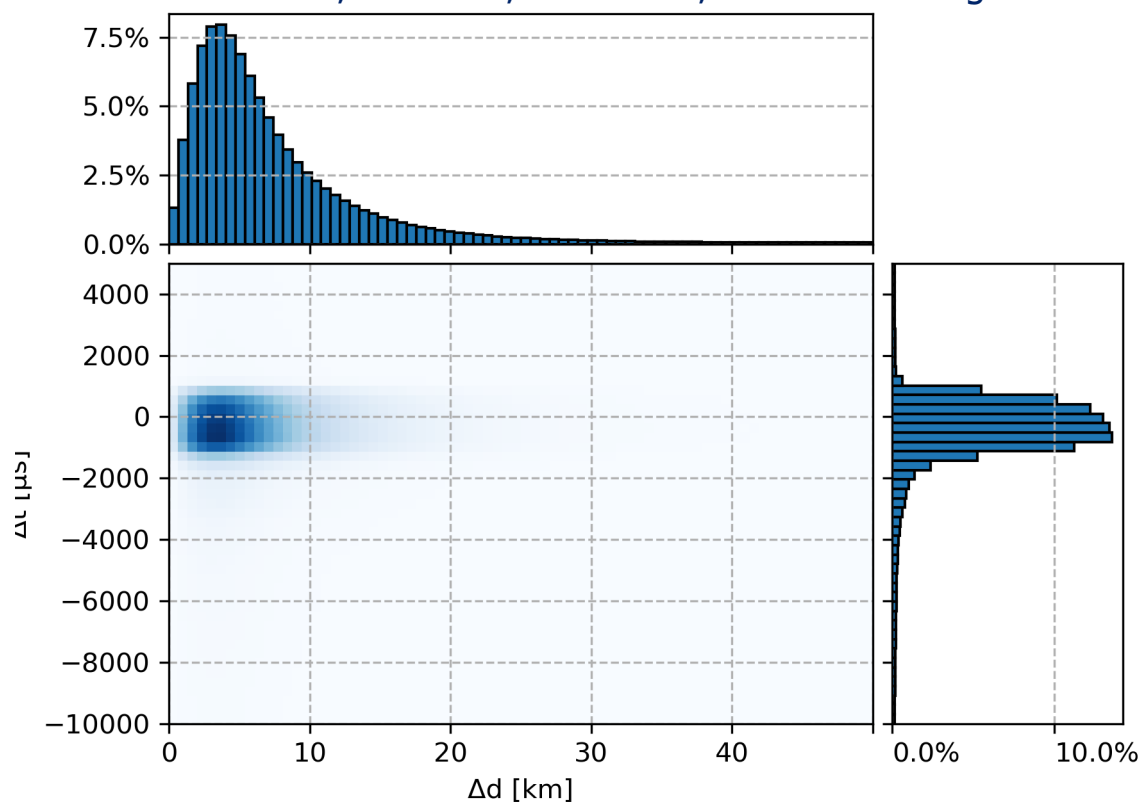
$\Delta t_b = 5\text{ms}$, $\Delta t_a = 10\text{ms}$, $\Delta z = 50\text{ km}$, 1-to-many matching



GLD-GLM 11/2020 one-to-one matching

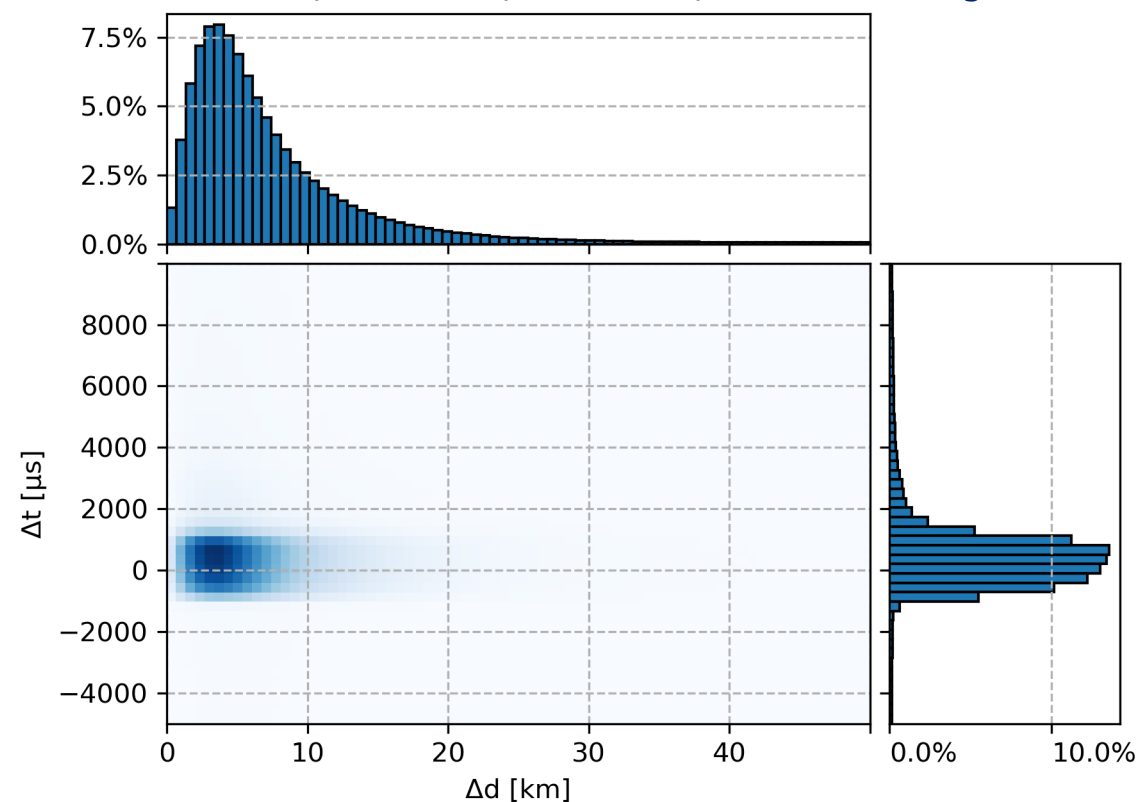
Spatio-Temporal Accuracy GLD360 vs. GLM Strokes Global 2020-11

$\Delta t_b = 10\text{ms}$, $\Delta t_a = 5\text{ms}$, $\Delta z = 50\text{ km}$, 1-to-1 matching



Spatio-Temporal Accuracy GLM vs. GLD360 Groups Global 2020-11

$\Delta t_b = 5\text{ms}$, $\Delta t_a = 10\text{ms}$, $\Delta z = 50\text{ km}$, 1-to-1 matching



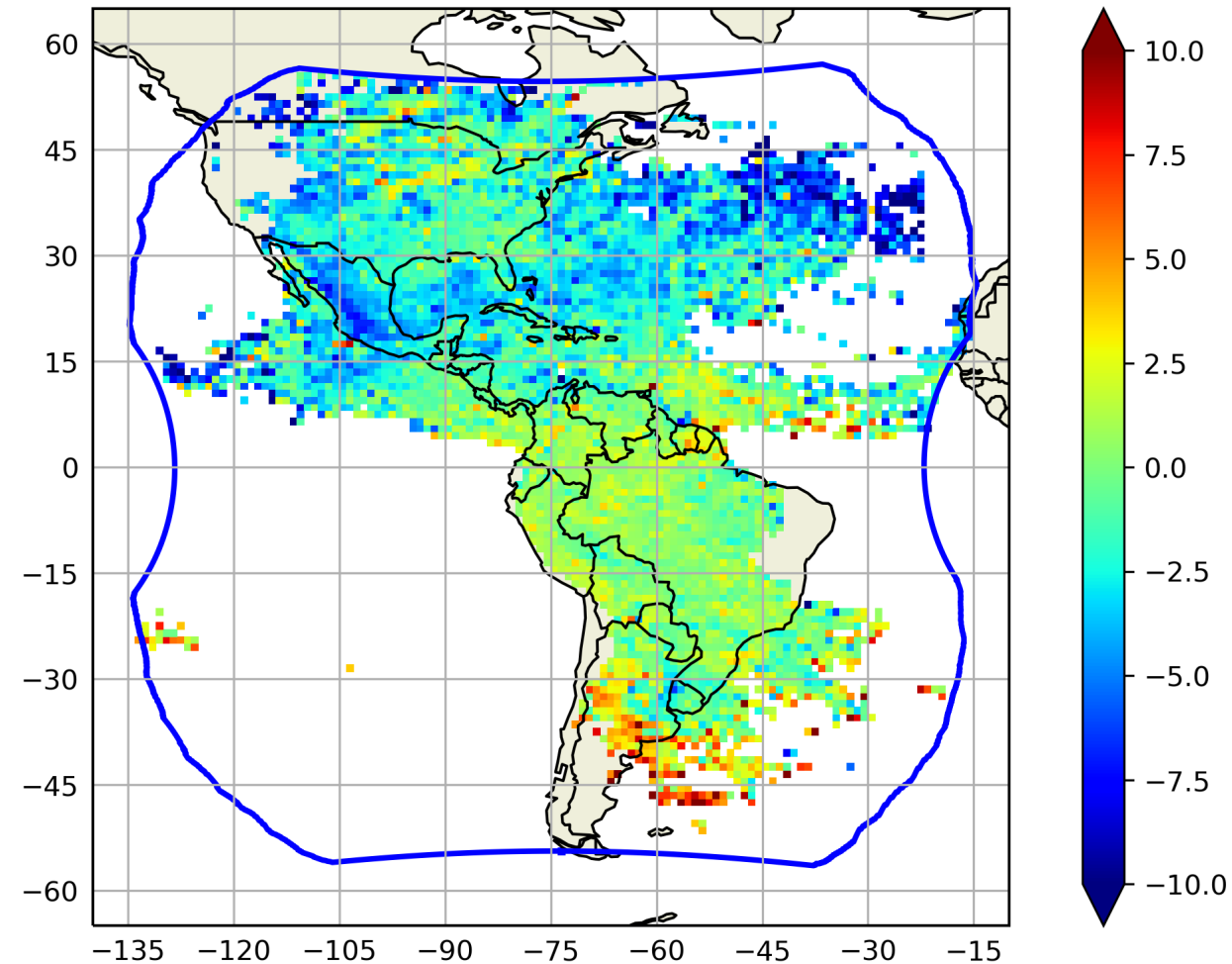
LISTAR matching type conclusion

- Two main possibilities.
- ‘Coarse’ approach:
 - Use one-to many relationship for all, i.e. detection efficiency, location offset and time offset.
- ‘Fine’ approach:
 - Use one to many relationship for detection efficiency.
 - Use one-to-one relationship for location and time offset.
- Any feedback?

GLM group offset relative to GLD360 September 2018

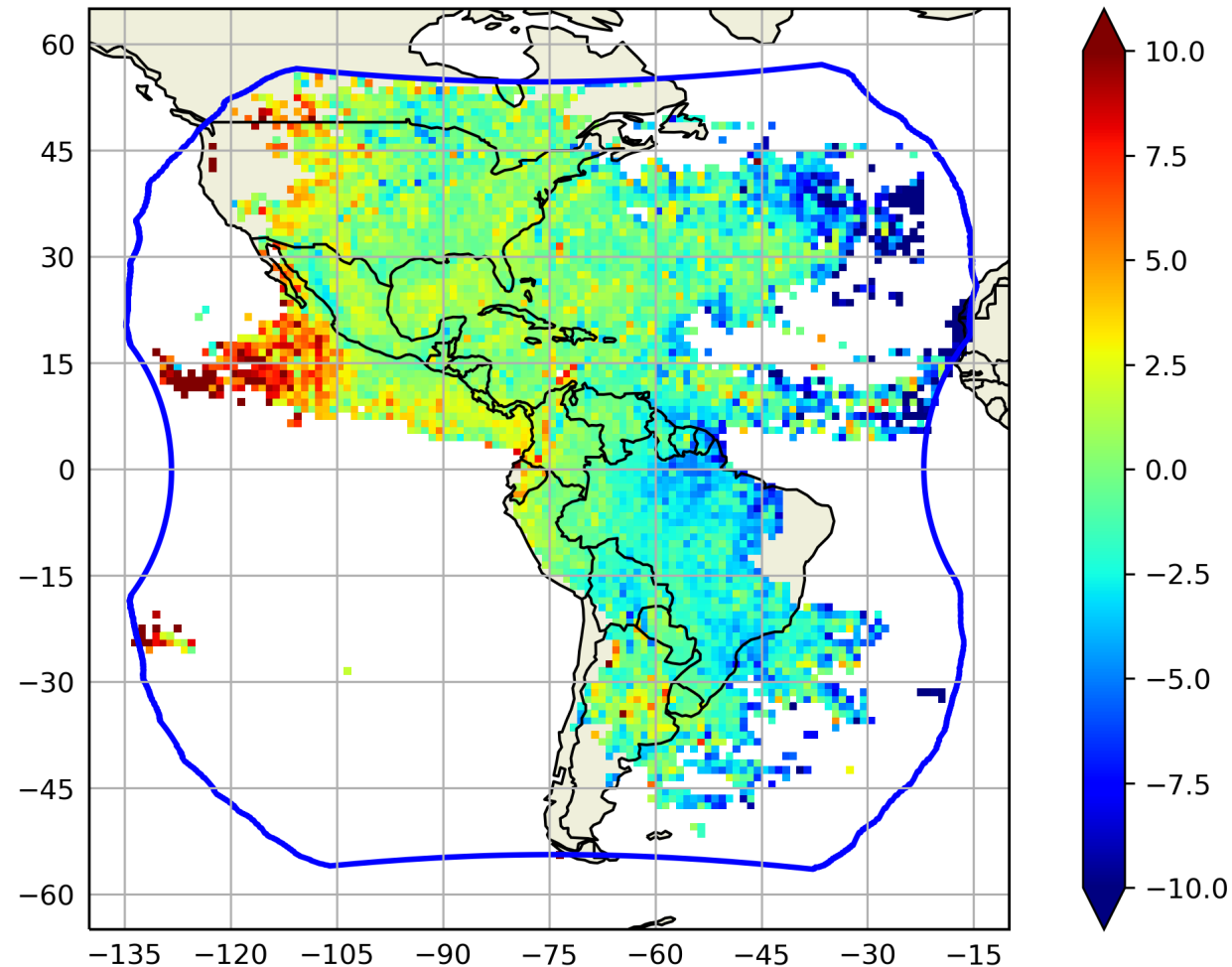
LA NS (GLM-GLD360) Group Gld 2018-09

$\Delta t_b=5\text{ms}$, $\Delta t_a=10\text{ms}$, $\Delta z=50\text{ km}$, 1-to-1 matches



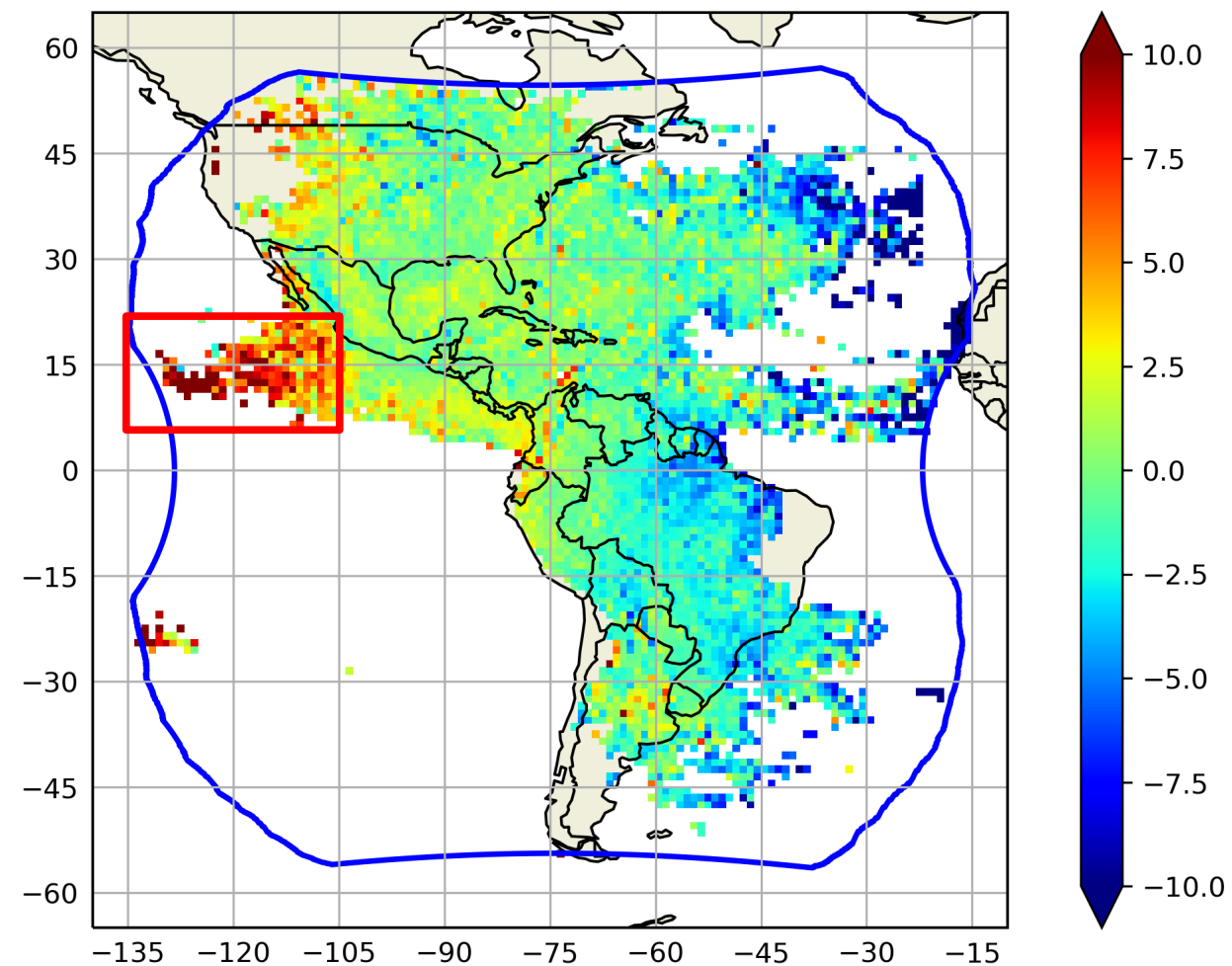
LA WE (GLM-GLD360) Group Gld 2018-09

$\Delta t_b=5\text{ms}$, $\Delta t_a=10\text{ms}$, $\Delta z=50\text{ km}$, 1-to-1 matches

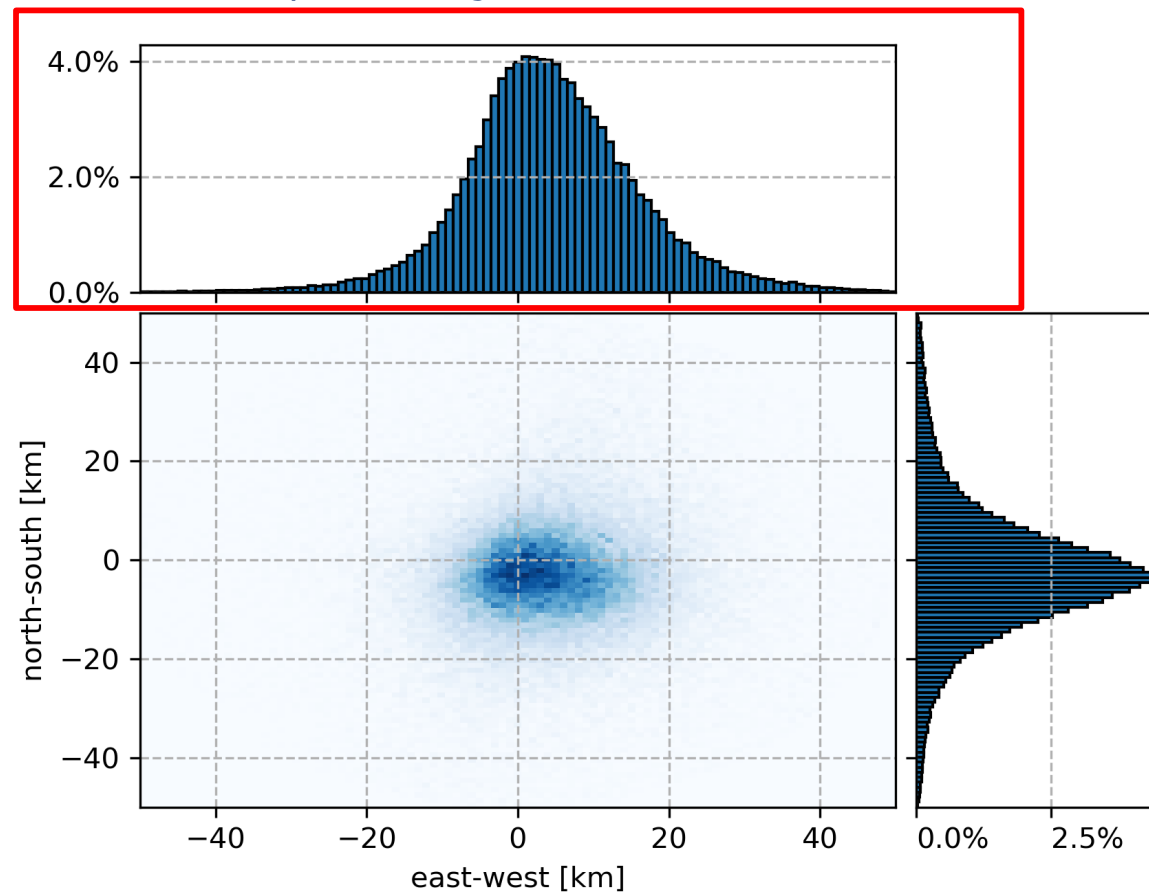


GLM group WE offset September 2018 area I

LA WE (GLM-GLD360) Group Gld 2018-09

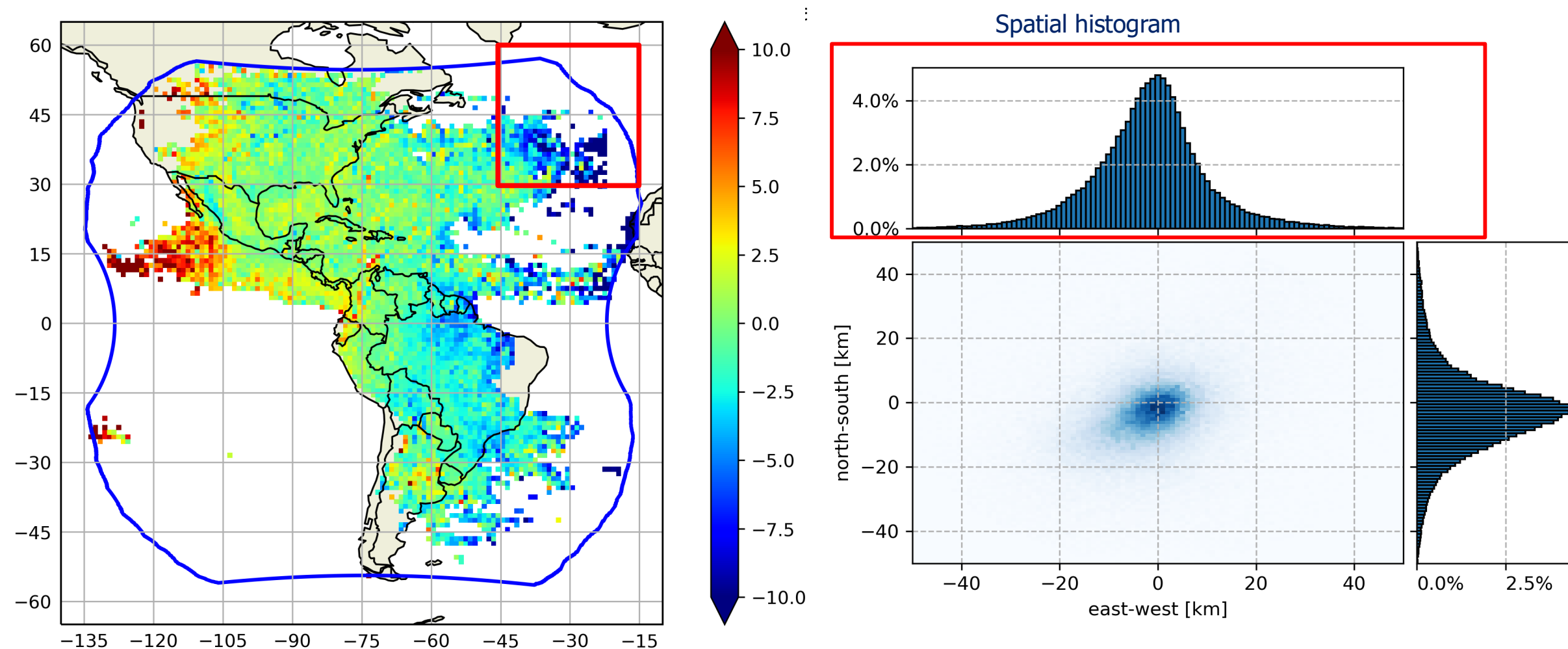


Spatial histogram



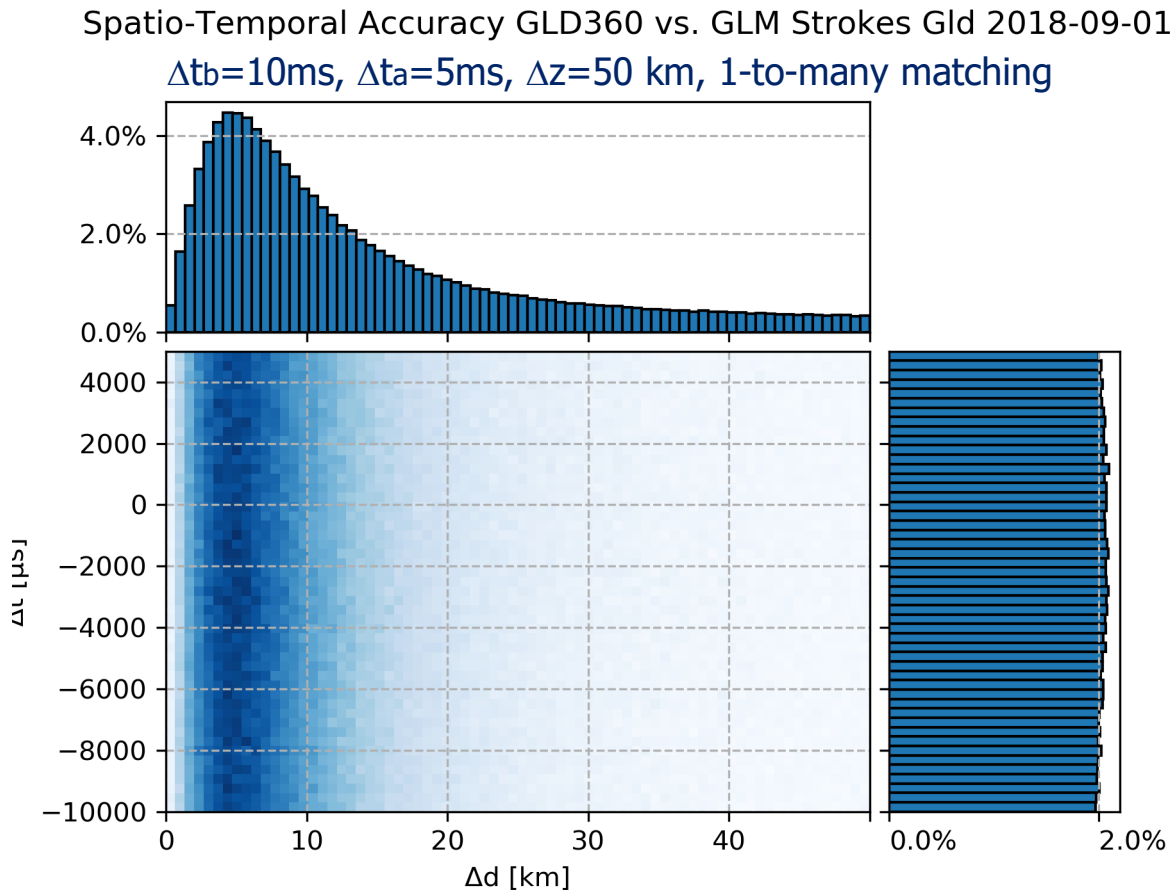
GLM group WE offset September 2018 area II

LA WE (GLM-GLD360) Group Gld 2018-09

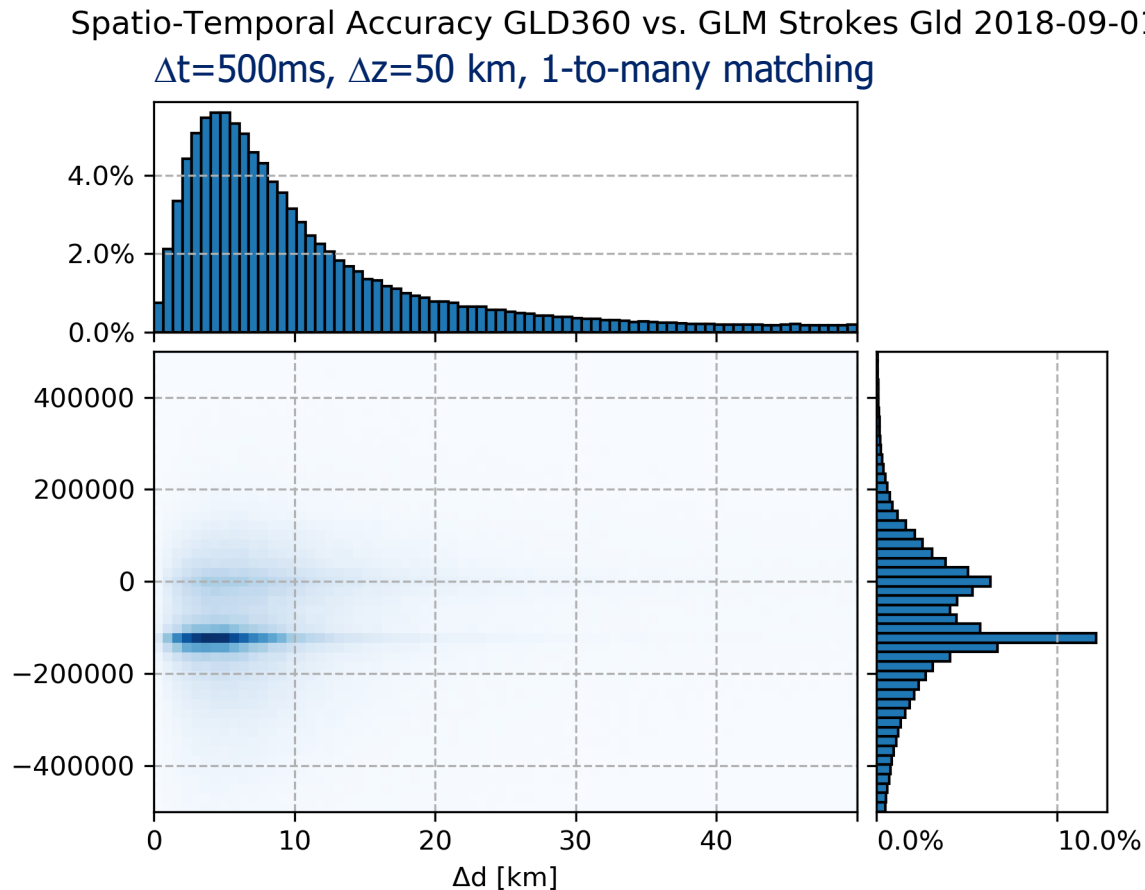


Spatiotemporal histogram in problem investigation

What if facing something like that?

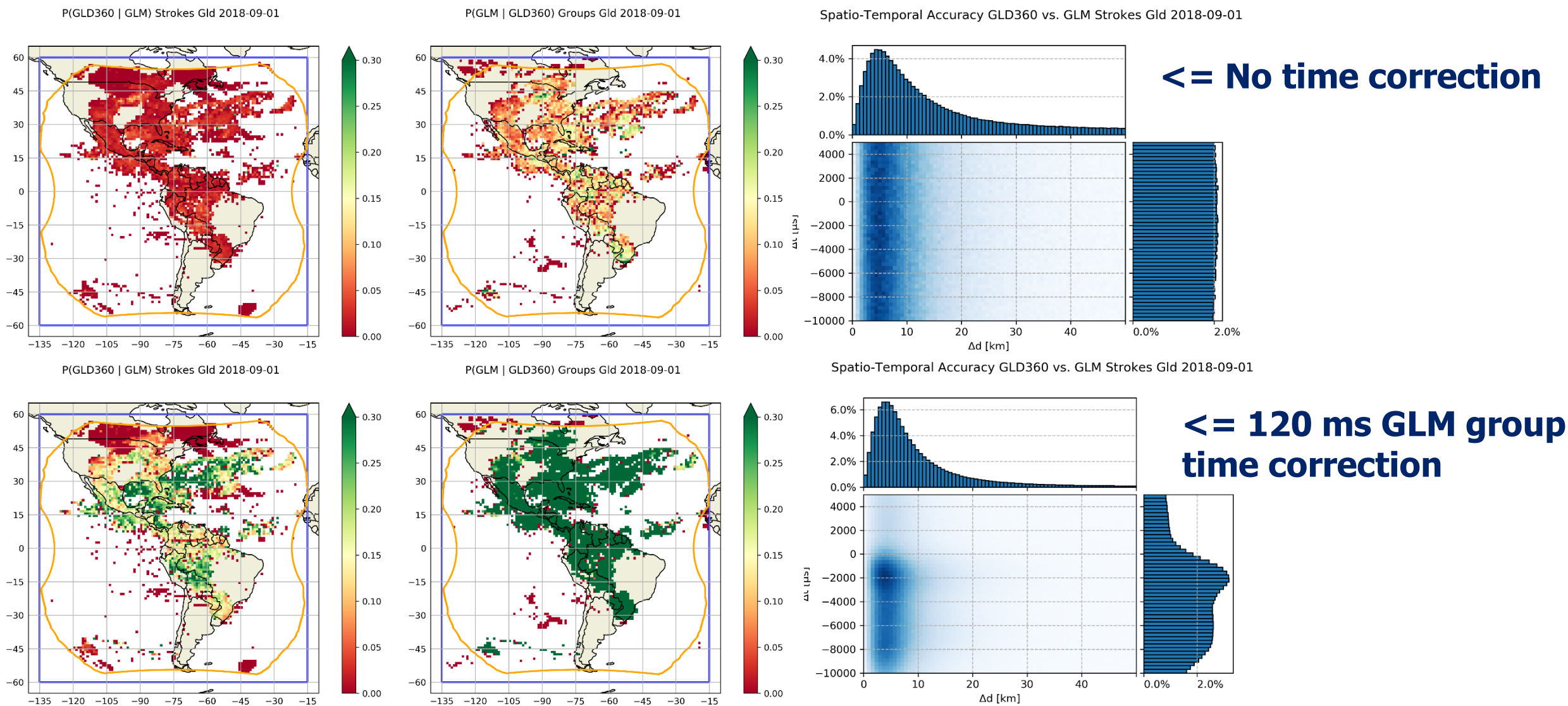


Can redo matching, using much wider time window, to check for possible timing errors/problems.



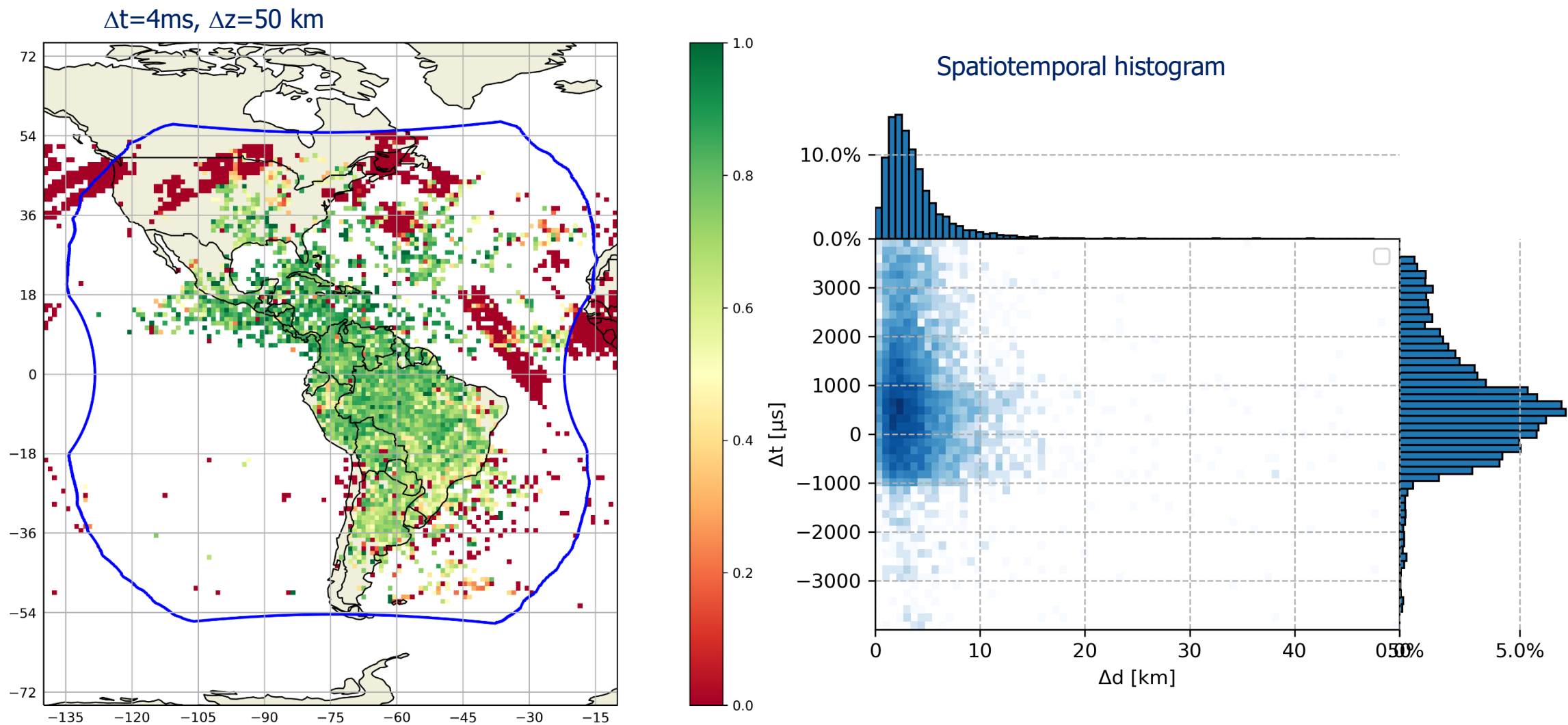
Conclusion – GLM times not photon travel time corrected (09/2018)

Testing coarse 120 ms GLM group time correction



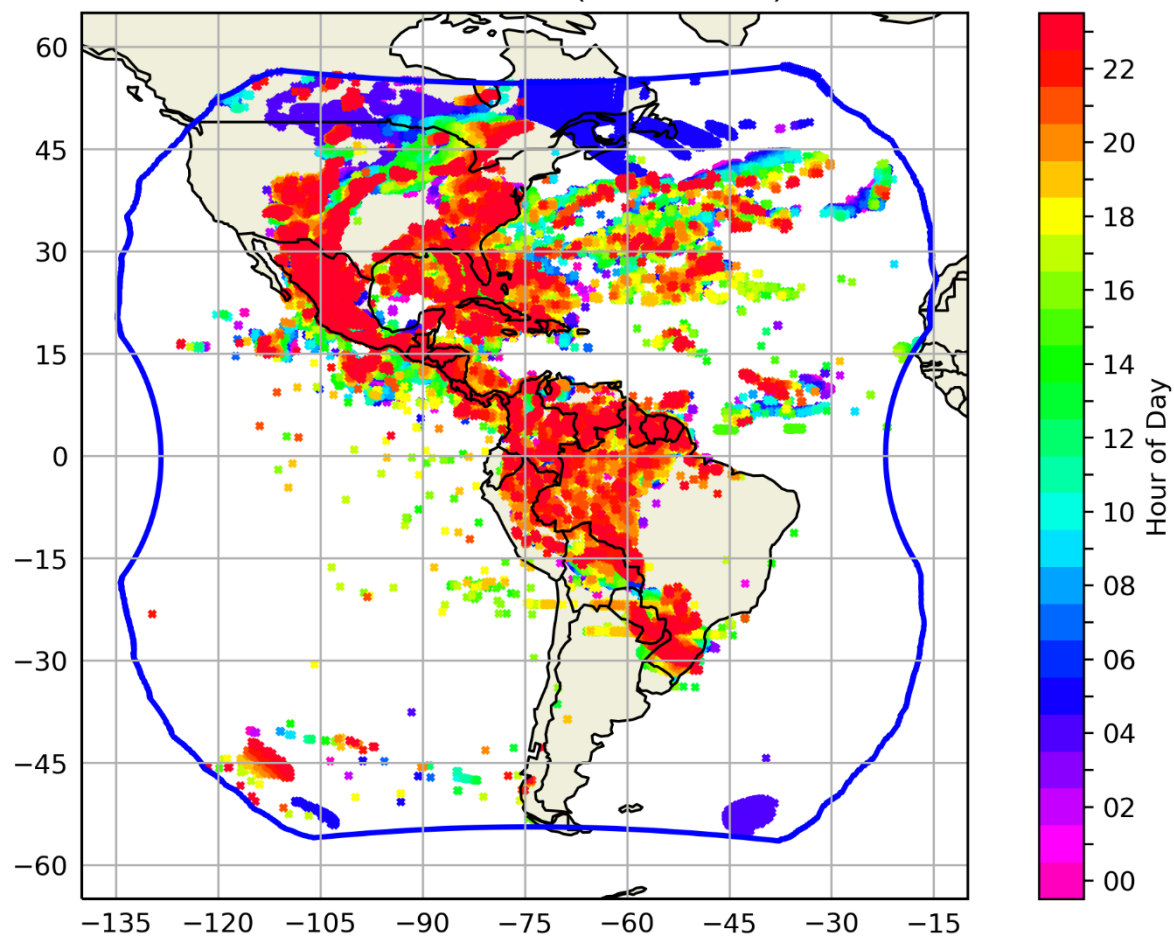
ISSLIS vs GLM group DE Oct-Nov 2020

P(GLM | ISS-LIS) Groups GIm 2020-10-01..2020-11-30



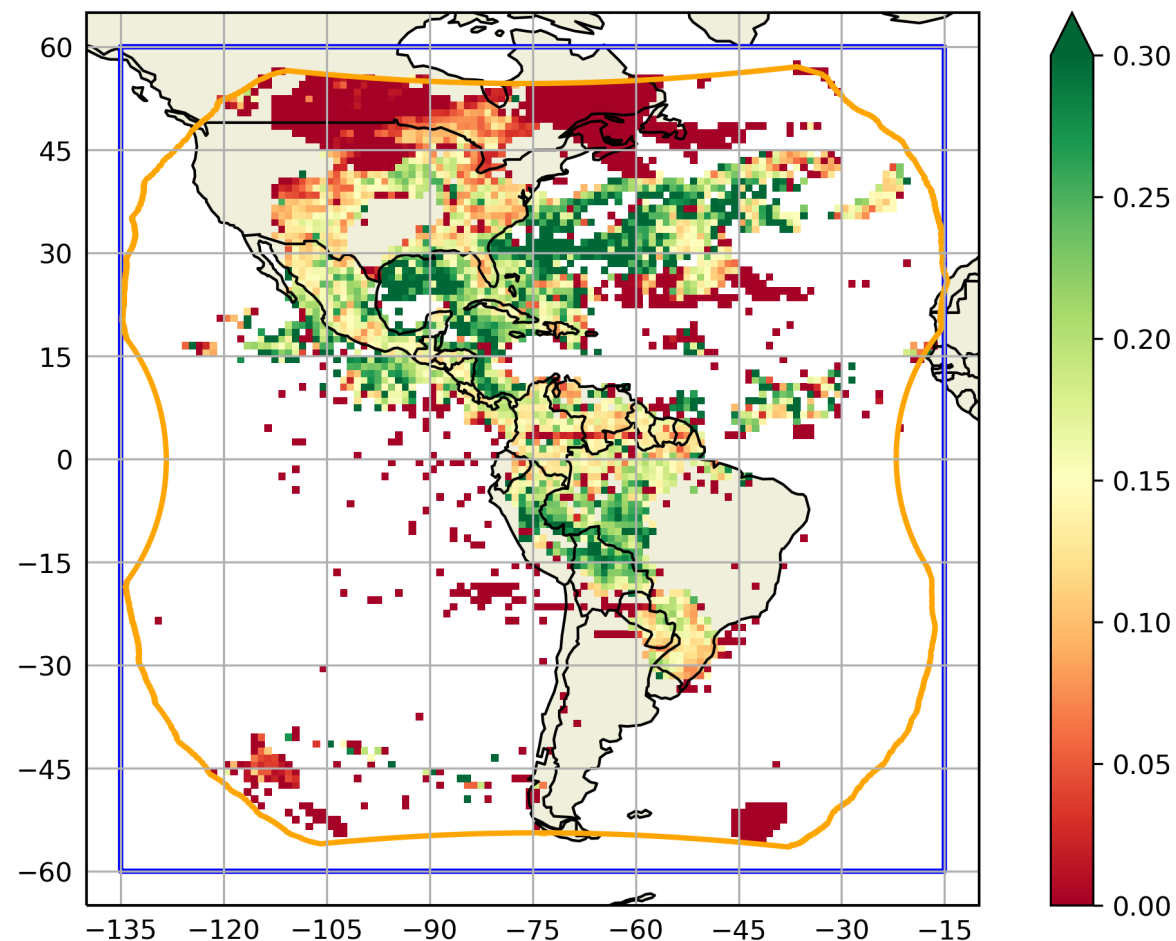
GLD360-GLM group DE 01-09-2018

GLM Flash Plot Gld (2018-09-01)



P(GLD360 | GLM) Strokes Gld 2018-09-01

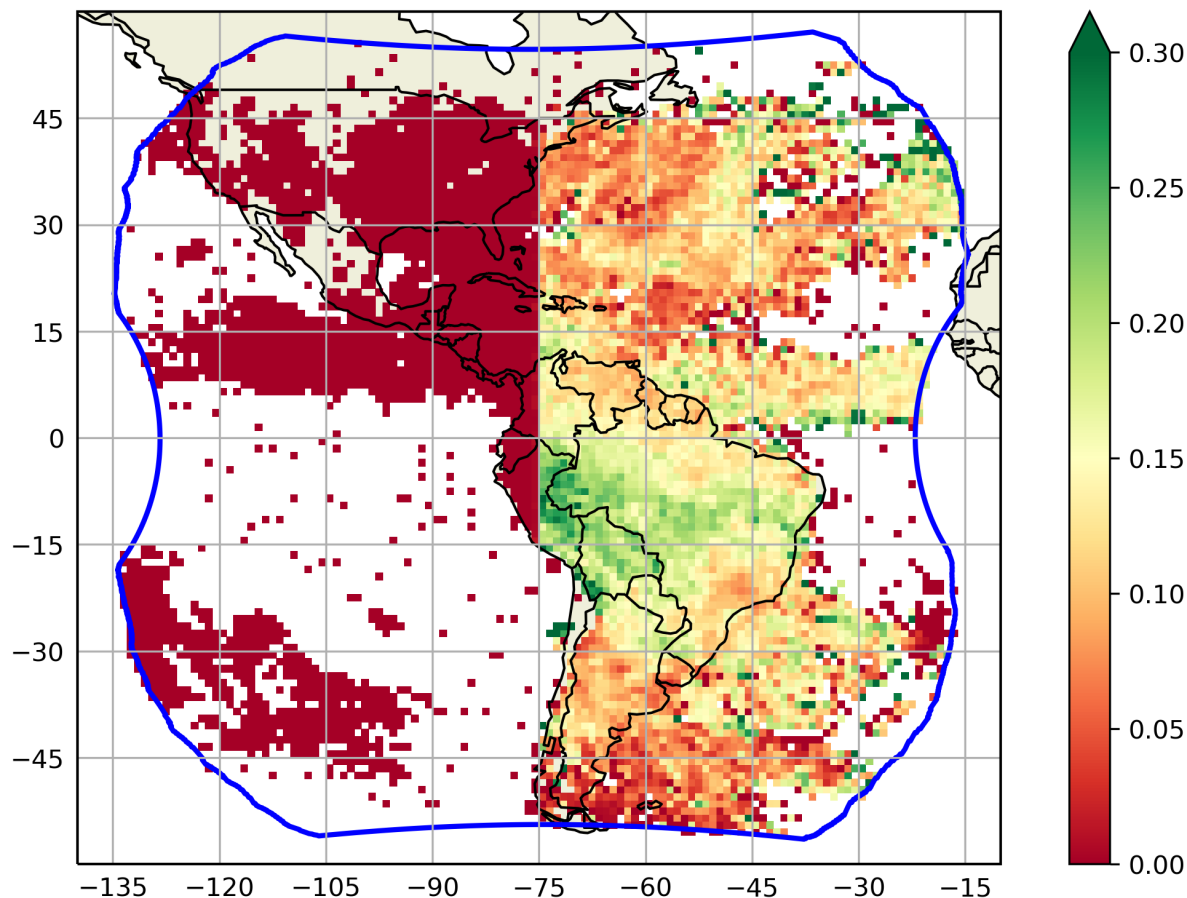
$\Delta t_b=10\text{ms}$, $\Delta t_a=5\text{ms}$, $\Delta z=50\text{ km}$, 1-to-many matching



GLD has higher GLM and lower ISS-LIS group DE?

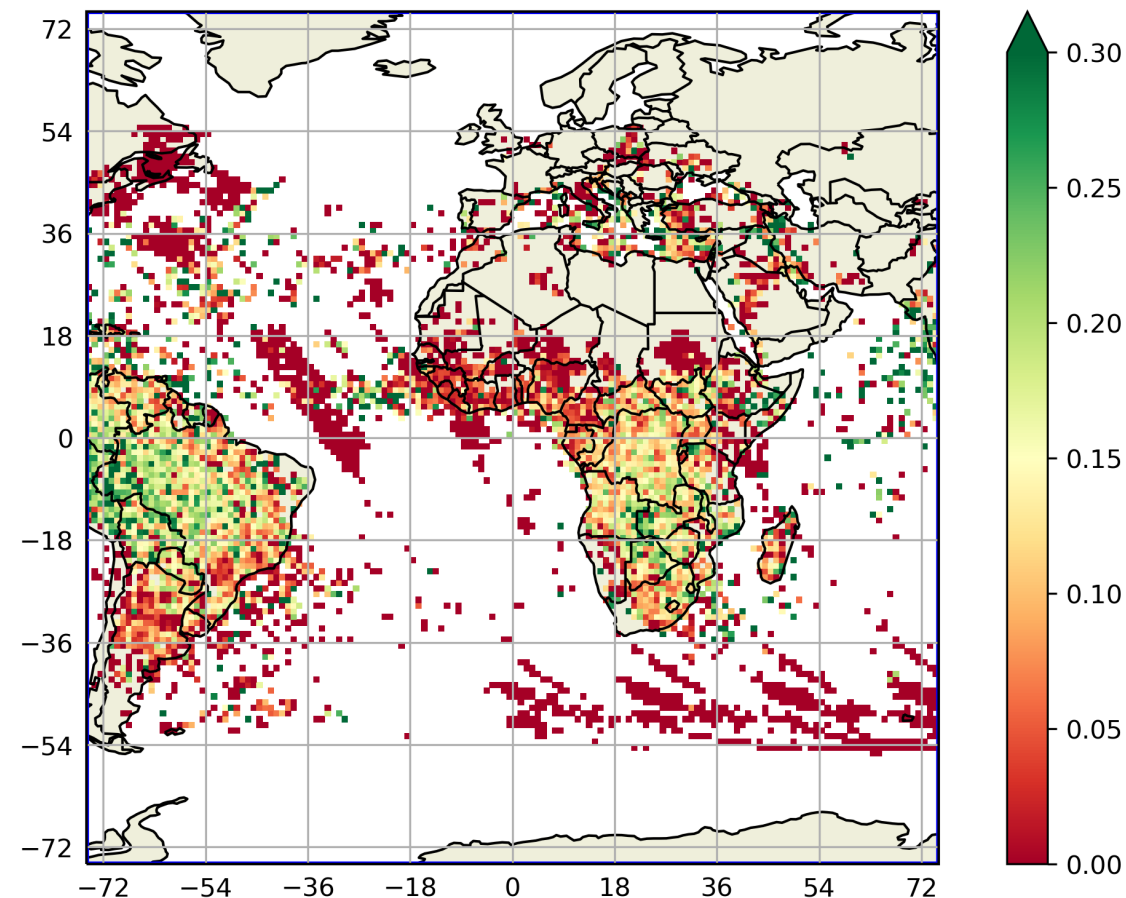
P(GLD360 | GLM) Strokes Americas 2020-11

$\Delta t_b=10\text{ms}$, $\Delta t_a=5\text{ms}$, $\Delta z=50\text{ km}$, 1-to-many matching



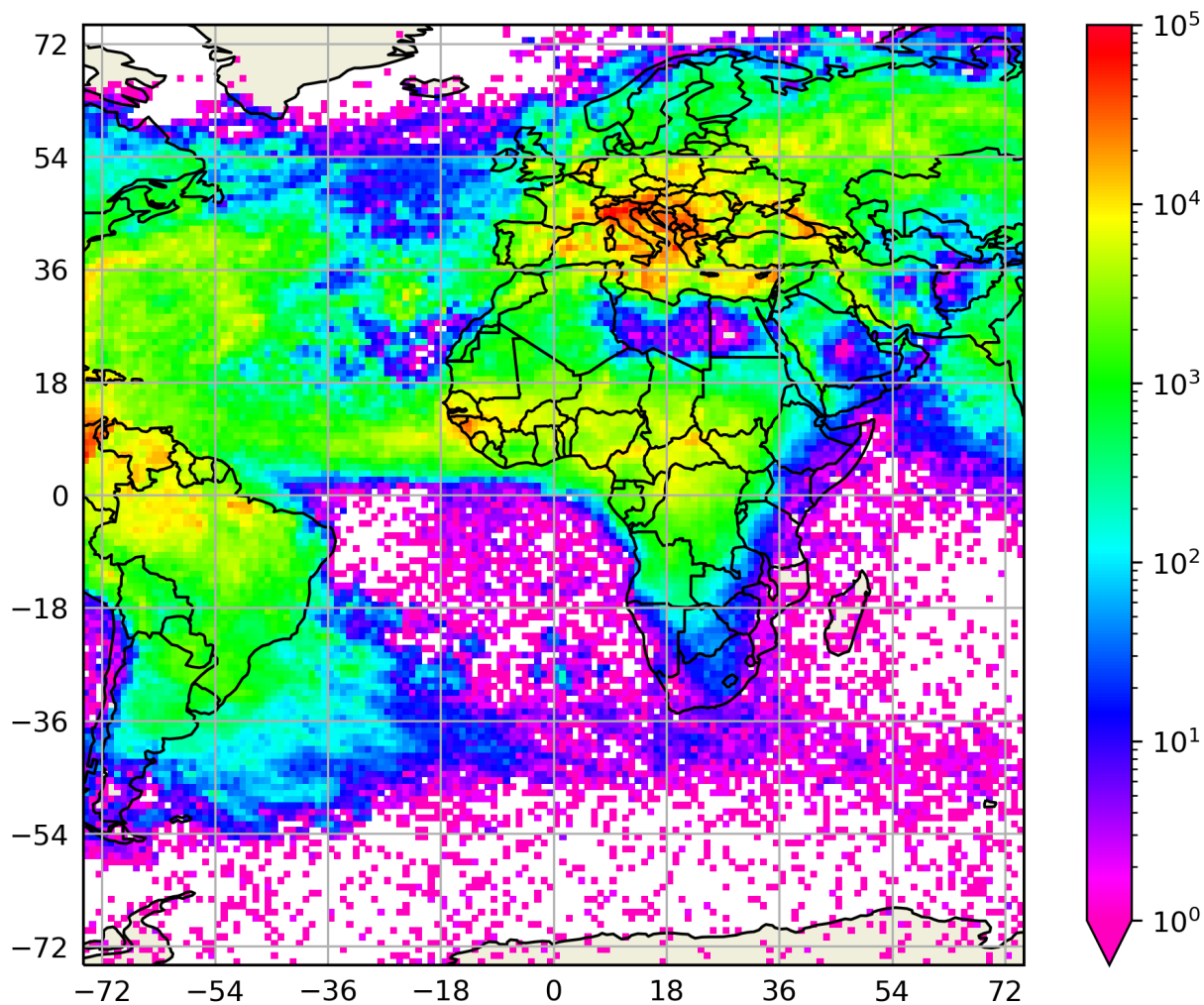
P(GLD360 | ISS-LIS) Strokes Gld 2020-10-01..2020-11-30

$\Delta t_b=10\text{ms}$, $\Delta t_a=5\text{ms}$, $\Delta z=50\text{ km}$, 1-to-many matching

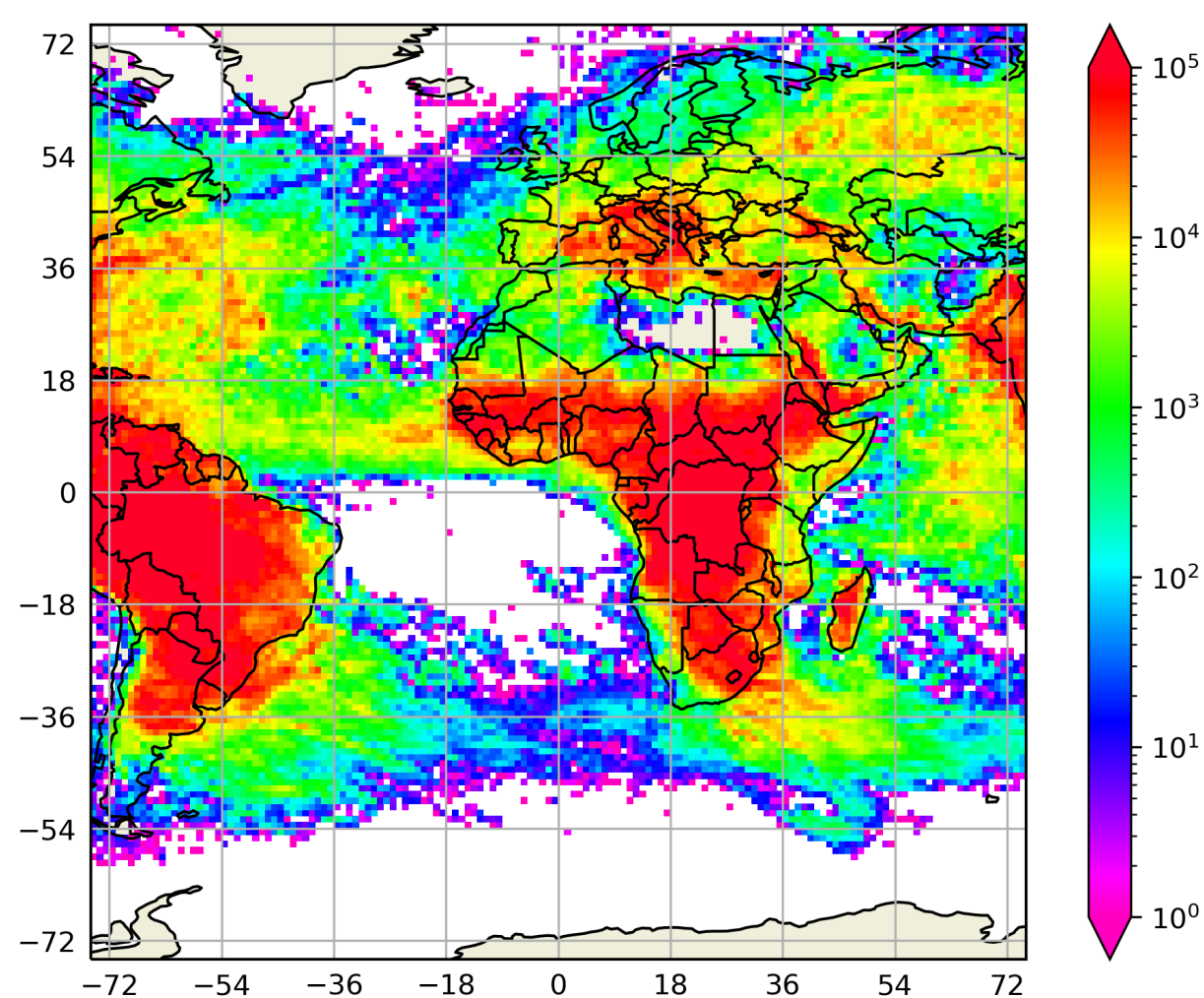


Assessing ATDnet spatial range I

ATDnet Accumulated Strokes Gld 2020-07-01..2020-11-30



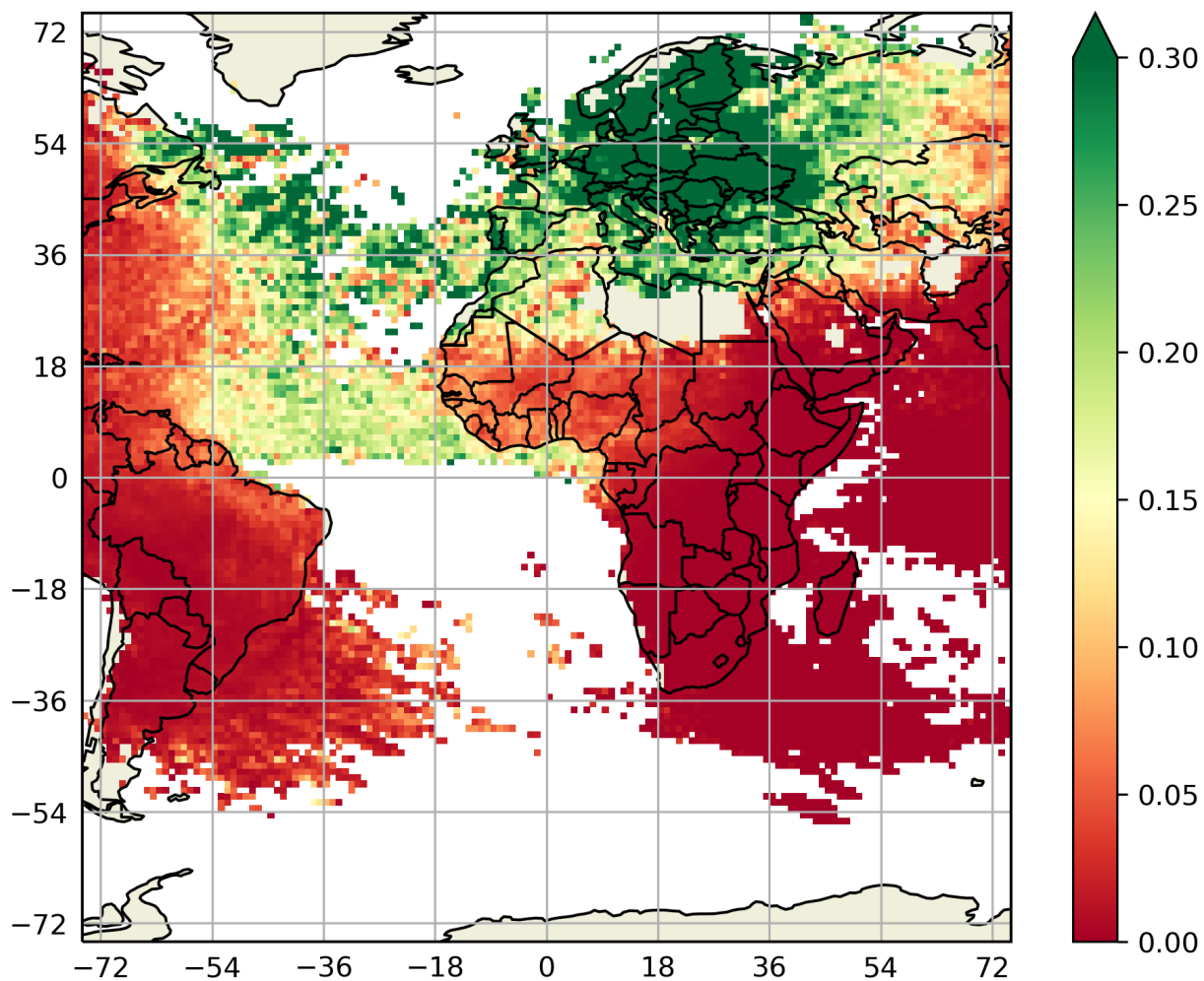
GLD360 Accumulated Strokes Gld 2020-07-01..2020-11-30



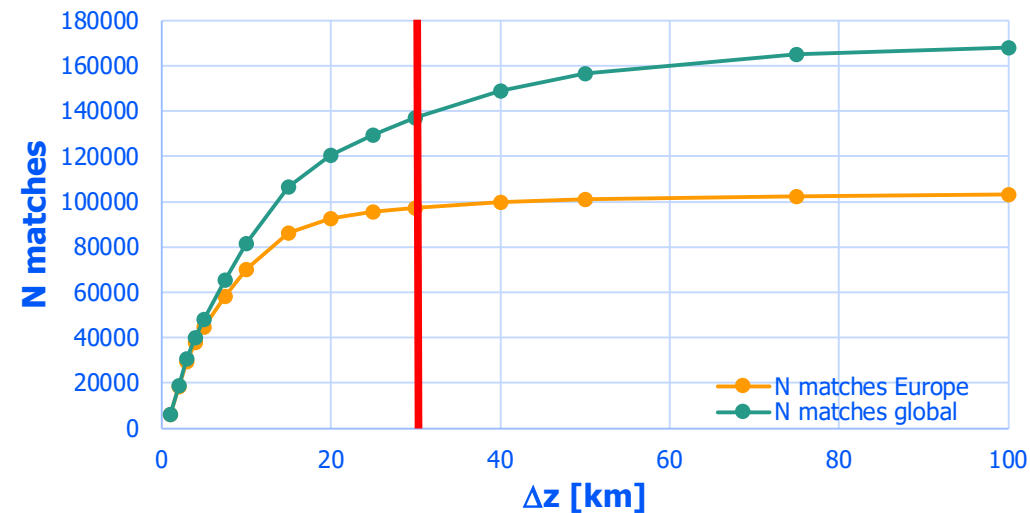
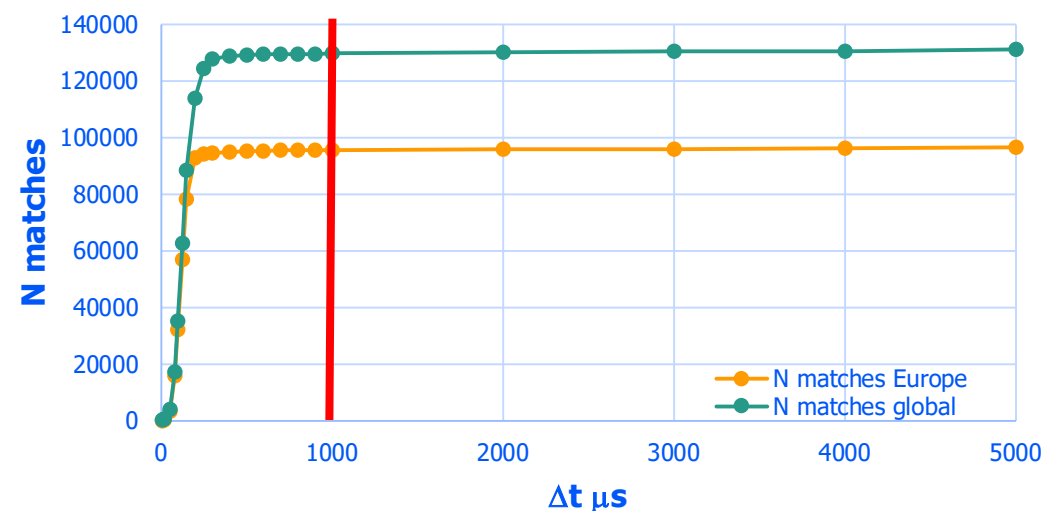
Assessing ATDnet spatial range II

P(ATDnet | GLD360) Strokes Gld 2020-07-01..2020-11-30

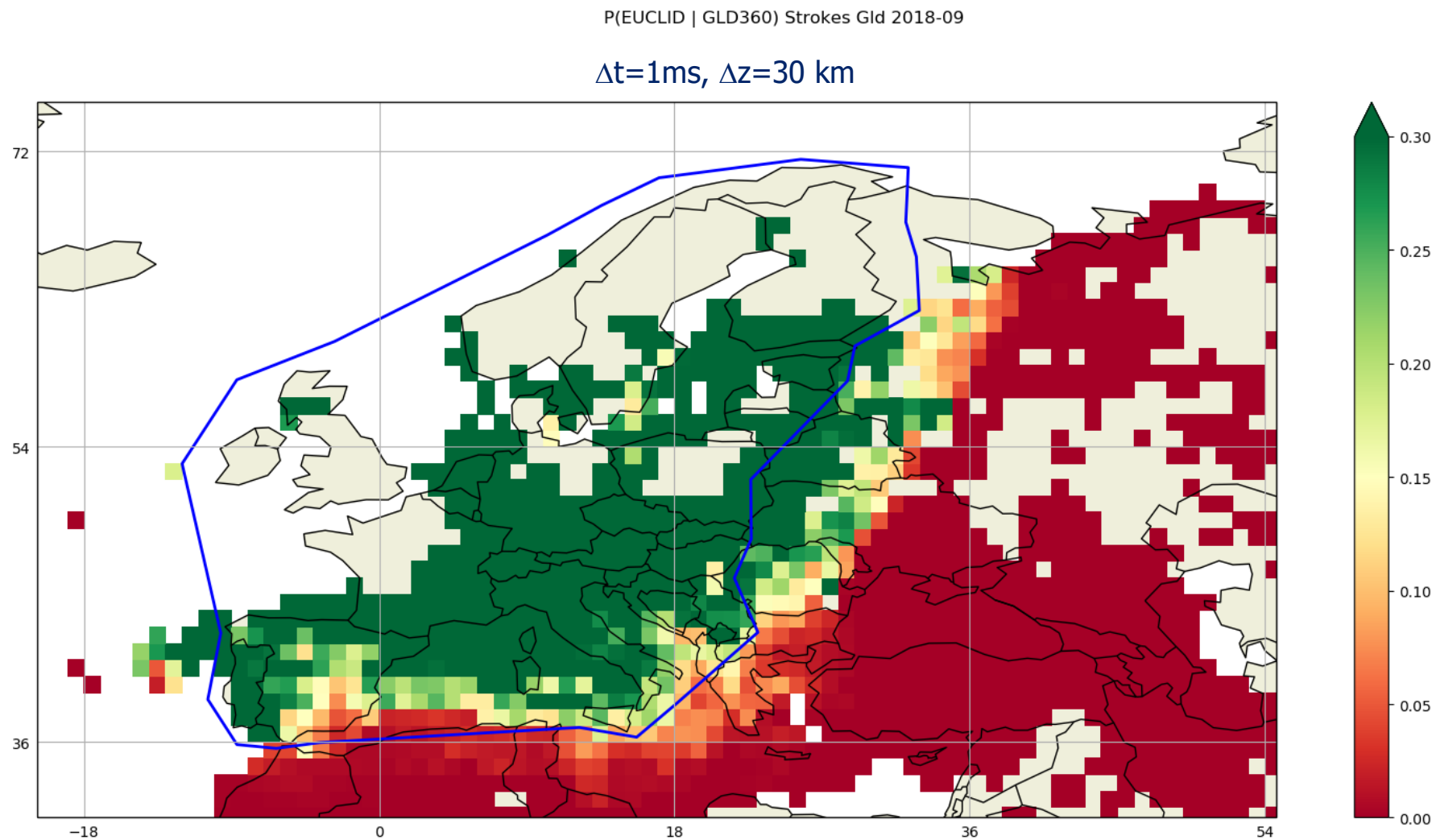
$\Delta t=1\text{ms}$, $\Delta z=30\text{ km}$



Why $\Delta t=1\text{ms}$, $\Delta z=30\text{ km}$?

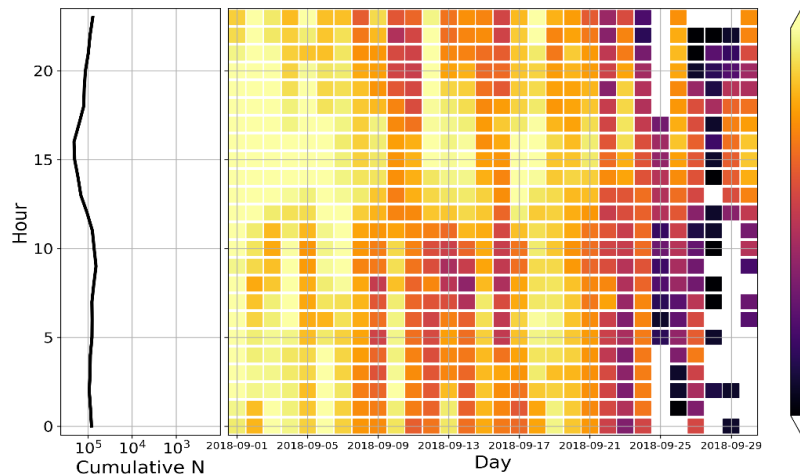


Checking EUCLID spatial range

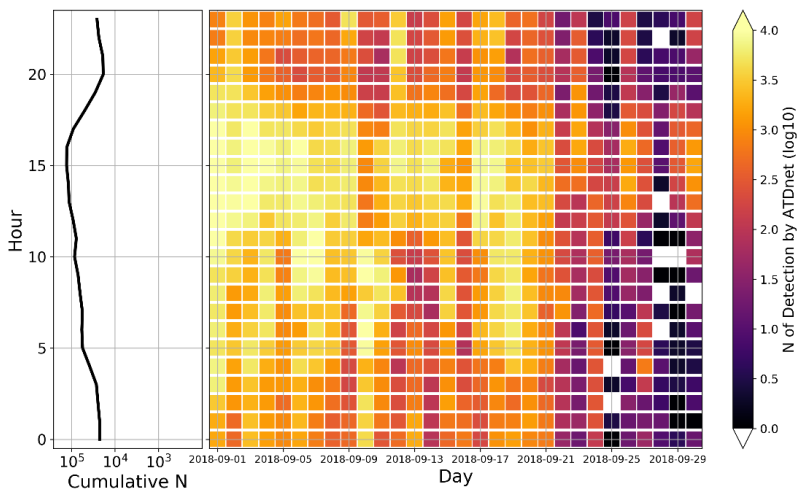


ATDnet vs EUCLID strokes September 2018 I

N of Detections (Strokes), EUCLID, Euclid,2018-09

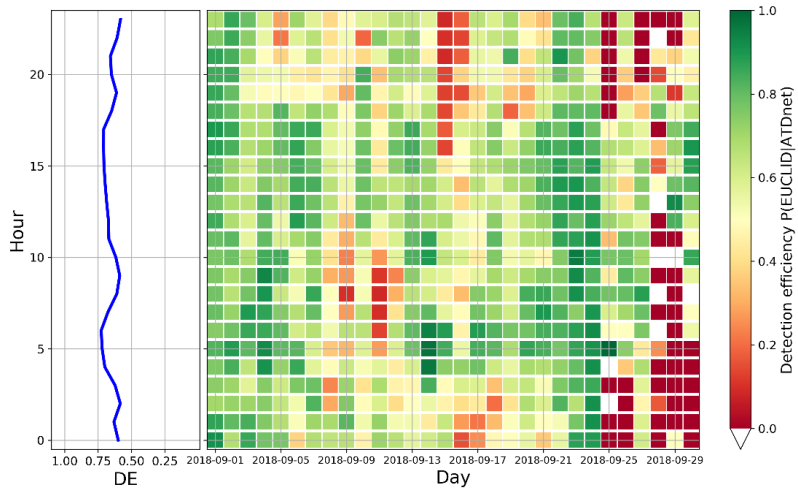


N of Detections (Strokes), ATDnet, Euclid,2018-09

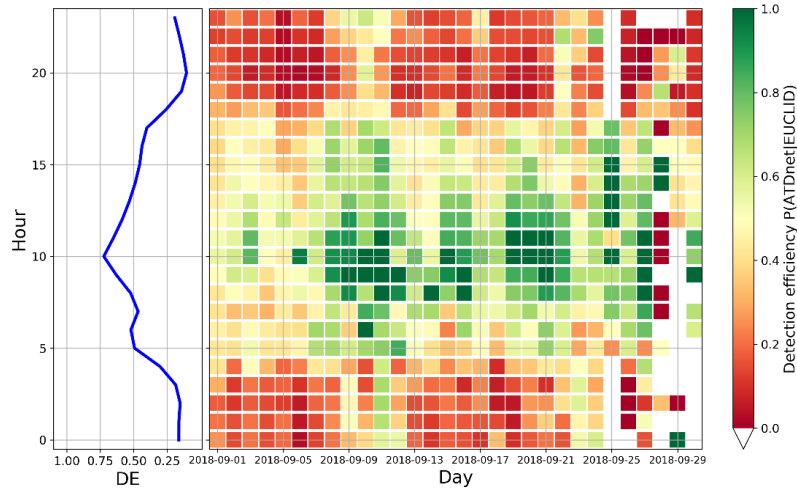


<= stroke counts

Detection Efficiency (Strokes), Euclid,2018-09



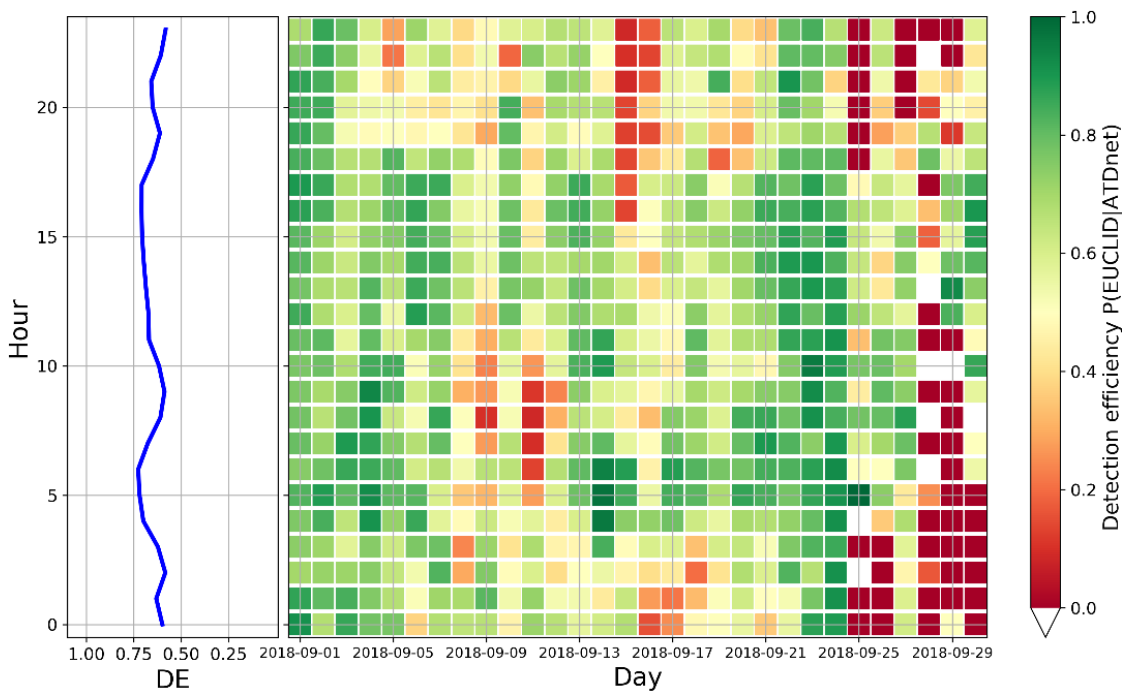
Detection Efficiency (Strokes), Euclid,2018-09



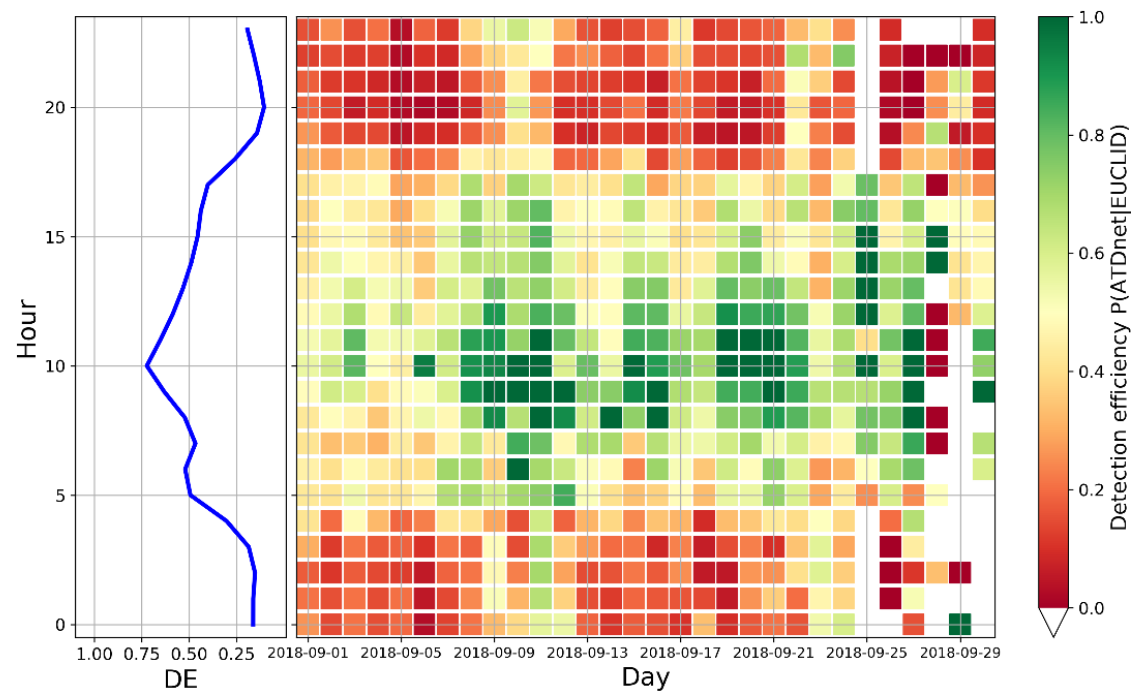
<= relative DE

ATDnet vs EUCLID strokes September 2018 II

Detection Efficiency (Strokes), Euclid,2018-09



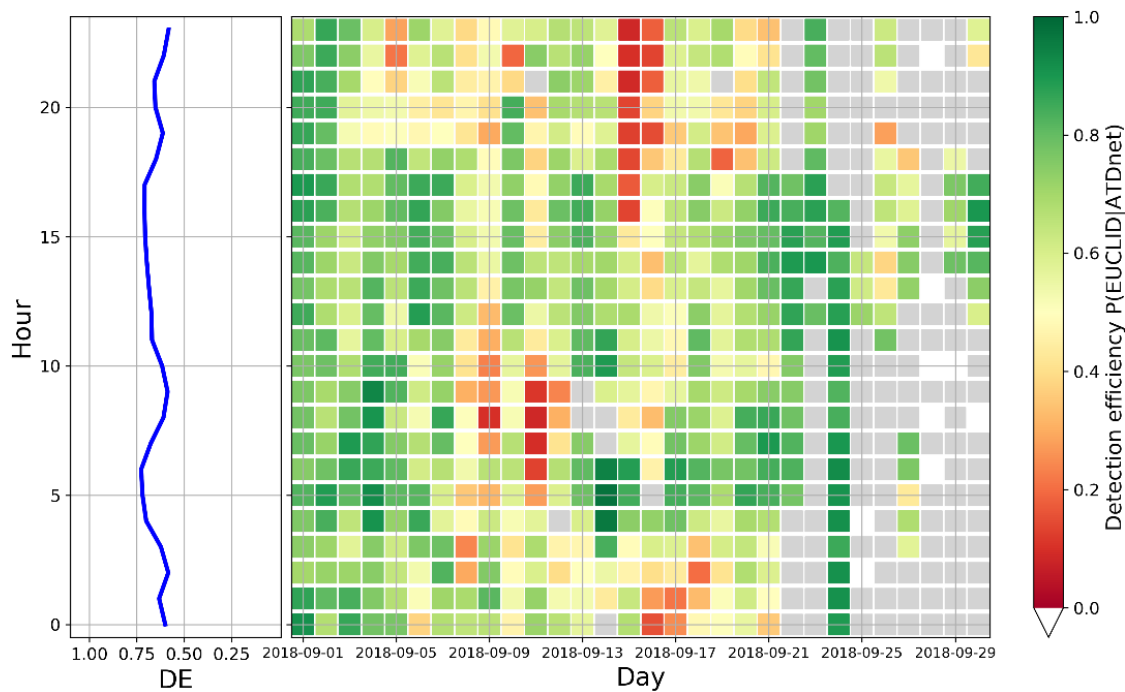
Detection Efficiency (Strokes), Euclid,2018-09



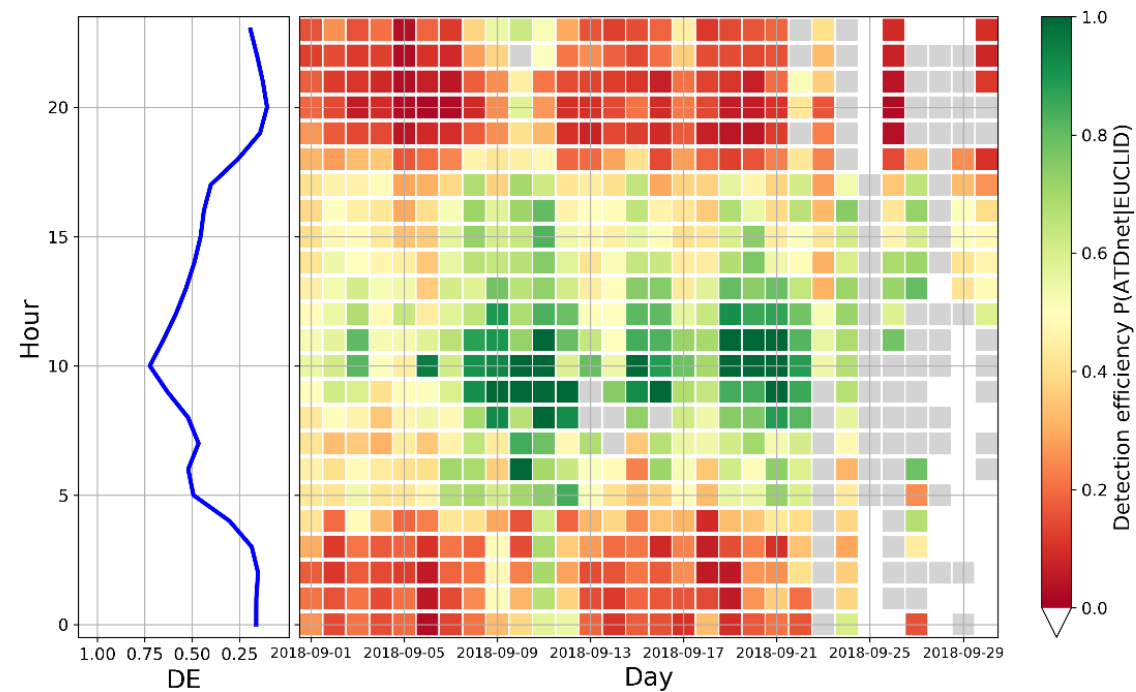
Hours with too low lightning activity will give unreliable results...

ATDnet vs EUCLID strokes September 2018 III

Detection Efficiency (Strokes), Euclid,2018-09, min_ref=100



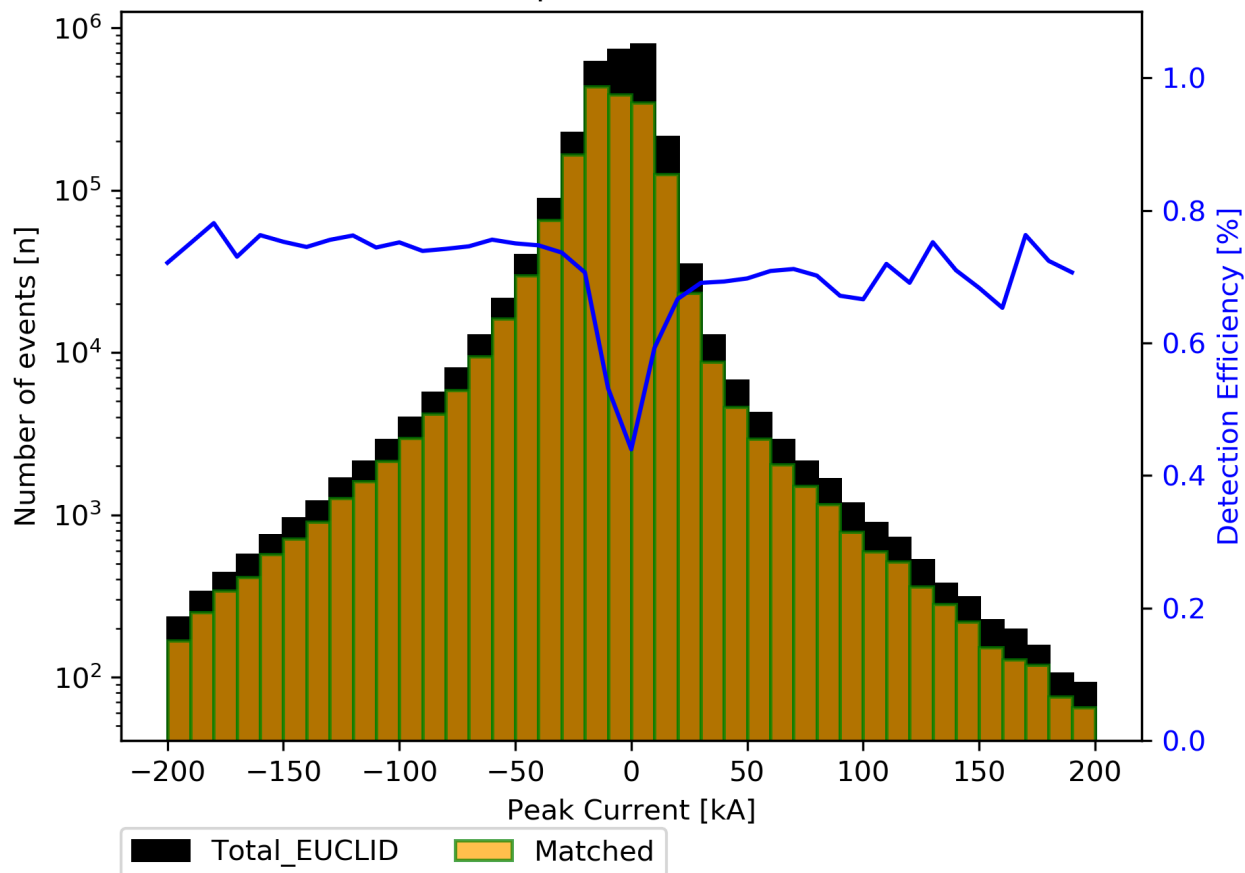
Detection Efficiency (Strokes), Euclid,2018-09, min_ref=100



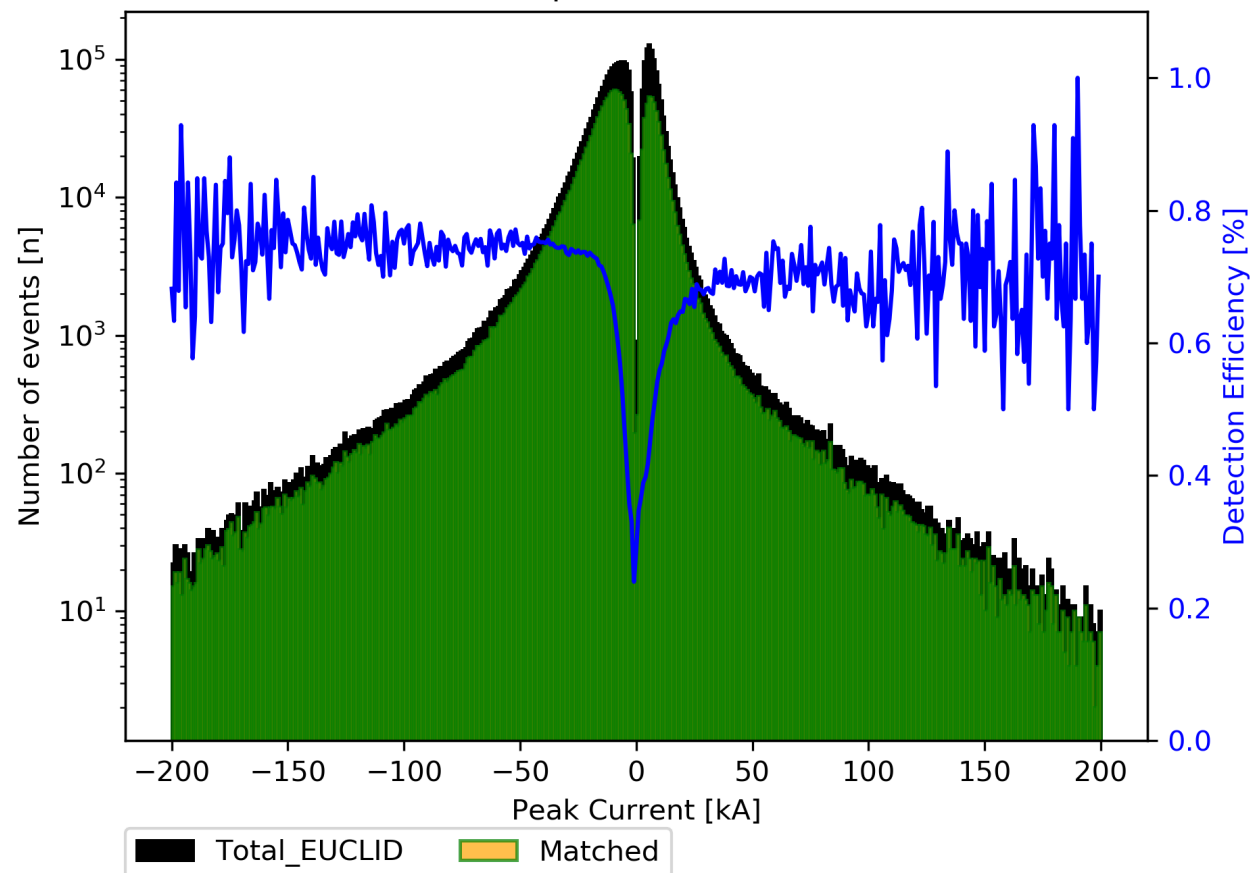
...so remove hours with <100 reference network strokes.

ATDnet vs EUCLID stroke DE vs peak current

EUCLID, Europe, 2018-09, bin= 10 kA



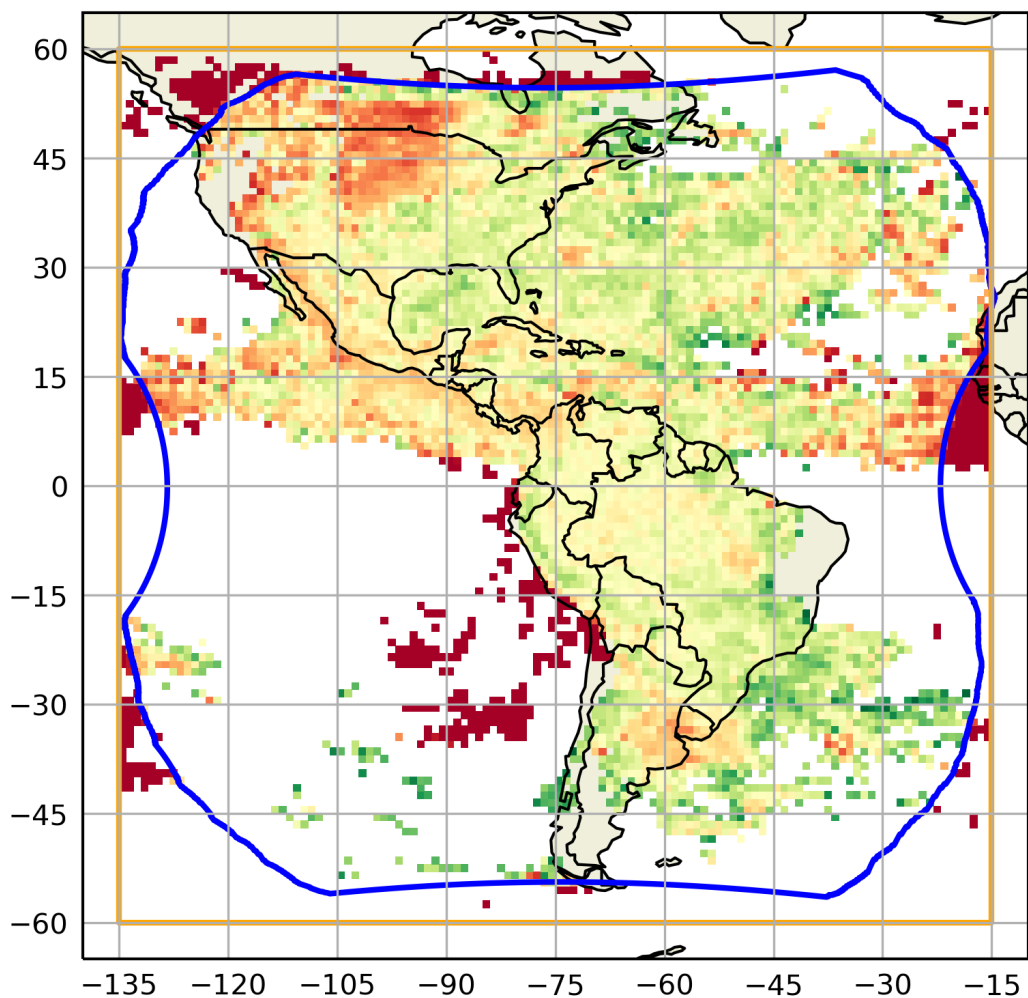
EUCLID, Europe, 2018-09, bin= 1 kA



GLD360 and GLM flash DE September 2018 - primitive

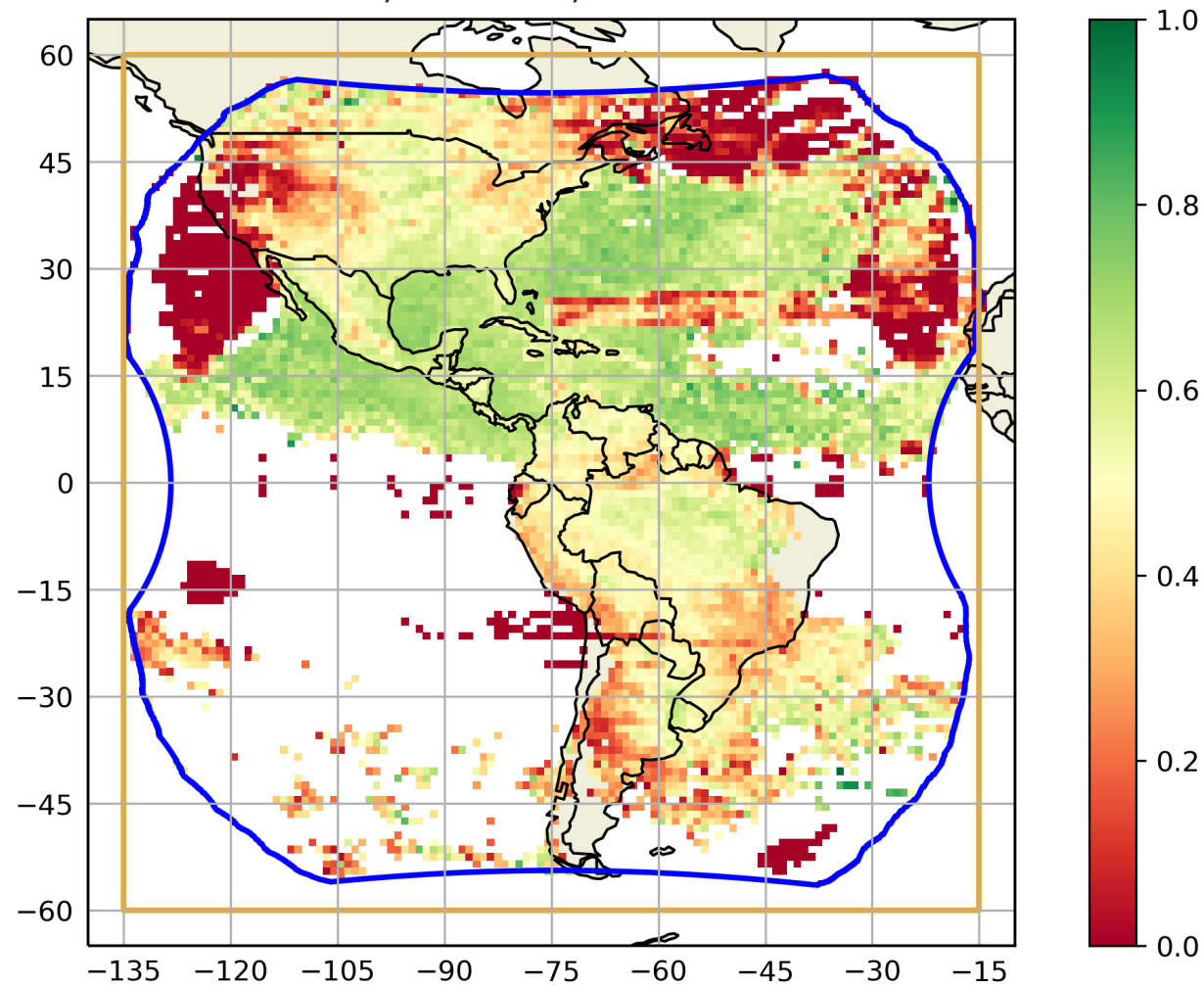
P(GLM | GLD360) Flashes Gld 2018-09

$\Delta t=0.2s$, $\Delta z=50$ km, 1-to-1 matches



P(GLD360 | GLM) Flashes Gld 2018-09

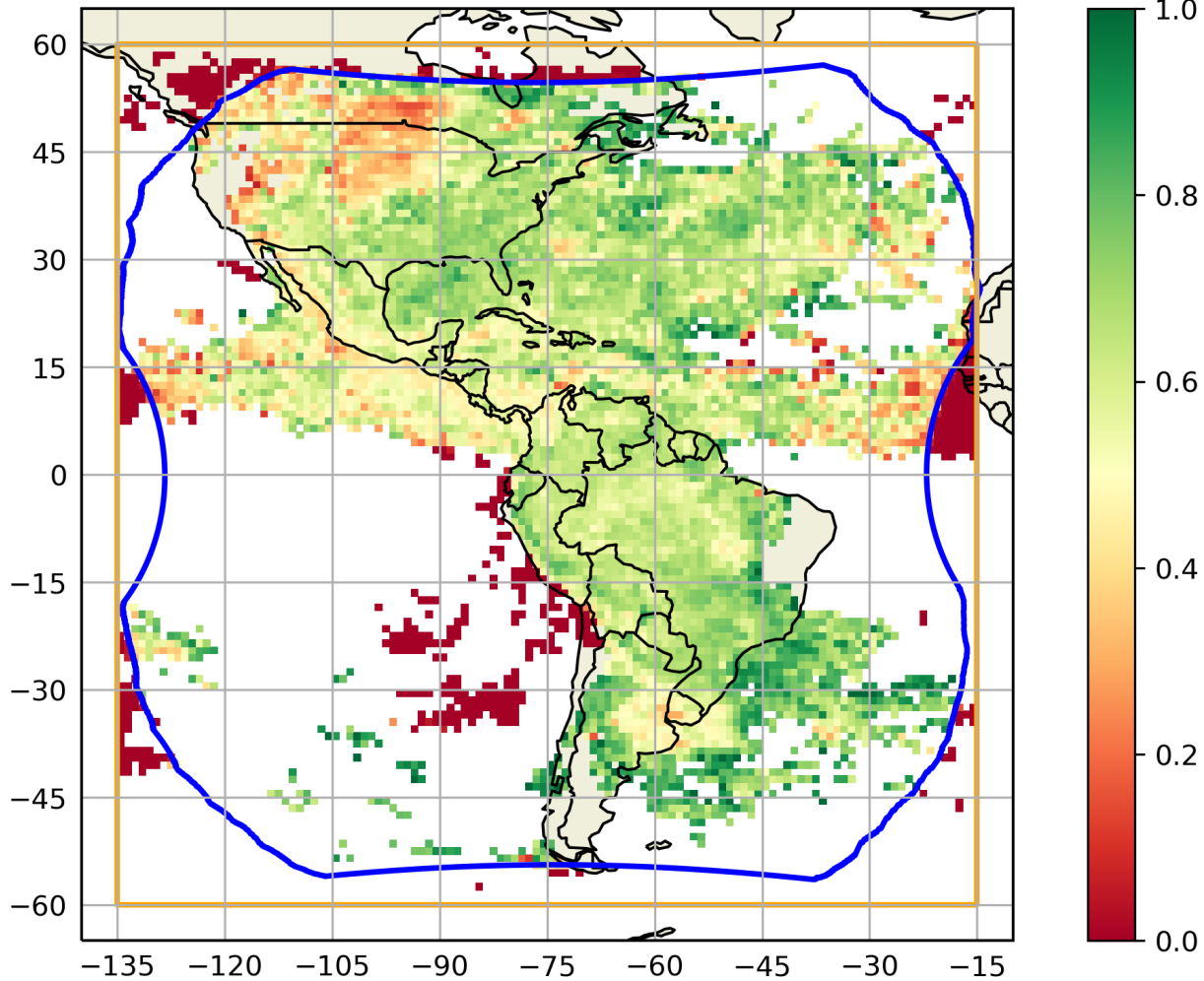
$\Delta t=0.2s$, $\Delta z=50$ km, 1-to-1 matches



GLD360 and GLM flash DE September 2018 - advanced

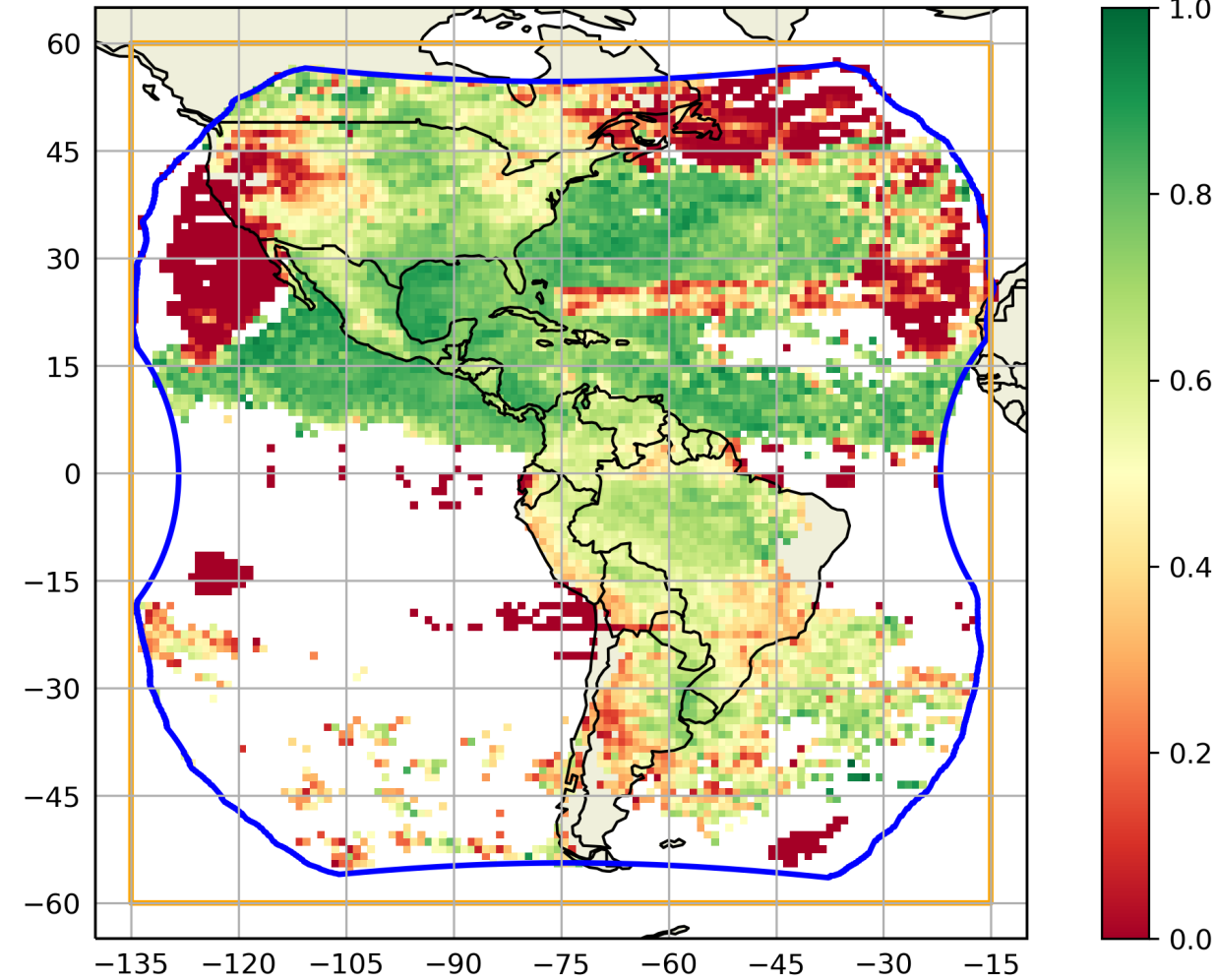
P(GLM | GLD360) Flashes Gld 2018-09

$\Delta t=0.2s$, $\Delta z=50$ km, 1-to-1 matches



P(GLD360 | GLM) Flashes Gld 2018-09

$\Delta t=0.2s$, $\Delta z=50$ km, 1-to-1 matches



The current stroke/pulse/group matching parameters

Reference network	Test network				
	ATDnet	EUCLID	GLD360	GLM	ISSLIS
ATDnet		$\Delta t = 1 \text{ ms}$ $\Delta z = 30 \text{ km}$	$\Delta t = 1 \text{ ms}$ $\Delta z = 30 \text{ km}$	$\Delta t_b = 5 \text{ ms}$ $\Delta t_a = 10 \text{ ms}$ $\Delta z = 50 \text{ km}$	$\Delta t_b = 5 \text{ ms}$ $\Delta t_a = 10 \text{ ms}$ $\Delta z = 50 \text{ km}$
EUCLID	$\Delta t = 1 \text{ ms}$ $\Delta z = 30 \text{ km}$		$\Delta t = 1 \text{ ms}$ $\Delta z = 20 \text{ km}$		$\Delta t_b = 5 \text{ ms}$ $\Delta t_a = 10 \text{ ms}$ $\Delta z = 50 \text{ km}$
GLD360	$\Delta t = 1 \text{ ms}$ $\Delta z = 30 \text{ km}$	$\Delta t = 1 \text{ ms}$ $\Delta z = 20 \text{ km}$		$\Delta t_b = 5 \text{ ms}$ $\Delta t_a = 10 \text{ ms}$ $\Delta z = 50 \text{ km}$	$\Delta t_b = 5 \text{ ms}$ $\Delta t_a = 10 \text{ ms}$ $\Delta z = 50 \text{ km}$
GLM	$\Delta t_b = 10 \text{ ms}$ $\Delta t_a = 5 \text{ ms}$ $\Delta z = 50 \text{ km}$		$\Delta t_b = 10 \text{ ms}$ $\Delta t_a = 5 \text{ ms}$ $\Delta z = 50 \text{ km}$		$\Delta t = 4 \text{ ms}$ $\Delta z = 50 \text{ km}$
ISSLIS	$\Delta t_b = 10 \text{ ms}$ $\Delta t_a = 5 \text{ ms}$ $\Delta z = 50 \text{ km}$	$\Delta t_b = 10 \text{ ms}$ $\Delta t_a = 5 \text{ ms}$ $\Delta z = 50 \text{ km}$	$\Delta t_b = 10 \text{ ms}$ $\Delta t_a = 5 \text{ ms}$ $\Delta z = 50 \text{ km}$	$\Delta t = 4 \text{ ms}$ $\Delta z = 50 \text{ km}$	

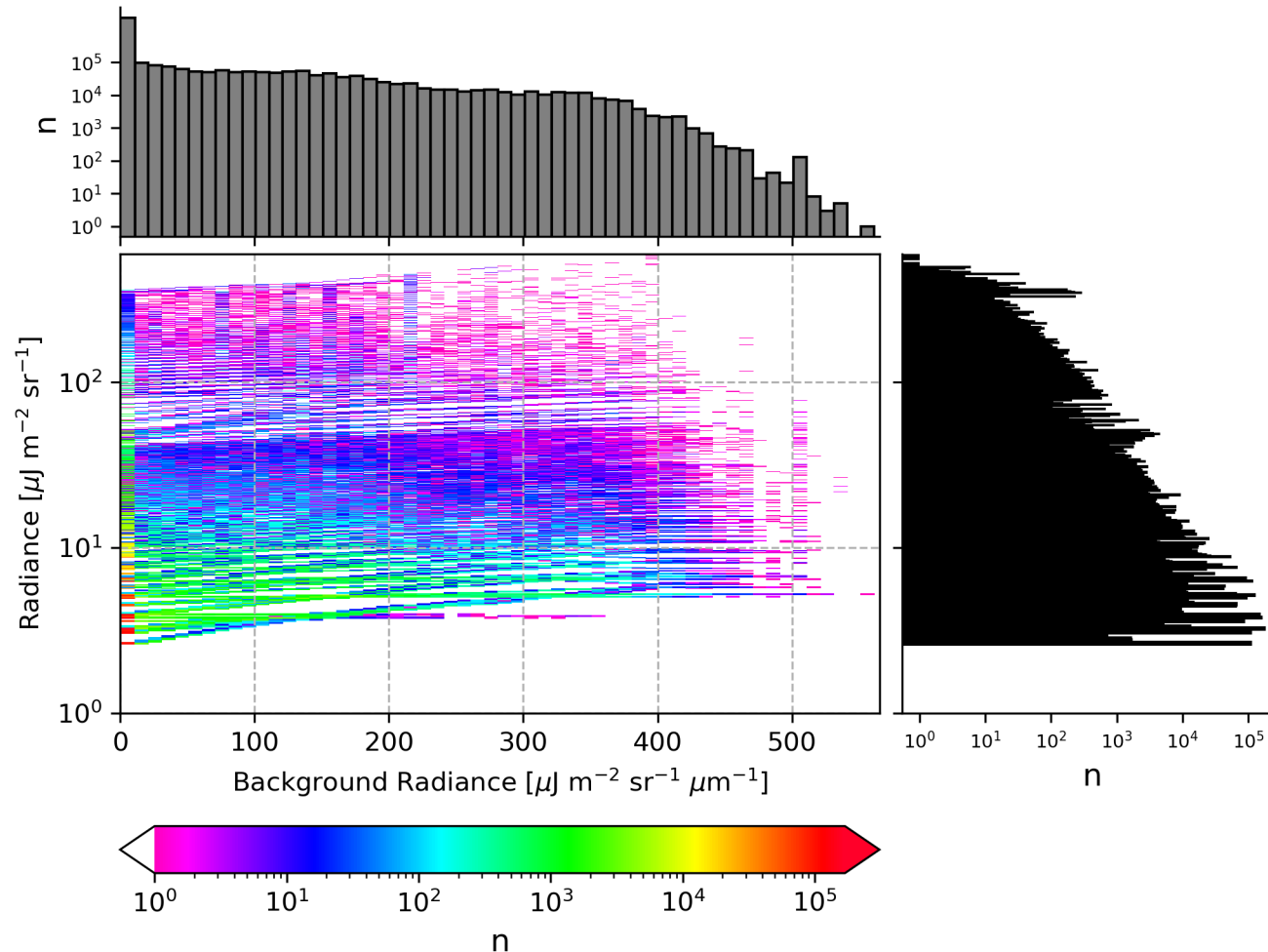
The current flash matching parameters

Reference network	Test network				
	ATDnet	EUCLID	GLD360	GLM	ISSLIS
ATDnet					
EUCLID					
GLD360				$\Delta t=200\text{ ms}$ $\Delta z=50\text{ km}$	$\Delta t=200\text{ ms}$ $\Delta z=50\text{ km}$
GLM			$\Delta t=200\text{ ms}$ $\Delta z=50\text{ km}$		$\Delta t=200\text{ ms}$ $\Delta z=50\text{ km}$
ISSLIS			$\Delta t=200\text{ ms}$ $\Delta z=50\text{ km}$	$\Delta t=200\text{ ms}$ $\Delta z=50\text{ km}$	

- ATDnet does not provide flashes, only pulses/strokes are available.
- EUCLID does not provide flash duration.
 - Could only do *group matching with extended time window* type of very primitive flash matching.

LI-STAR additional capabilities I

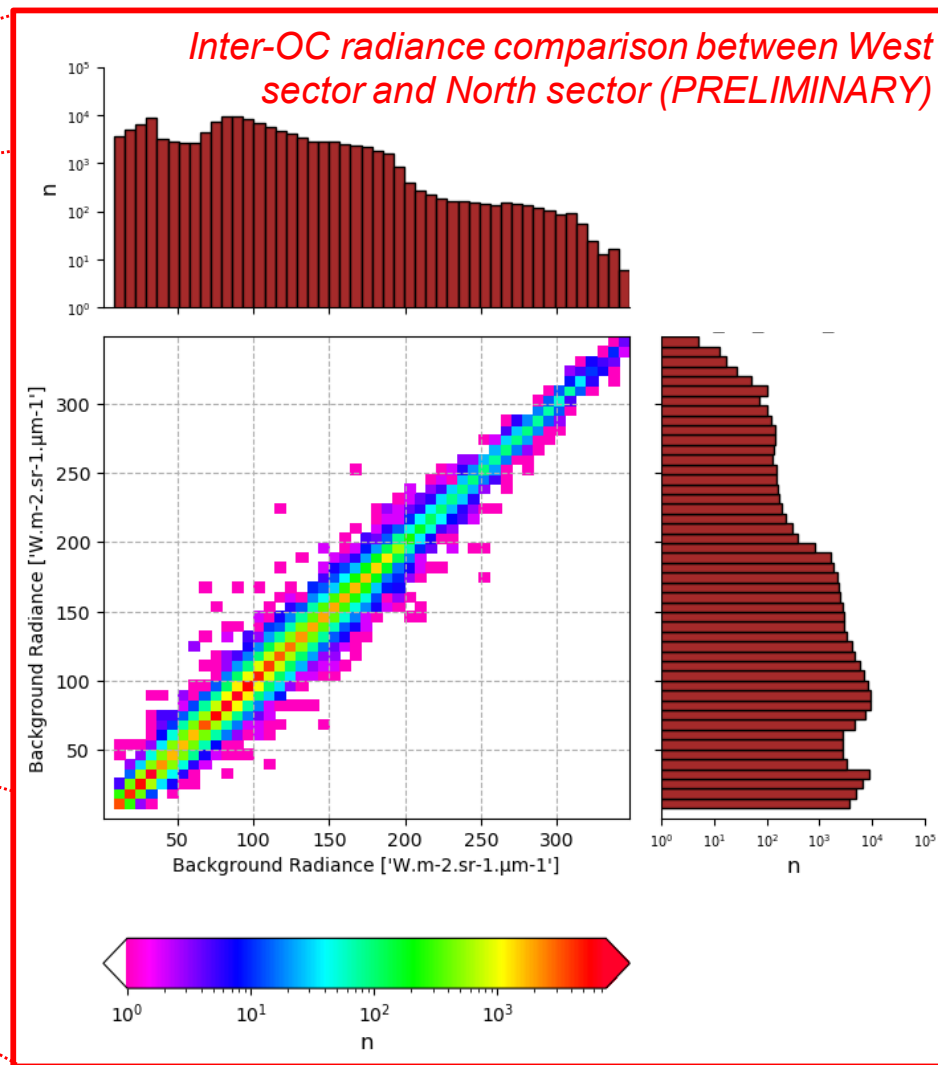
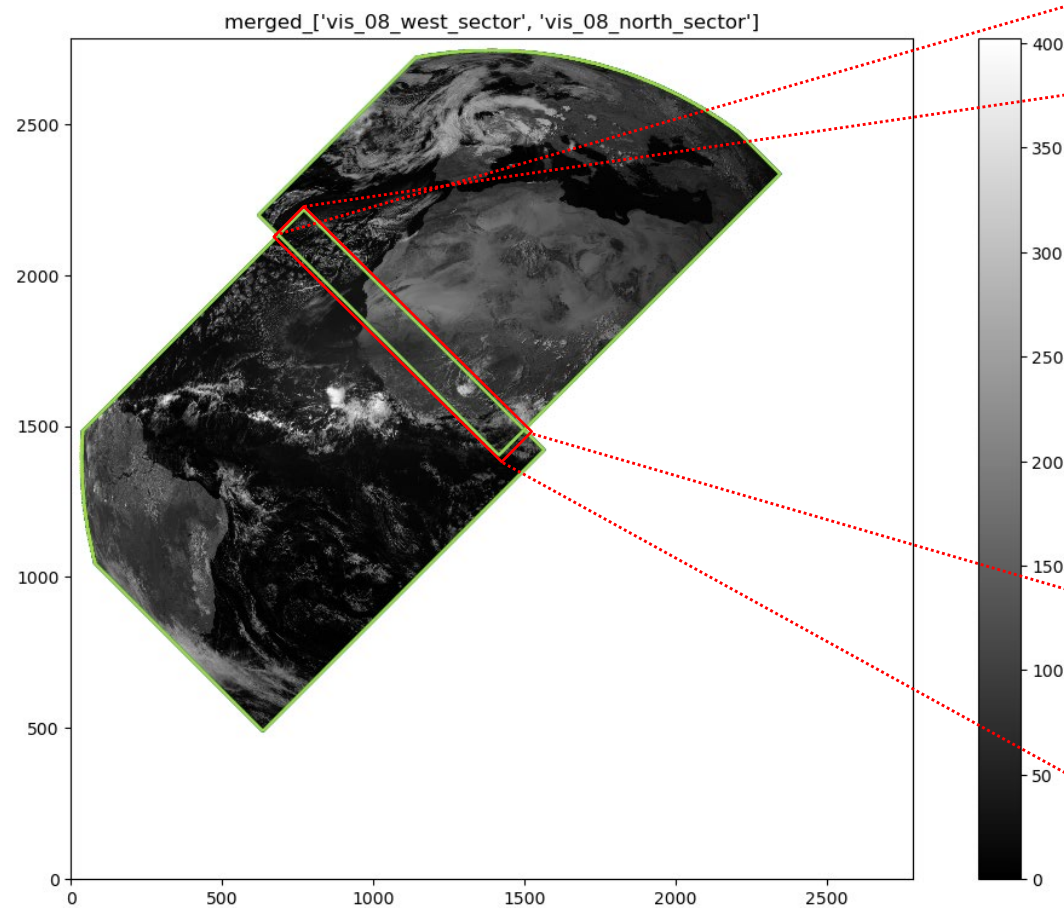
2D_dens_Histo_event__ISS-LIS_(2018-09-01..2018-09-30)_loc=Global



Monitoring event radiance
vs background radiance.
In this example using
ISS-LIS event information.

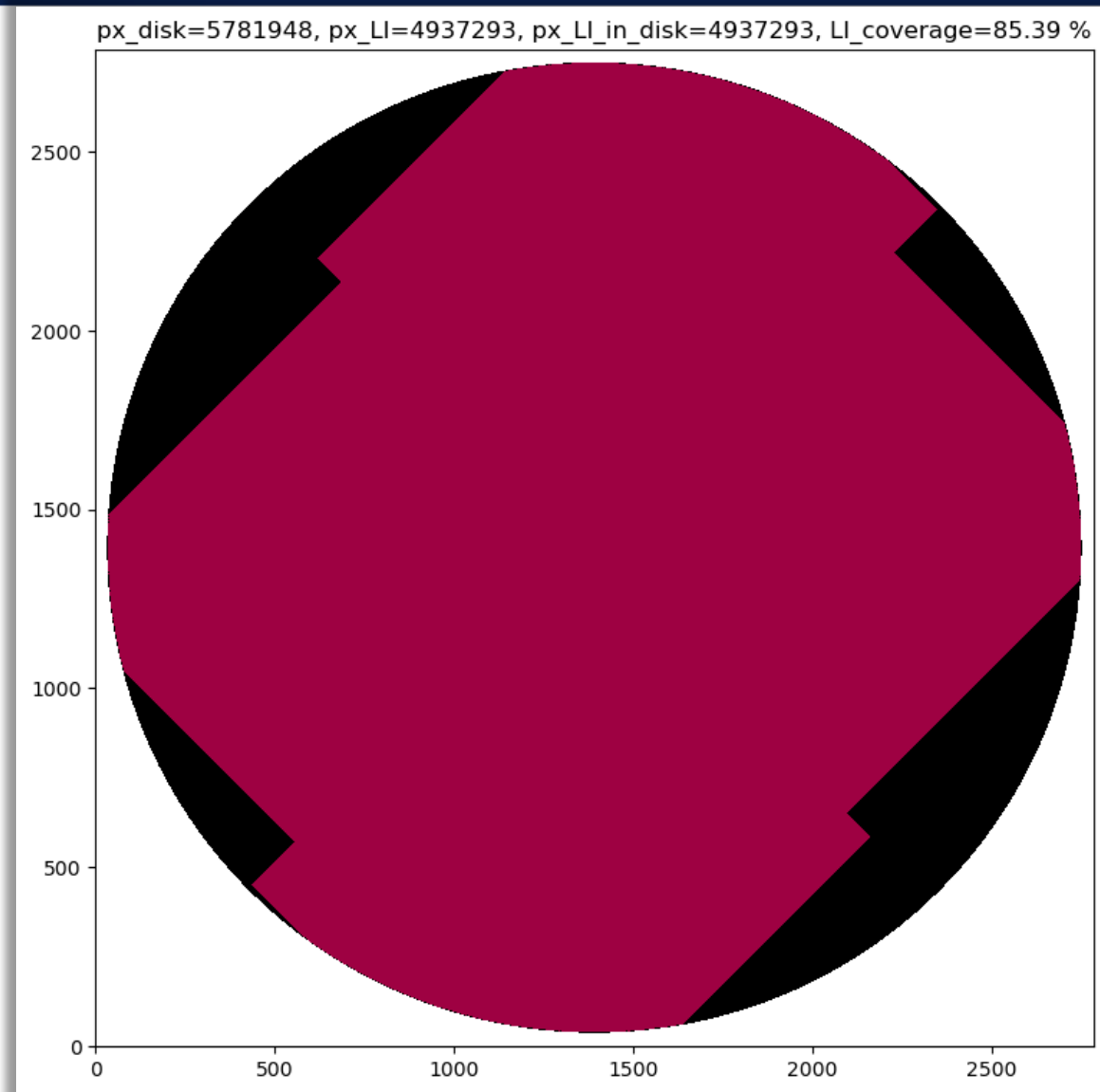
LI-STAR additional capabilities II

Monitoring of both radiometric and geometric performances



LI-STAR additional capabilities III

Monitoring the coverage against geo disk and against the member states



MISSING INFORMATION

- The LMA data over the US to compare against GLM.
- The description of the FOV of the LMA networks.
- Time information description in UTC (all external lightning datasets).
- Is GLM with photo travel time correction available for 09/2018?